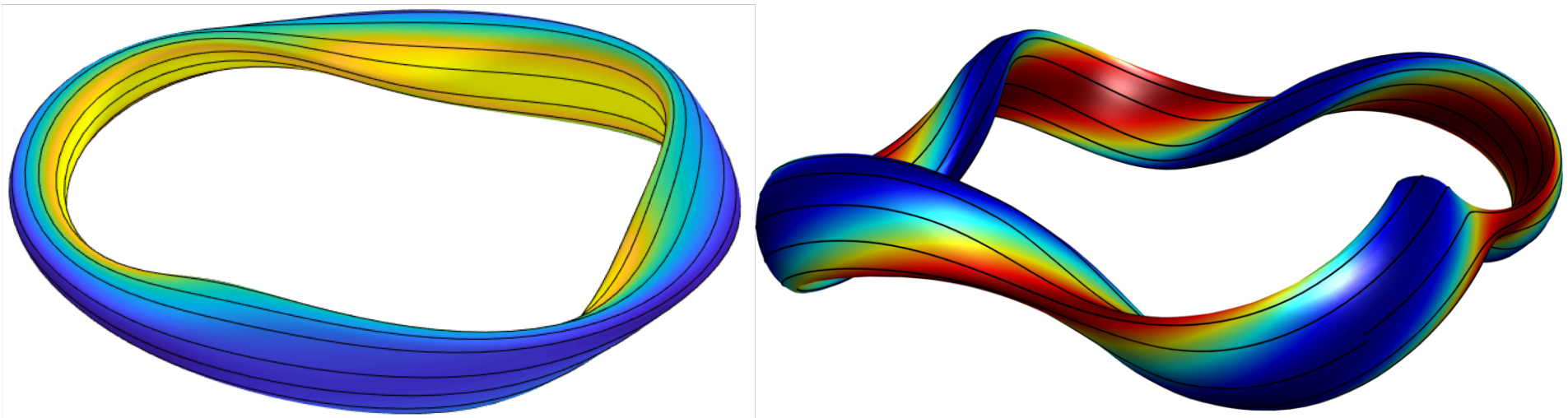


Confining charged particle orbits using hidden symmetry

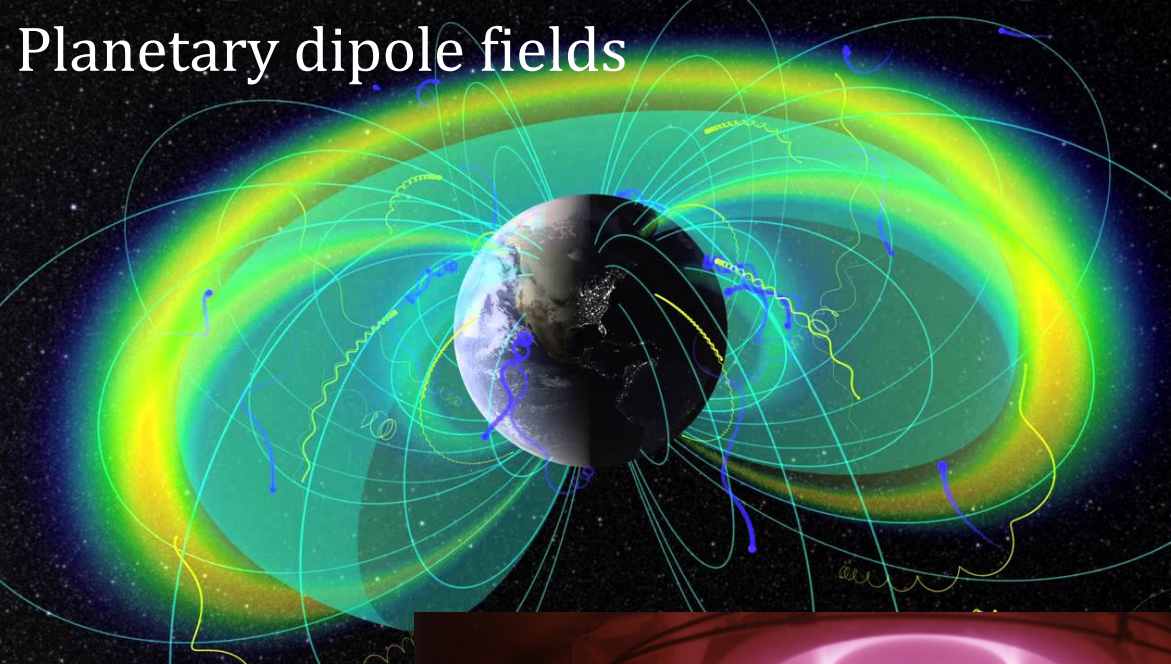


Matt Landreman, UMD IREAP

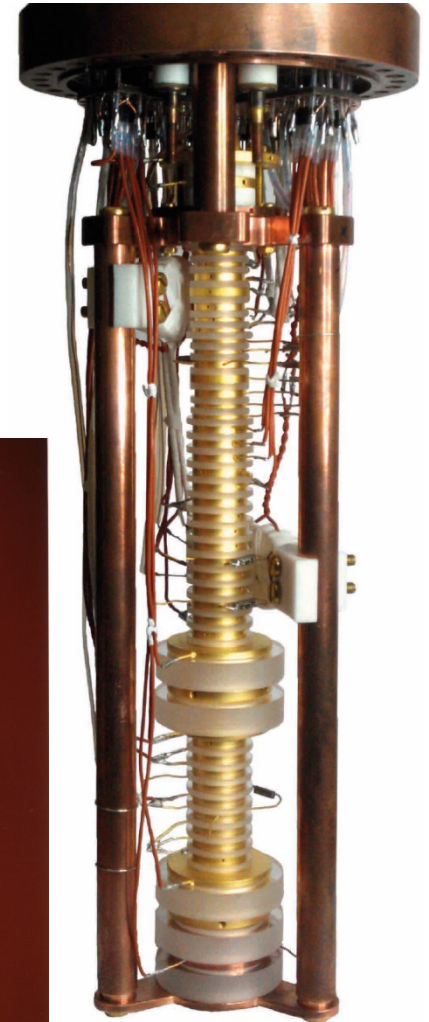
Outline

- Magnetic confinement, & pros/cons of axisymmetry
- Integrability of non-axisymmetric **B** fields
- Quasi-symmetry
- Finding quasi-symmetric fields

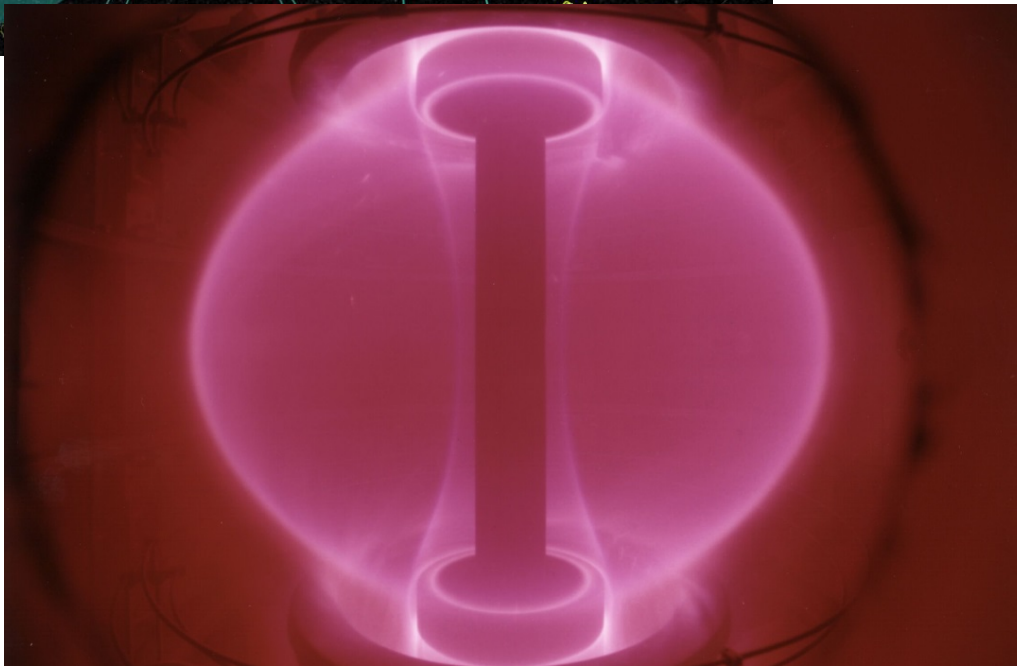
Charged particles can be confined by magnetic fields in many contexts.



Particle traps for basic physics:

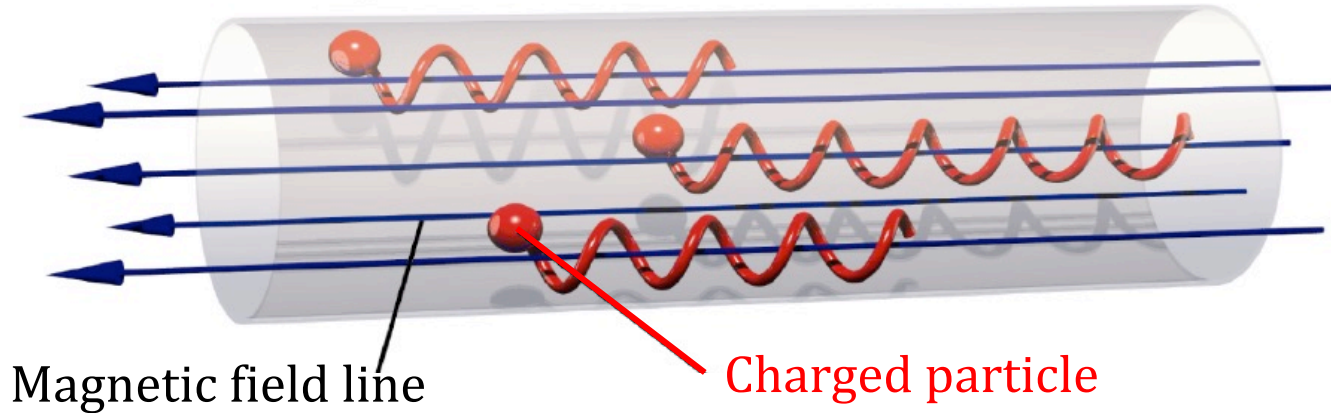


Hot
laboratory
plasmas,
fusion



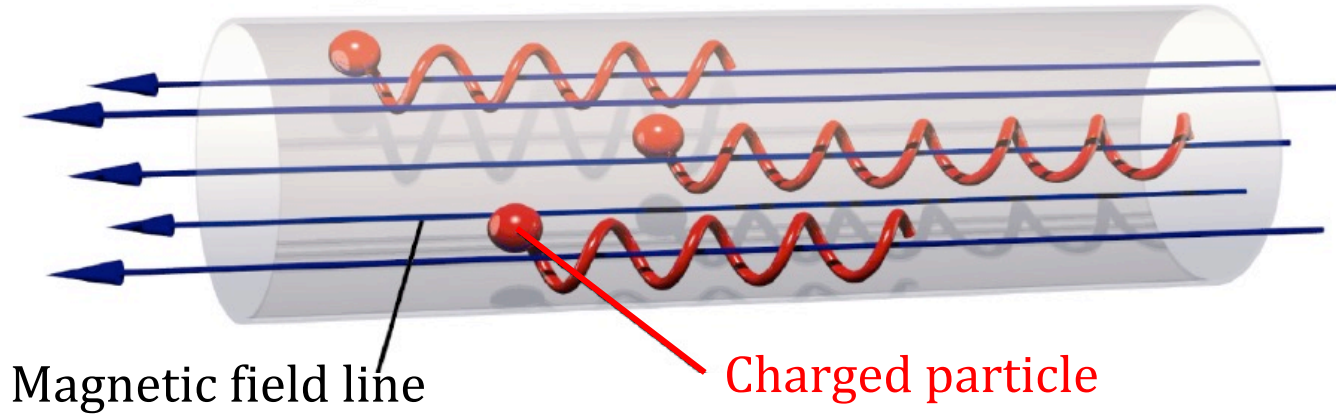
Confining charged particles with a magnetic field is tricky.

Uniform straight \mathbf{B} : confinement \perp to \mathbf{B} , but end losses.



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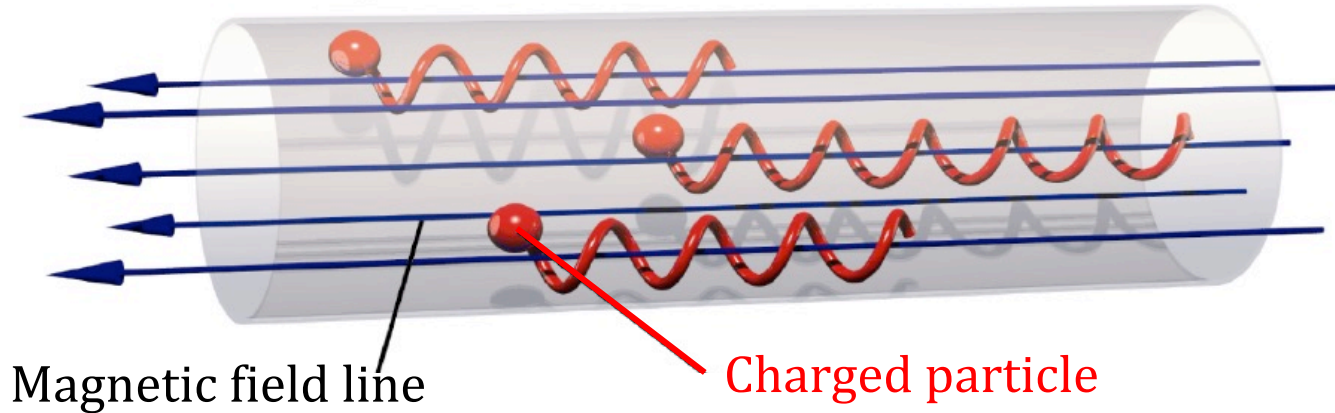


But if field lines are bent, particles drift off them.

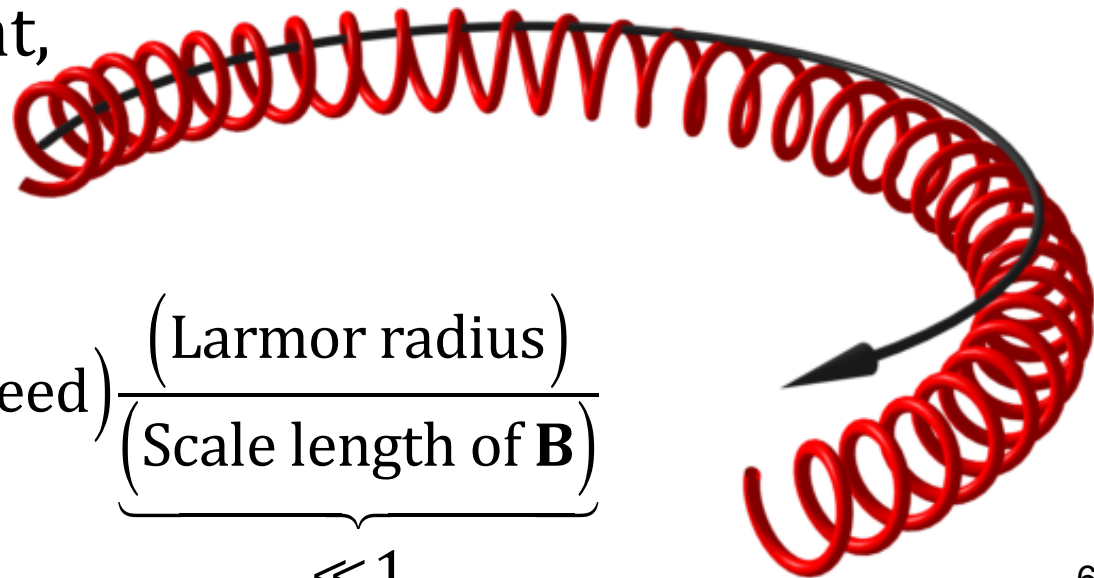


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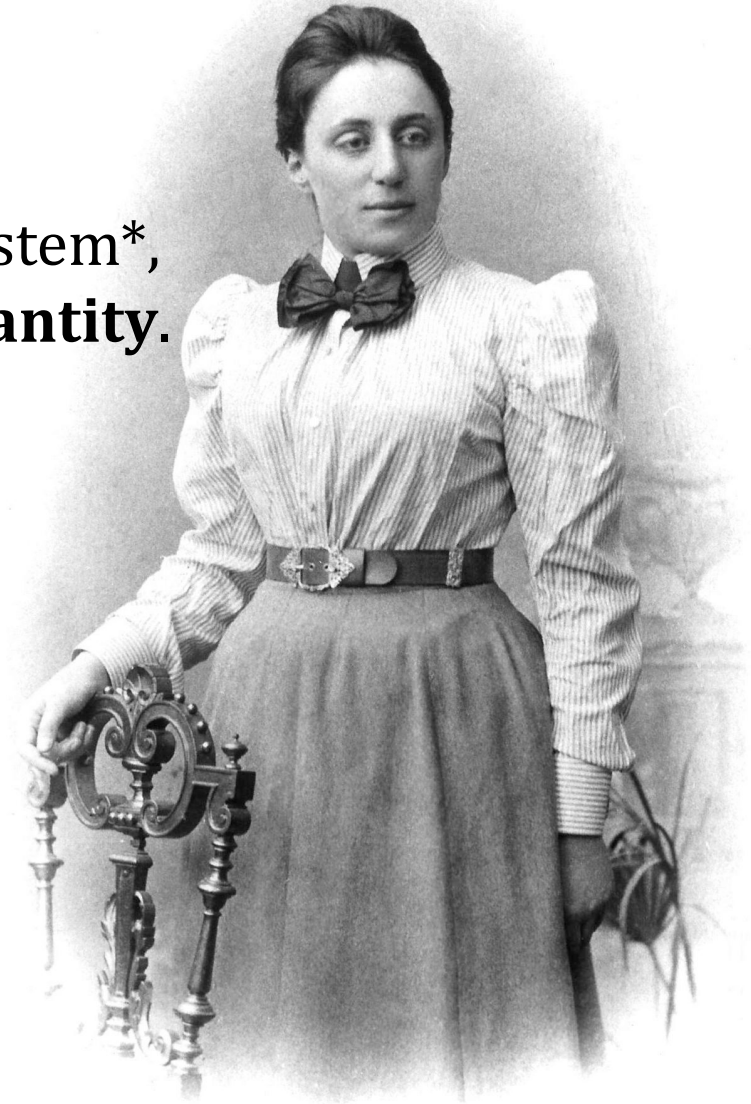
$$\text{(Drift velocity)} \sim \underbrace{\text{(Particle speed)}}_{\ll 1} \frac{\text{(Larmor radius)}}{\text{(Scale length of } \mathbf{B} \text{)}}$$

To confine particles, we can constrain their position with a conservation law.

Noether's theorem:

For each **continuous symmetry** of a system*, there is a corresponding **conserved quantity**.

* For this talk: Lagrangian is independent of a coordinate.



Emmy Noether (1882-1935) 7

Axisymmetry + Noether's Theorem is one way to achieve magnetic confinement.

Continuous rotational symmetry \Rightarrow Canonical angular momentum is conserved.

$$L_\phi = mv_\phi R + qA_\phi R = \text{constant}$$

 vector potential: $\mathbf{B} = \nabla \times \mathbf{A}$

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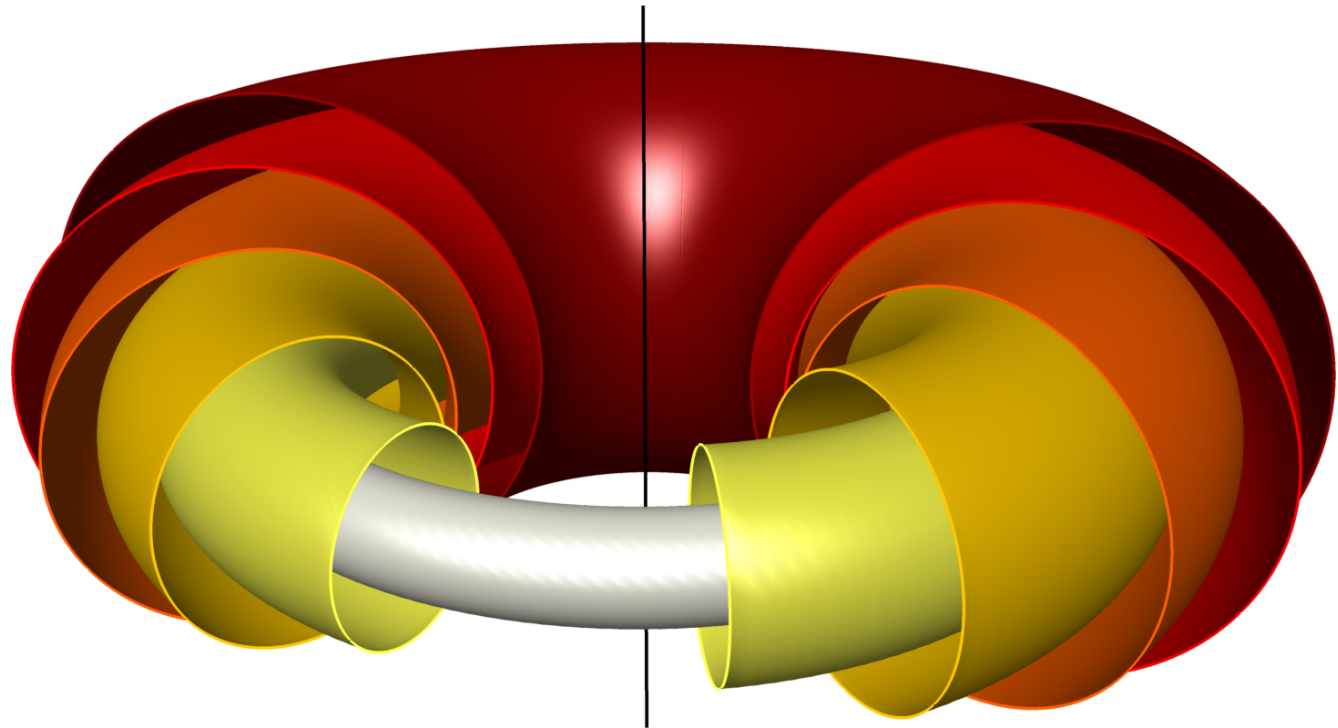
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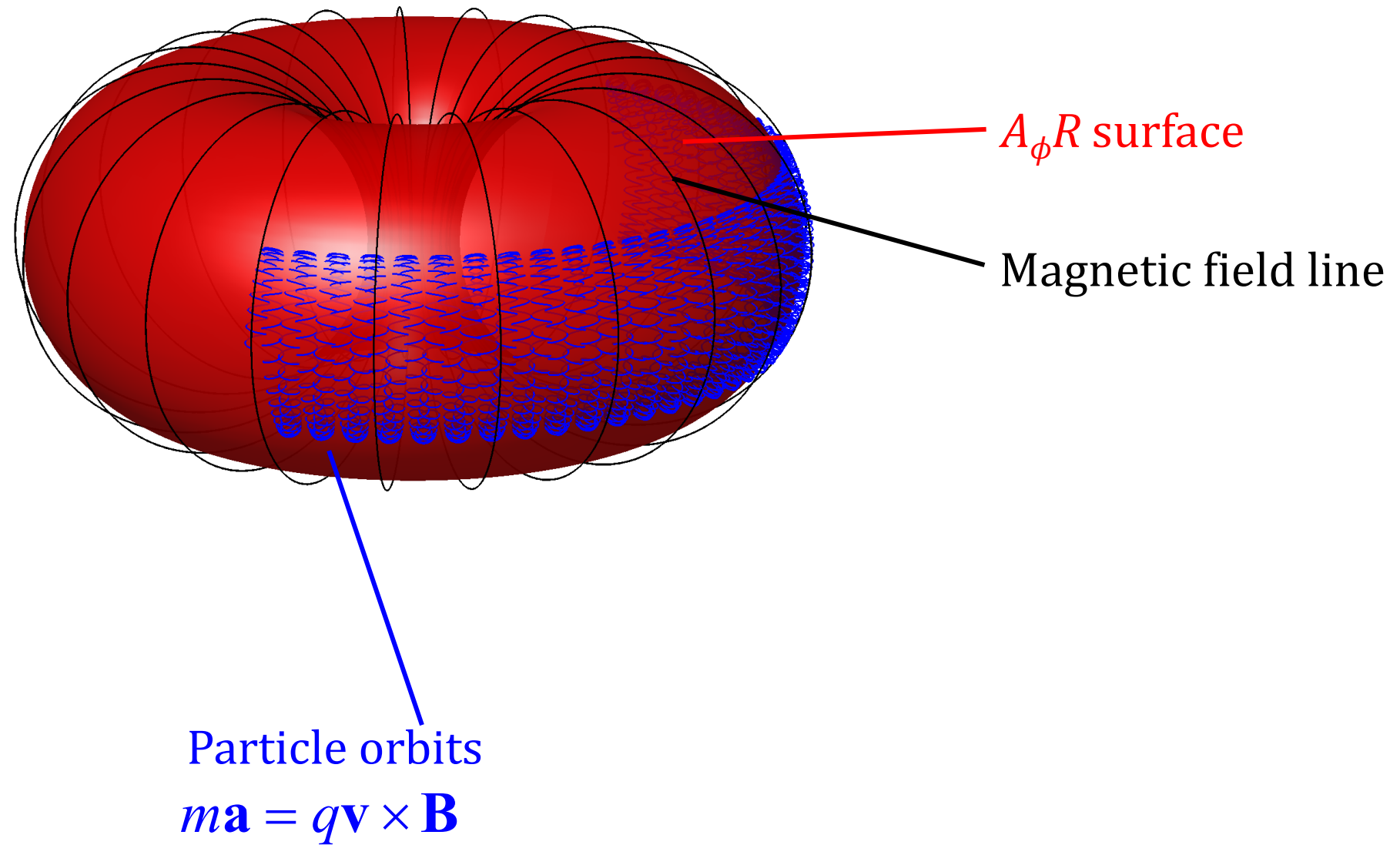
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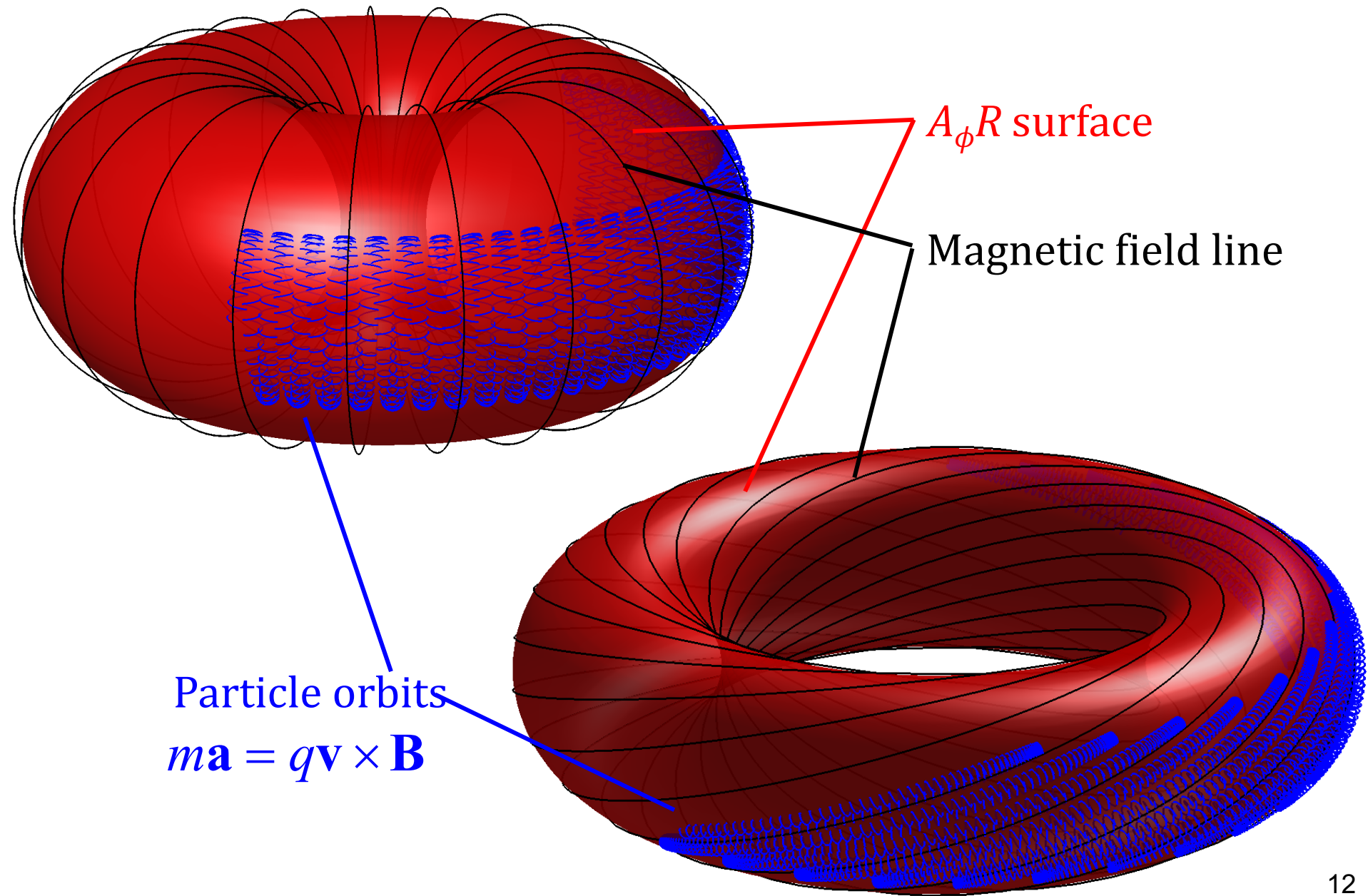
If $A_\phi R$ surfaces are bounded like this, then particles will be confined:



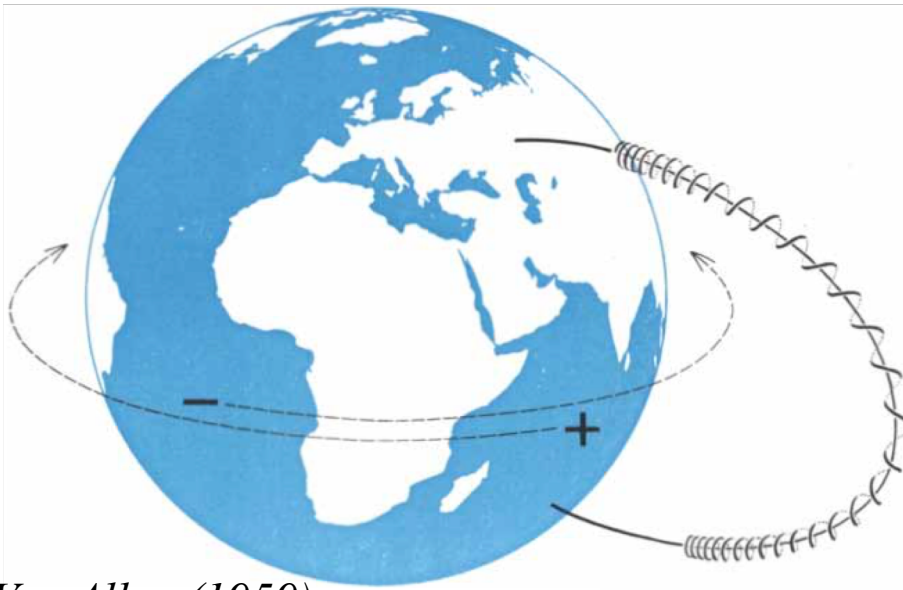
In axisymmetry, particles are confined (close) to $A_\phi R$ surfaces, despite complicated orbits.



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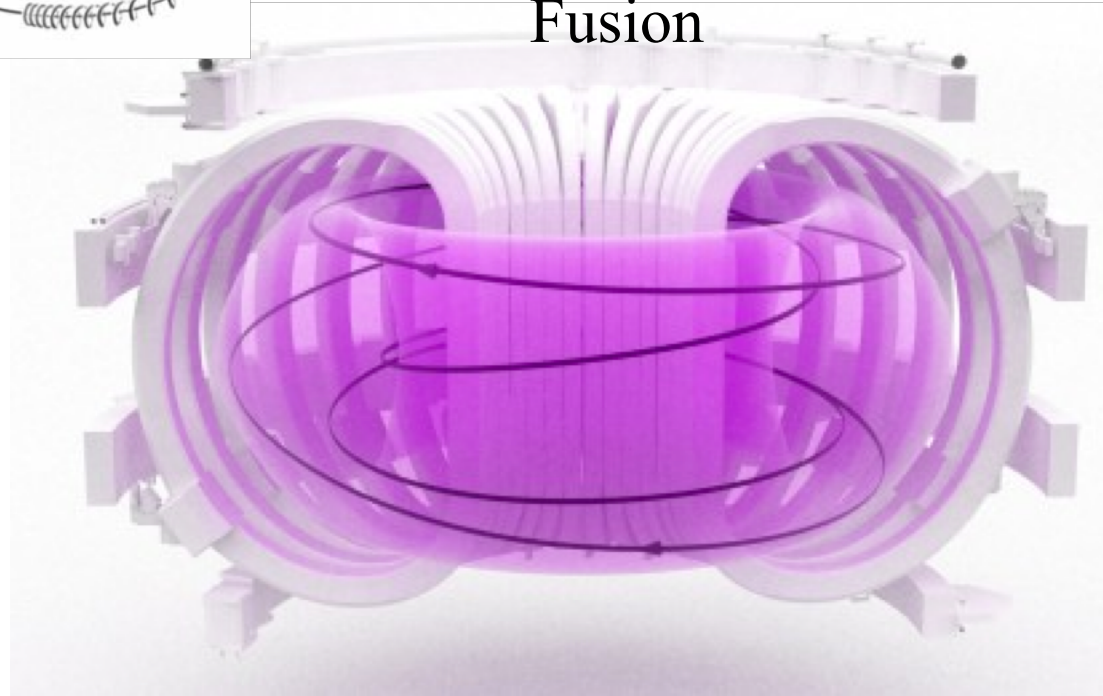


**Particles are actually confined this way
in nature and in the laboratory.**



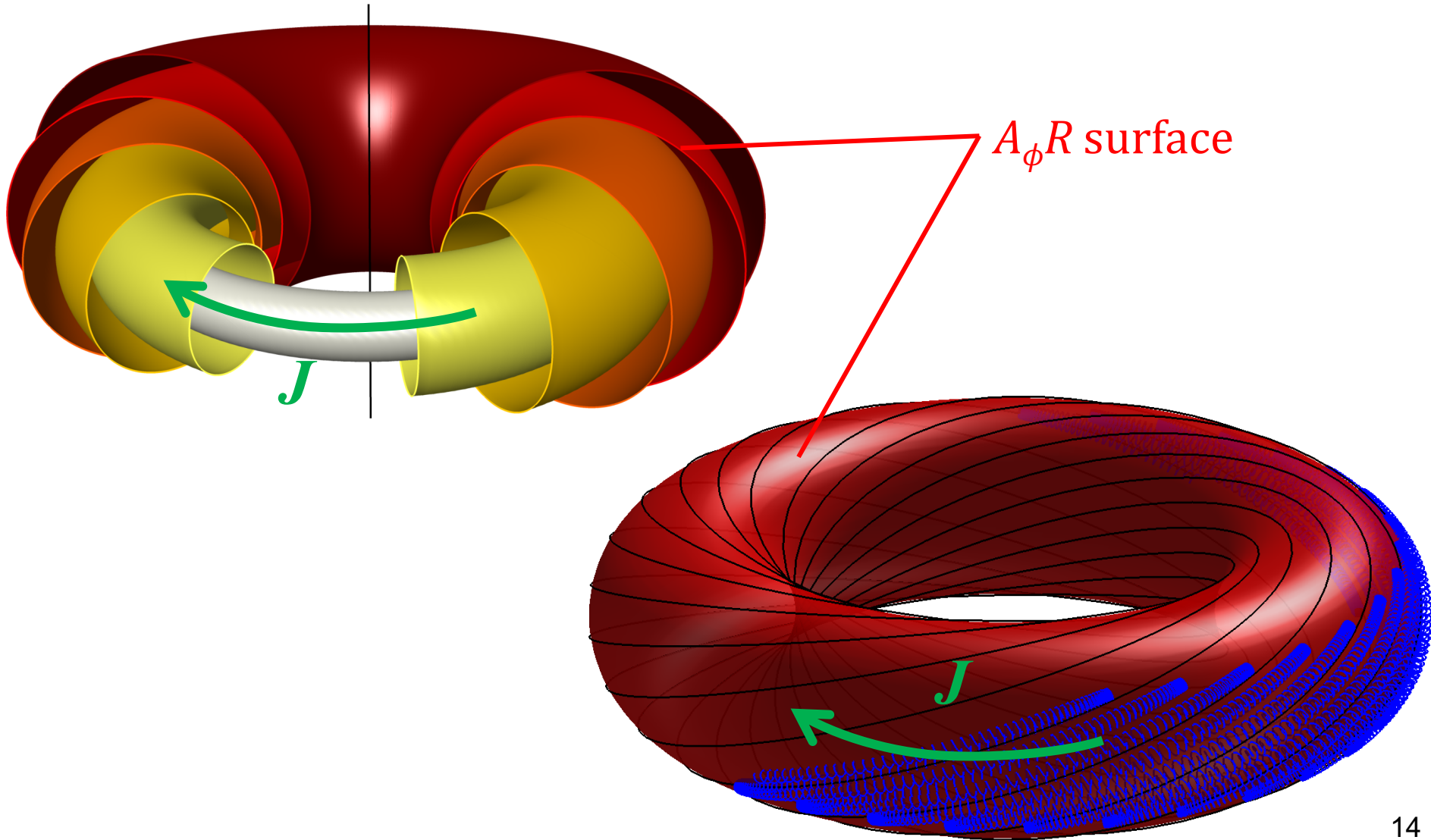
Van Allen (1959)

Fusion



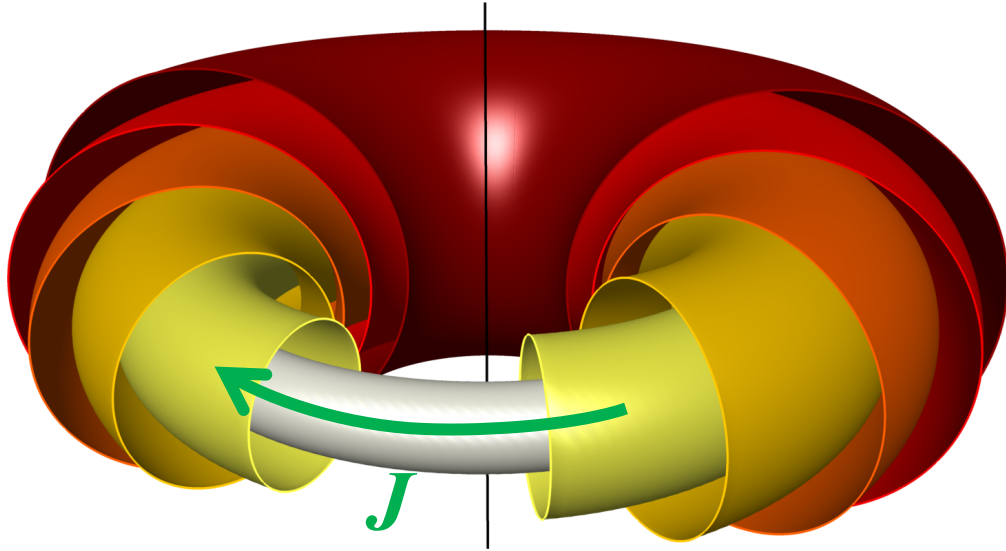
**But, axisymmetric confinement has a big problem:
requires an internal current.**

$\nabla \times (\nabla \times \mathbf{A}) = \nabla \times \mathbf{B} = \mu_0 \mathbf{J}$ so nested $A_\phi R$ surfaces require a J_ϕ .



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$\nabla \times (\nabla \times \mathbf{A}) = \nabla \times \mathbf{B} = \mu_0 \mathbf{J}$ so nested $A_\phi R$ surfaces require a J_ϕ .



- Sustaining this current in steady-state is hard.
- This current drives instabilities.
- Not possible for low plasma density.

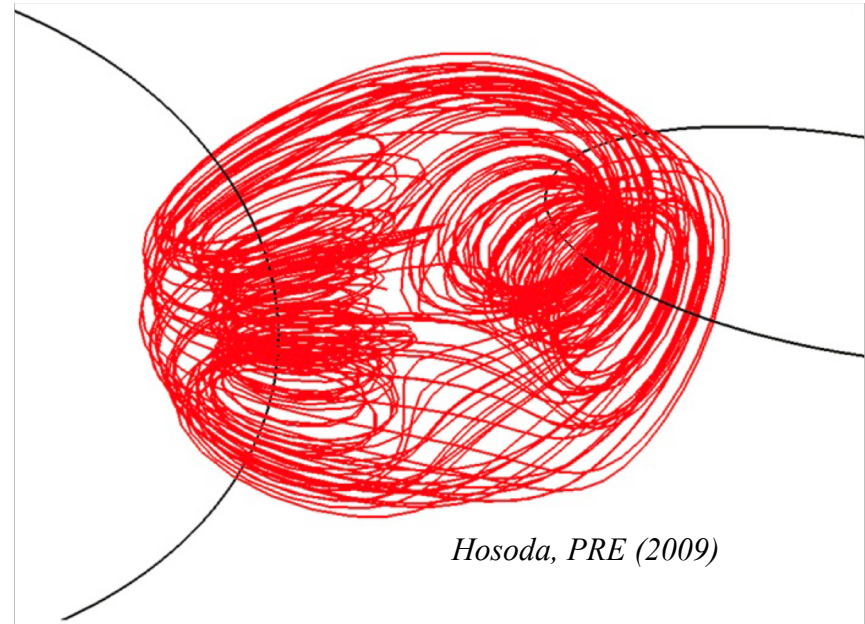
Can we achieve similar confinement without axisymmetry to avoid these problems?

Outline

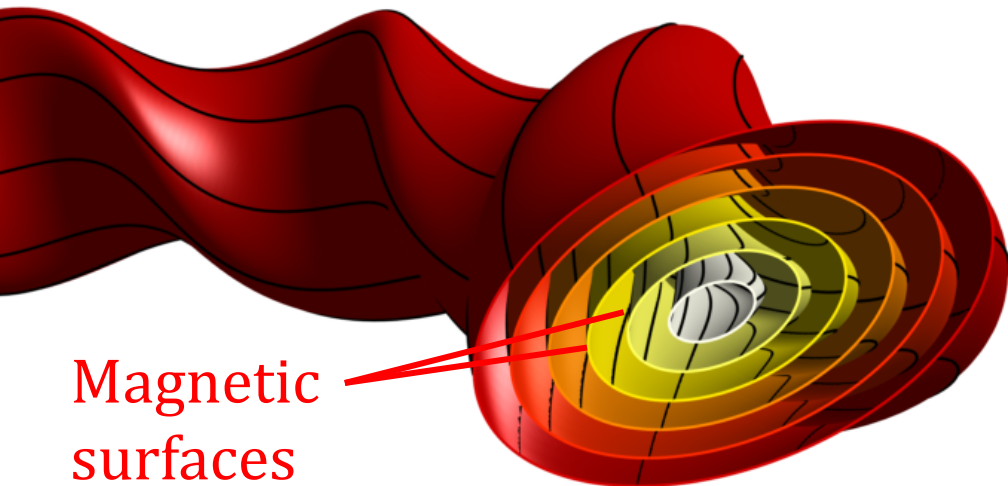
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When axisymmetry is broken, we want field lines to still lie on surfaces.

BAD: Particle motion along \mathbf{B} allows inside & outside to mix even without cross- \mathbf{B} drift.

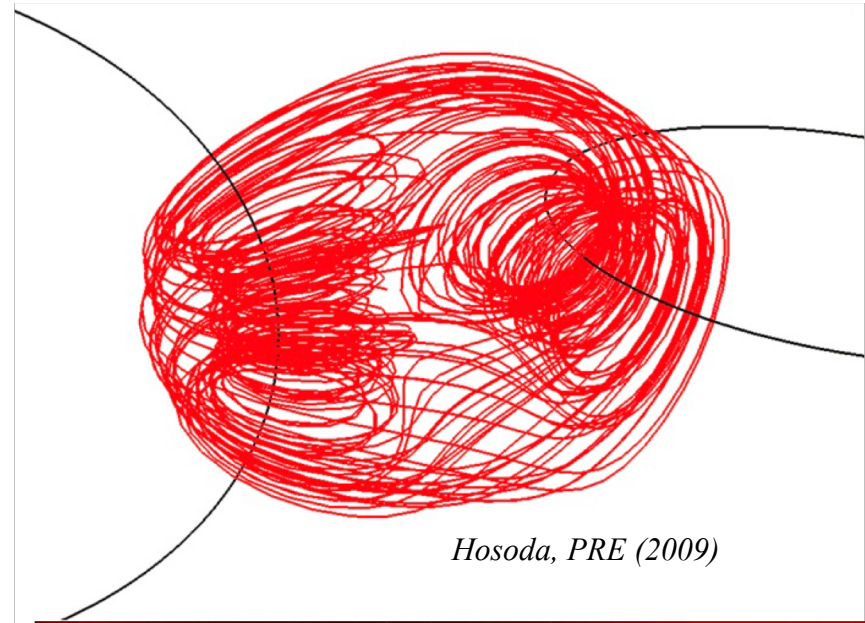


GOOD: \mathbf{B} is “integrable”



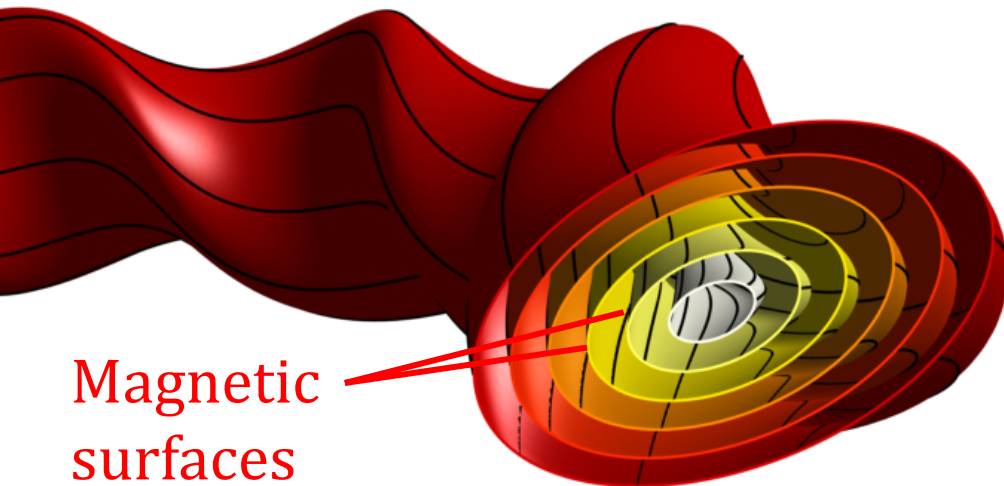
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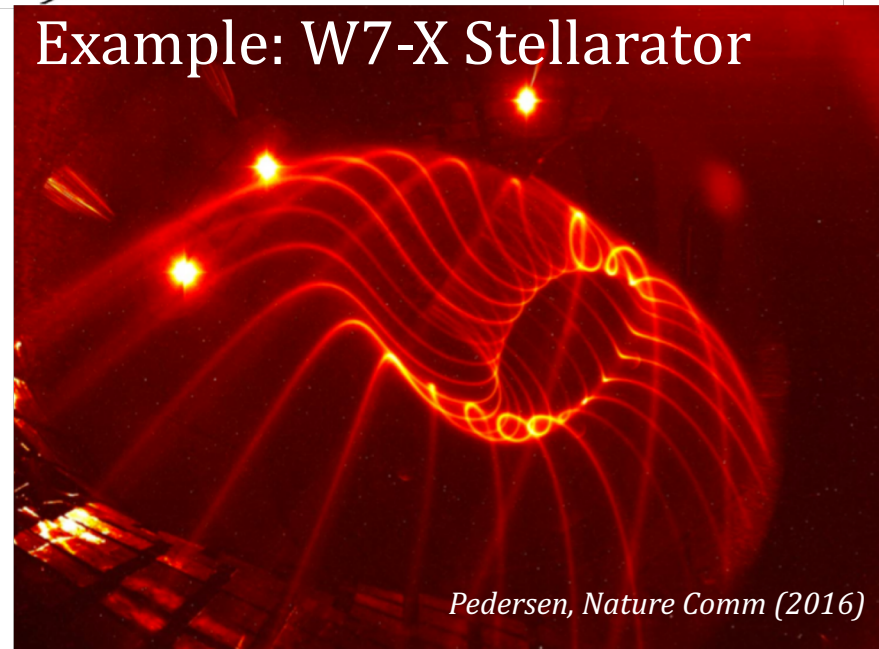
Hosoda, PRE (2009)

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Magnetic surfaces

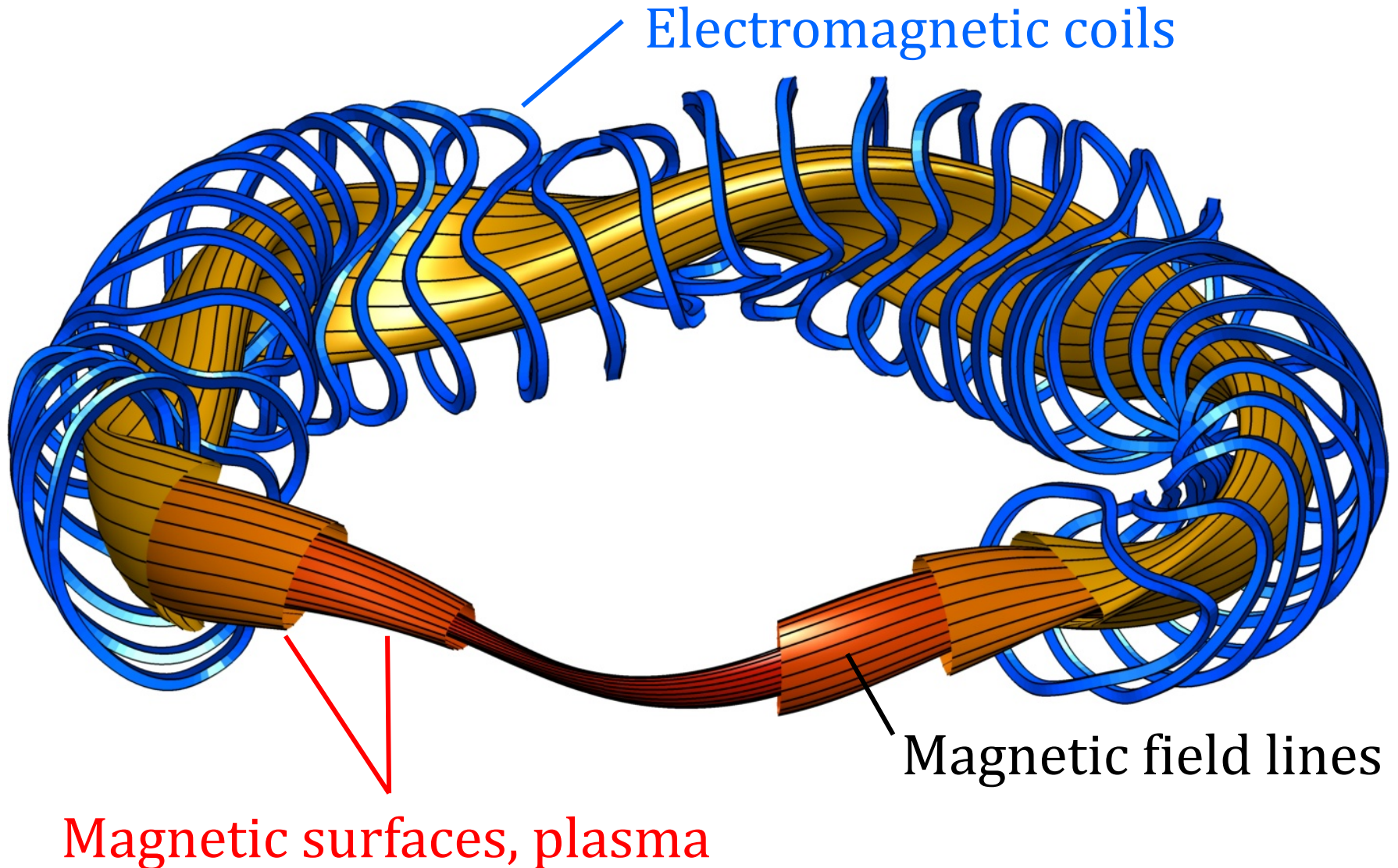
Example: W7-X Stellarator



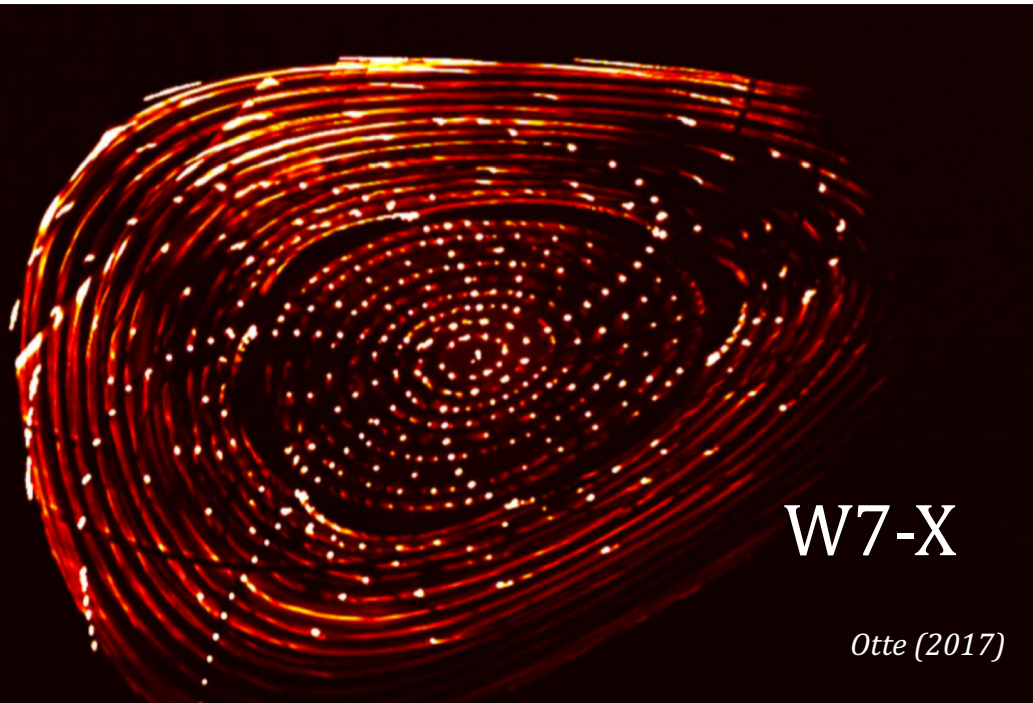
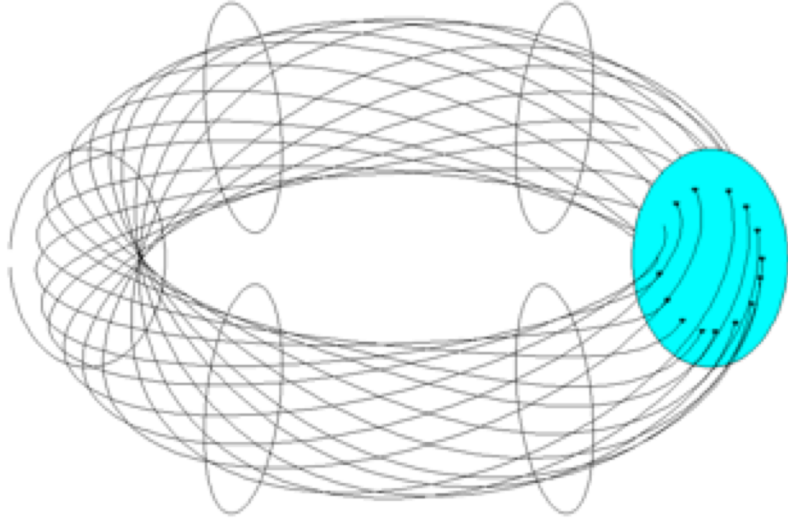
Pedersen, Nature Comm (2016)

A magnetic confinement device with a non-axisymmetric but integrable magnetic field is a "stellarator"

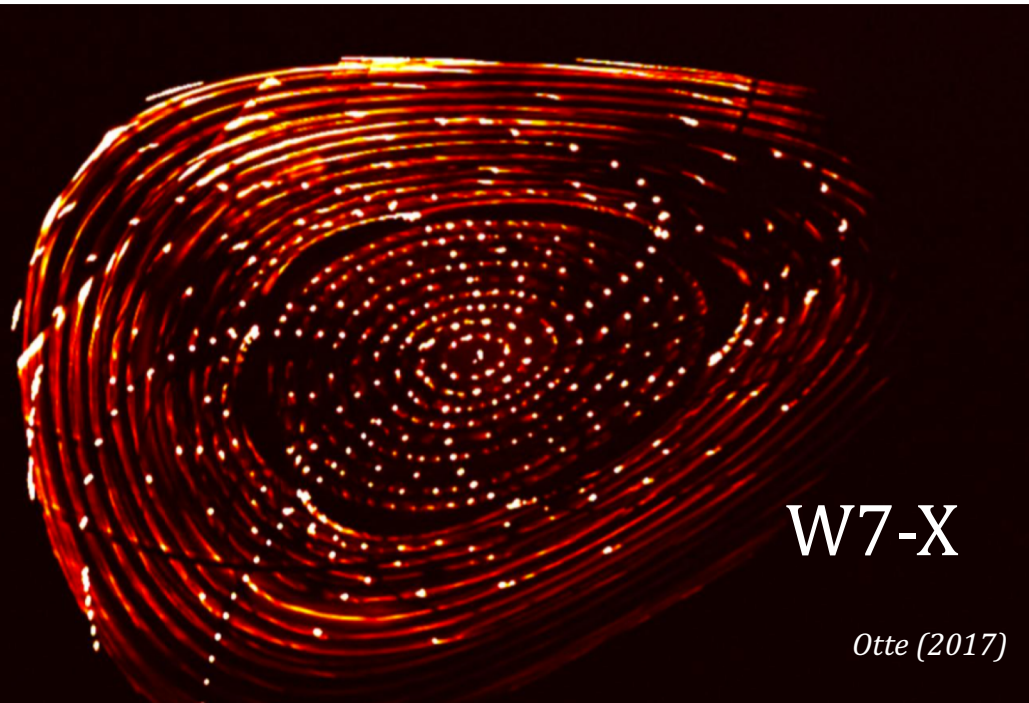
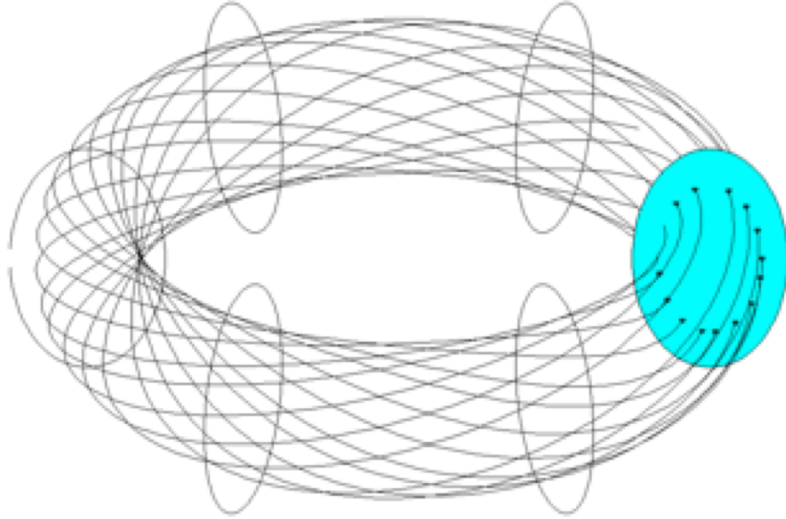
E.g. Wendelstein 7-X (Germany):



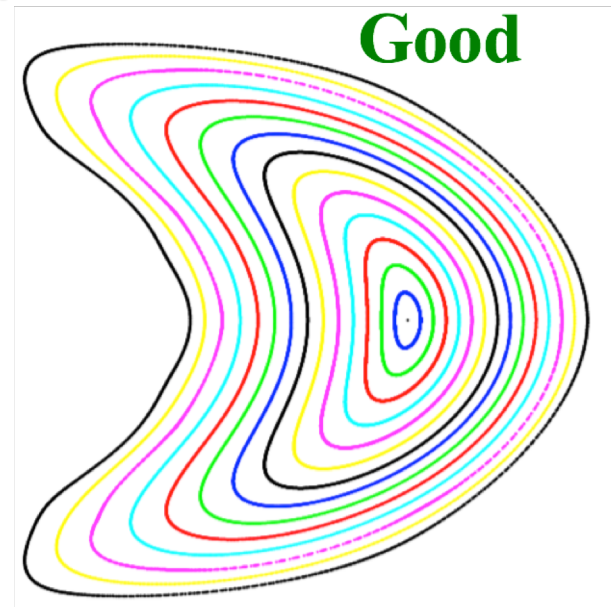
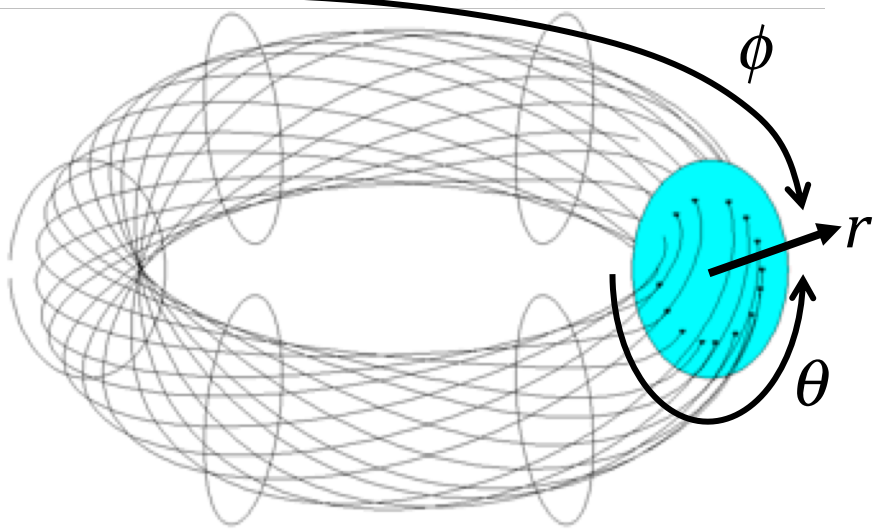
Integrability of magnetic fields can be viewed using Poincare plots



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Magnetic field lines can be described by a Hamiltonian, where “time” is ϕ



$$\frac{d\theta}{d\phi} = \frac{\partial H}{\partial r}, \quad \frac{dr}{d\phi} = -\frac{\partial H}{\partial \theta}$$

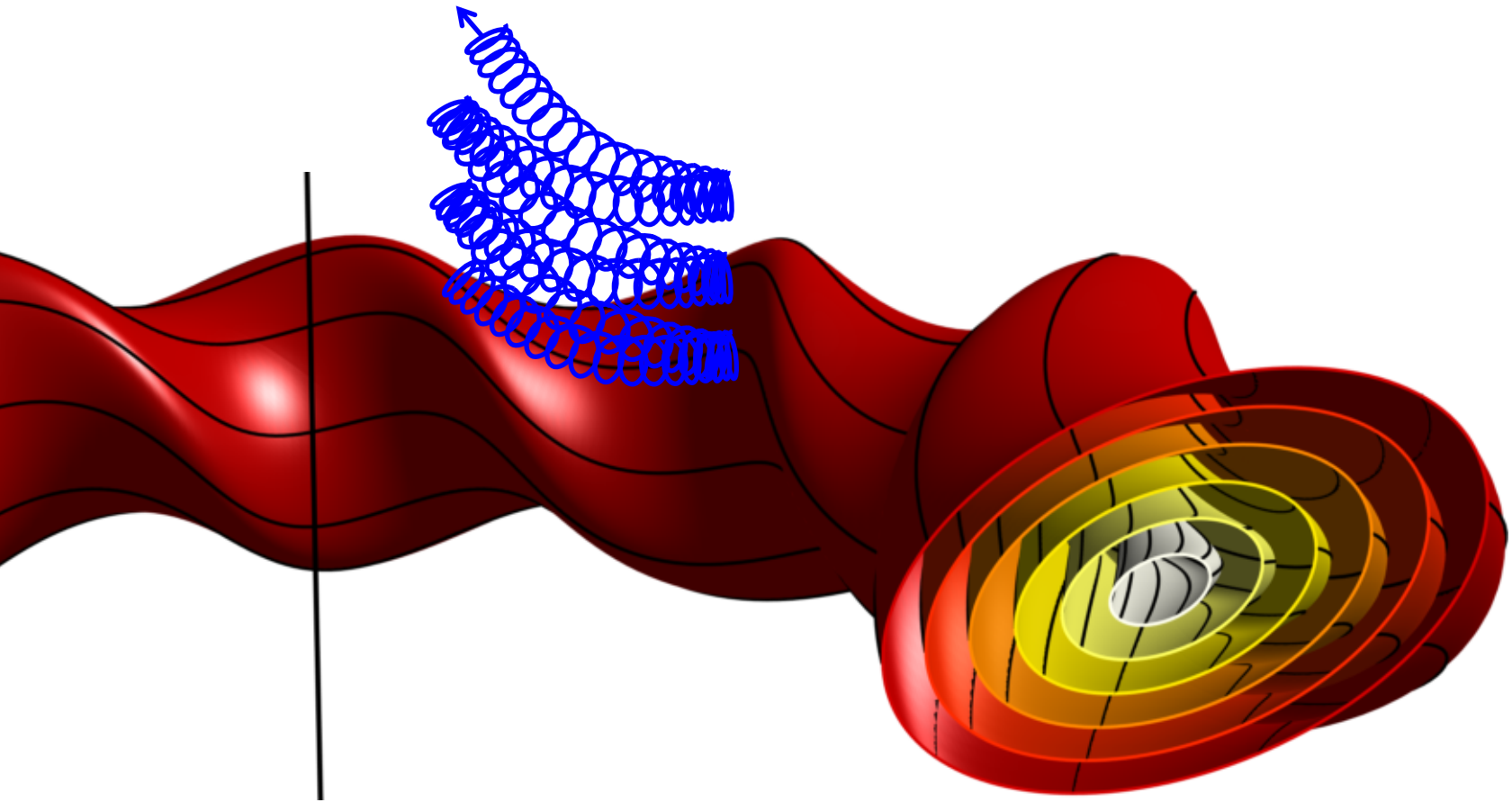
So tools from Hamiltonian systems like KAM apply.



Outline

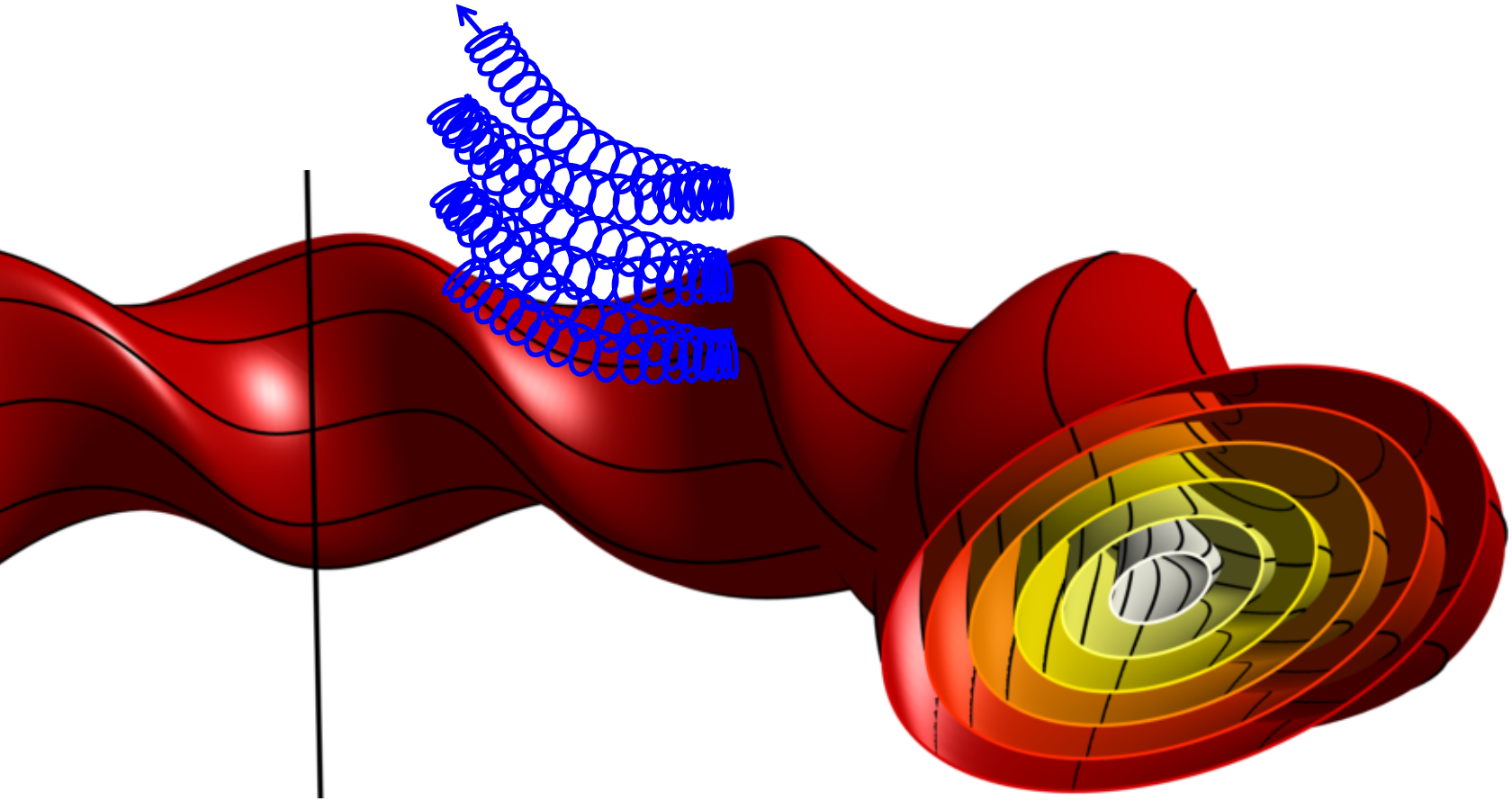
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“Quasi-symmetry” provides a solution.

Quasi-symmetry is a continuous symmetry in $|\mathbf{B}|$ (not vector \mathbf{B}) that implies confinement.

- When the Lagrangian is (1) expanded for large $B=|\mathbf{B}|$ and (2) written in a special coordinate system (“Boozer angles”), it depends on position only through the surface and B .

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- Therefore a symmetry in B implies a conserved quantity, even if \mathbf{B} has no obvious symmetry.
- This conserved quantity resembles canonical angular momentum, so it implies confinement just as in axisymmetry.

Averaging over gyration, Lagrangian depends on \mathbf{B} only through surface and $|\mathbf{B}|$.

Lagrangian for particle in magnetic field: $\mathcal{L} = q\mathbf{A} \cdot \dot{\mathbf{x}} + \frac{m}{2} |\dot{\mathbf{x}}|^2$ (Neglect \mathbf{E})

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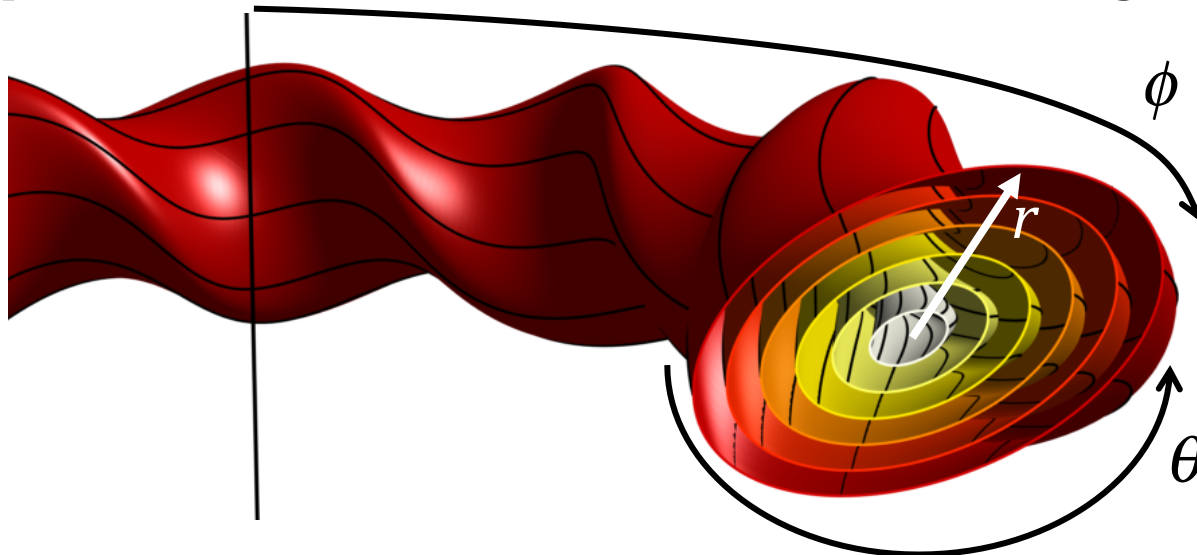
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Depends only on r

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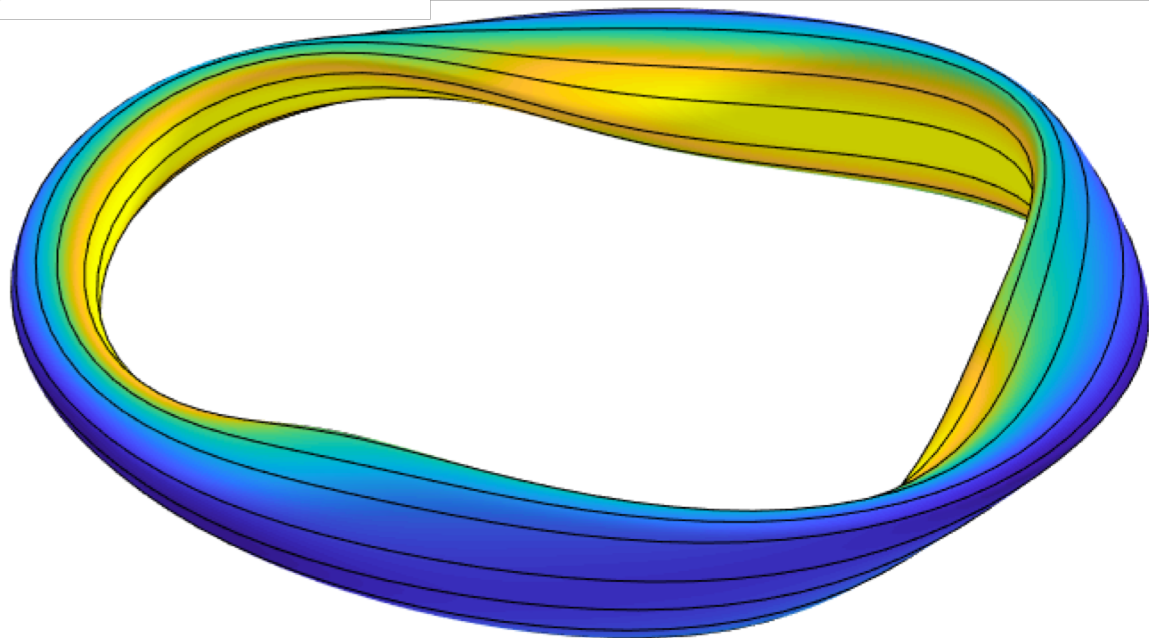
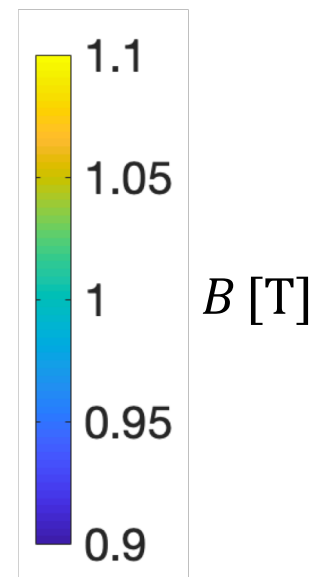
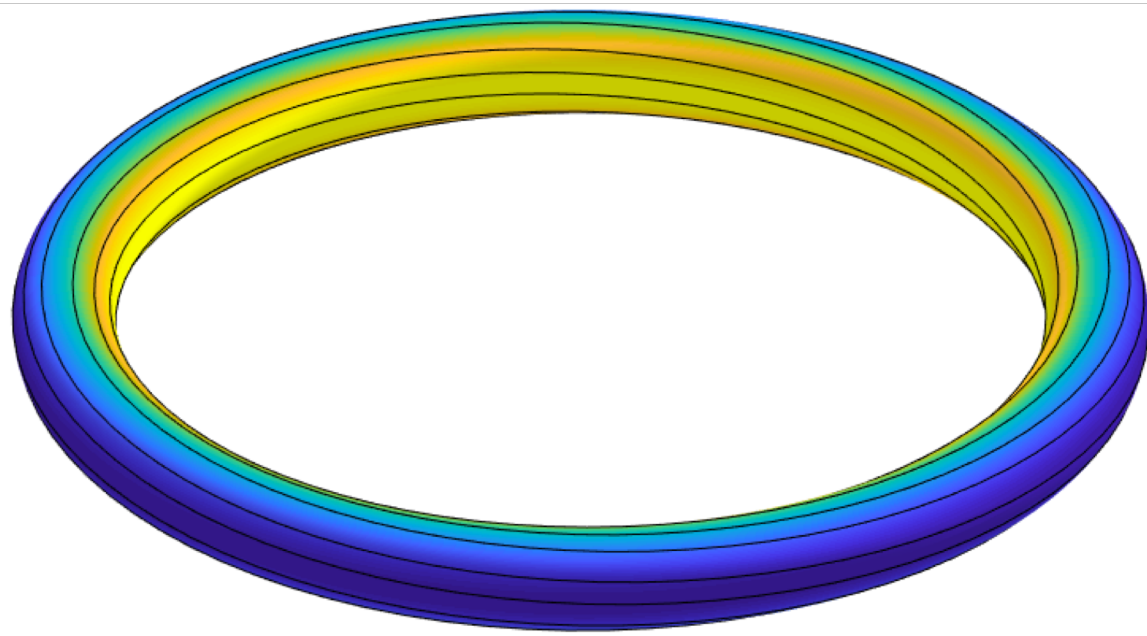
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Depends only on r

$$\frac{\partial B}{\partial \phi} = 0 \Rightarrow \text{Conservation of } \frac{\partial \mathcal{L}}{\partial \dot{\phi}} = -q\psi_p + \frac{mv_{\parallel} B_{\phi}}{B} \approx -q\psi_p \Rightarrow \text{Confinement!}$$

Due to quasisymmetry, different B fields can have isomorphic particle orbits.

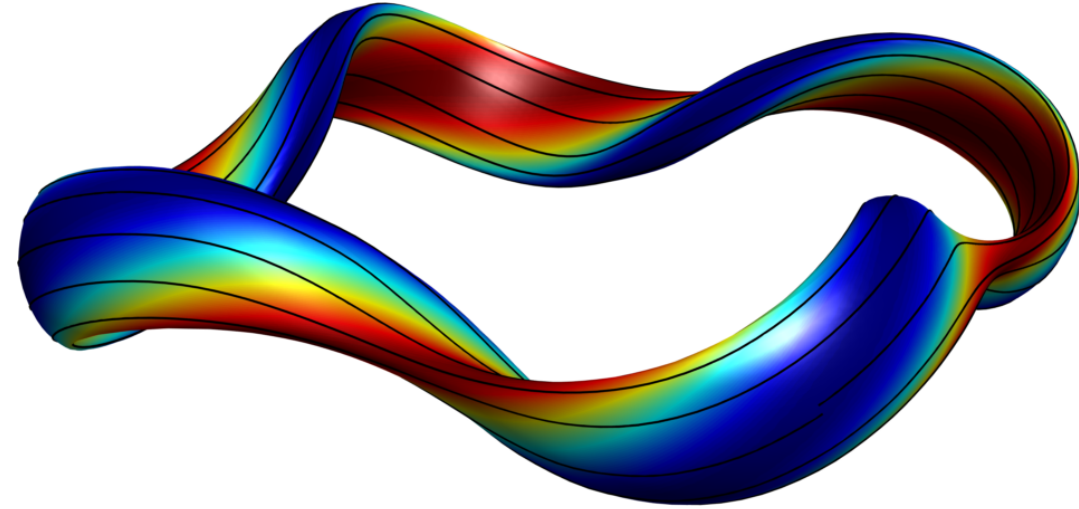


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Several quasi-symmetric confinement experiments have been designed using optimization.

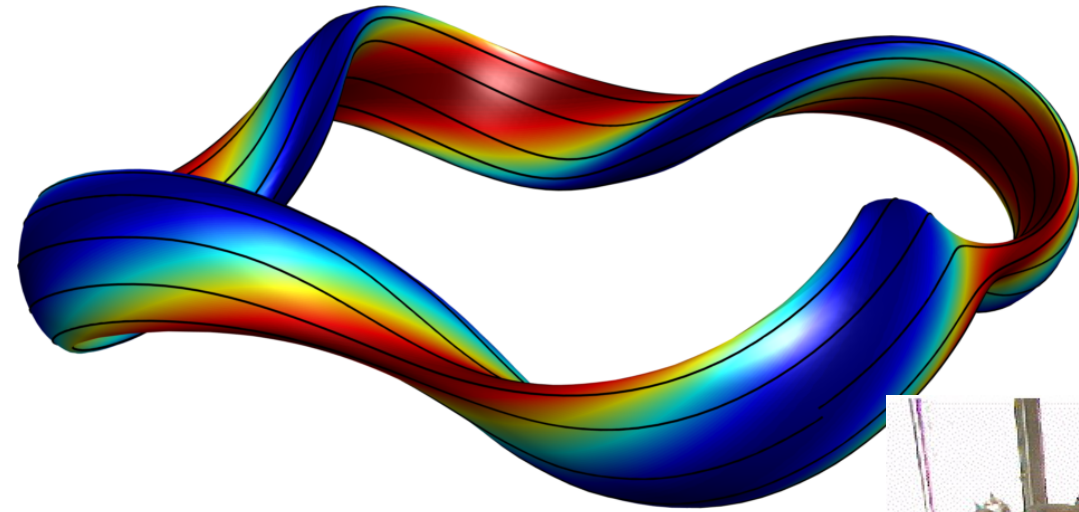
Boundary shape varied to minimize symmetry-breaking in $|B|$.



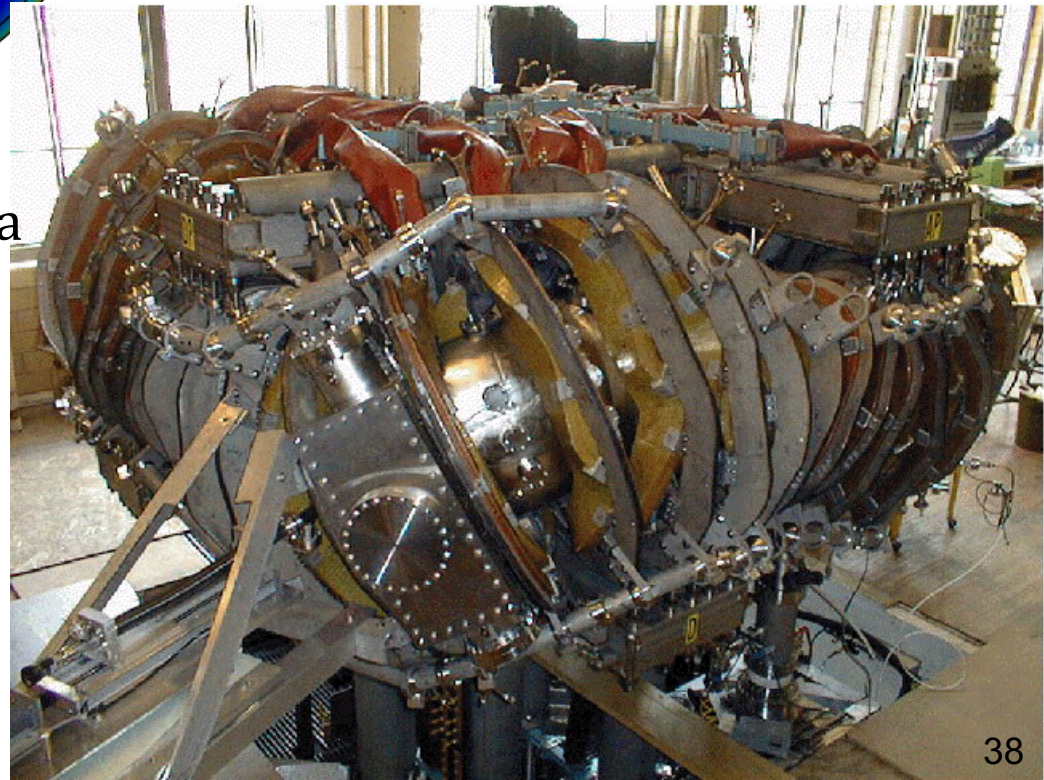
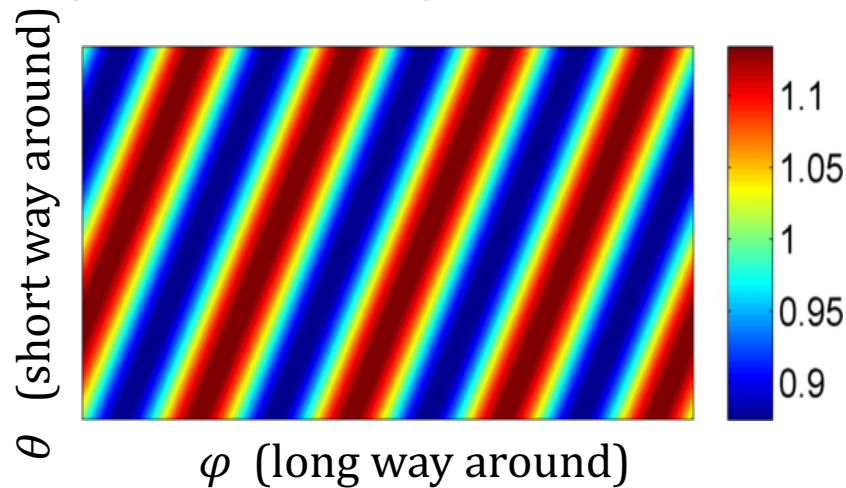
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HSX:
Helically Symmetric
eXperiment
(Univ. Wisconsin)



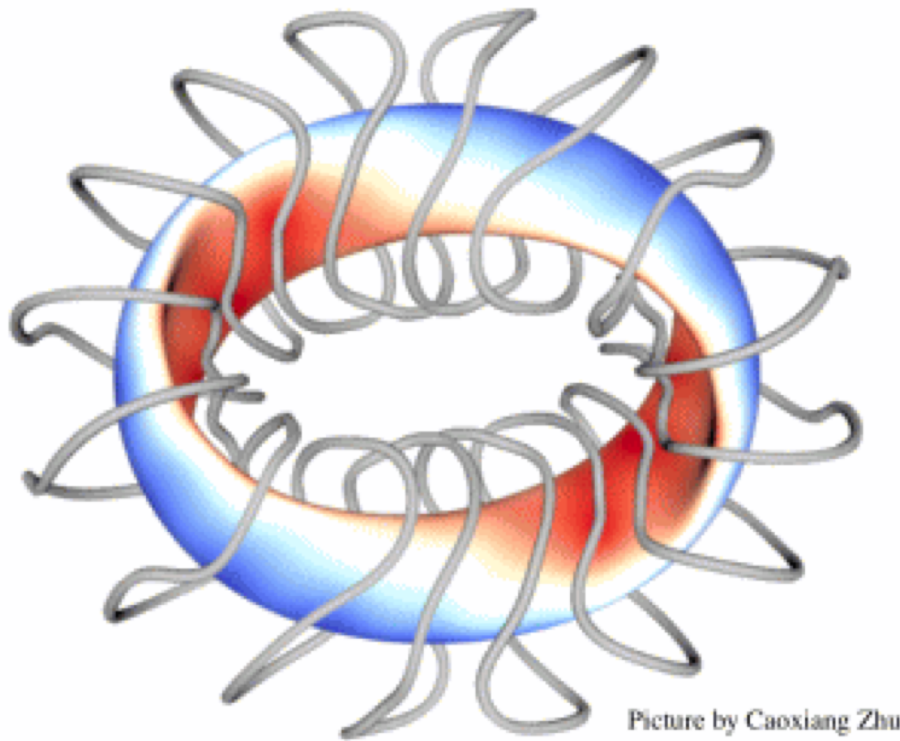
Magnetic field magnitude $|B|$ in Tesla



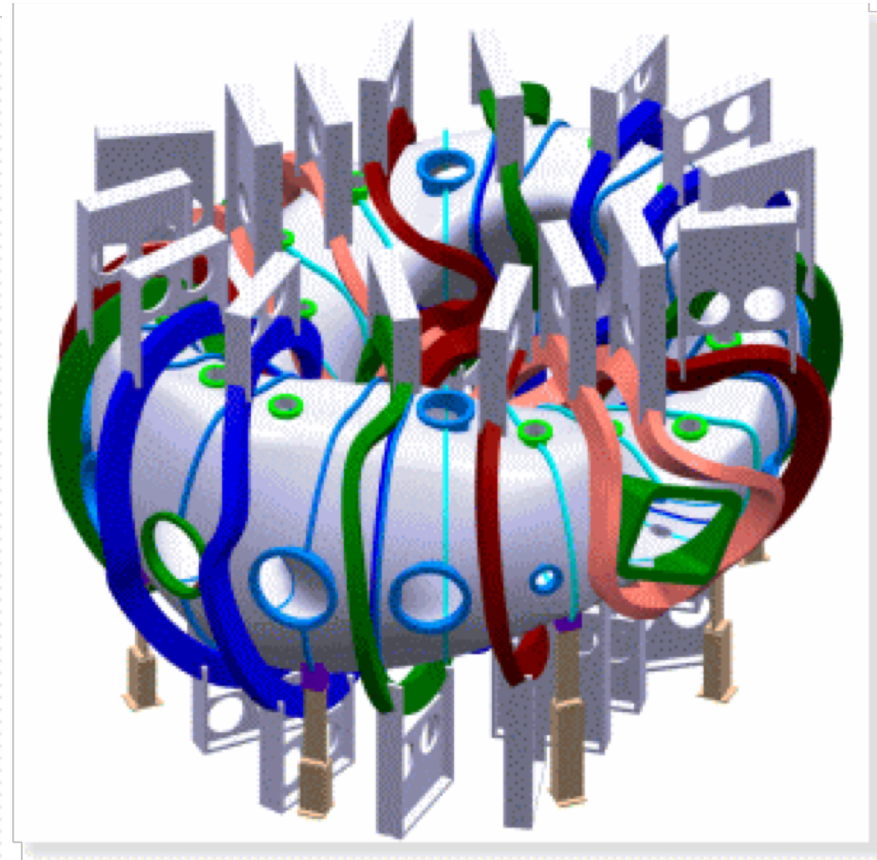
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CFQS (Chinese First Quasi-symmetric Stellarator),
Under construction



CFQS modular coil shape and plasma



CFQS, coils and vacuum chamber

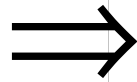
We have developed a new procedure to construct quasi-symmetric configurations.

Landreman, Sengupta, & Plunk, J. Plasma Physics (2018)

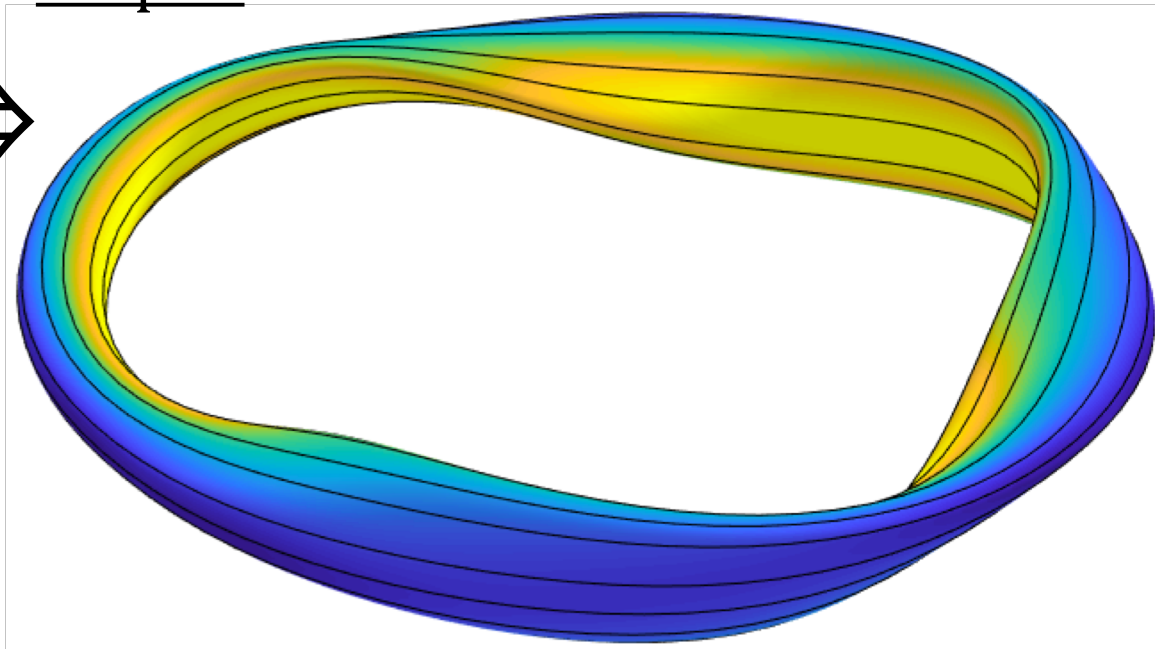
Directly solve equations for magnetohydrodynamic equilibrium
& $\partial B / \partial \phi = 0$, expanding in aspect ratio. $>10^6 \times$ faster!

Input:

Shape of the magnetic axis, & few other constants.



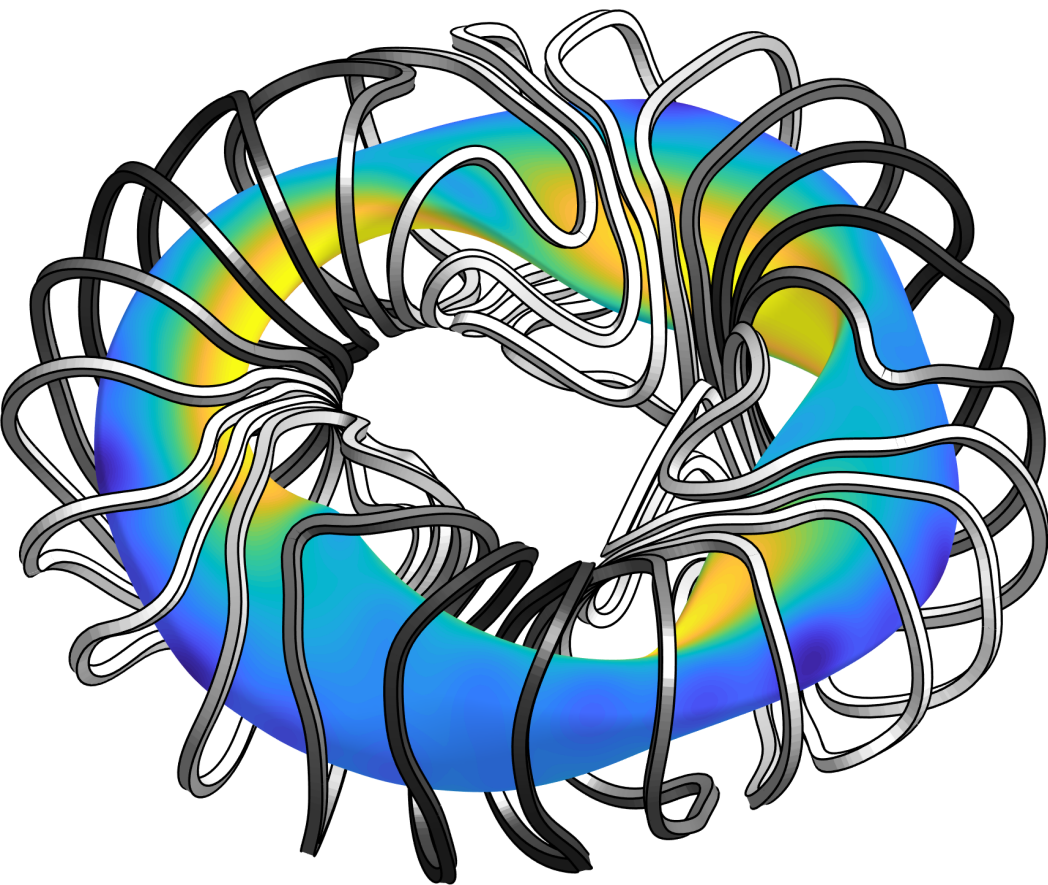
Output:



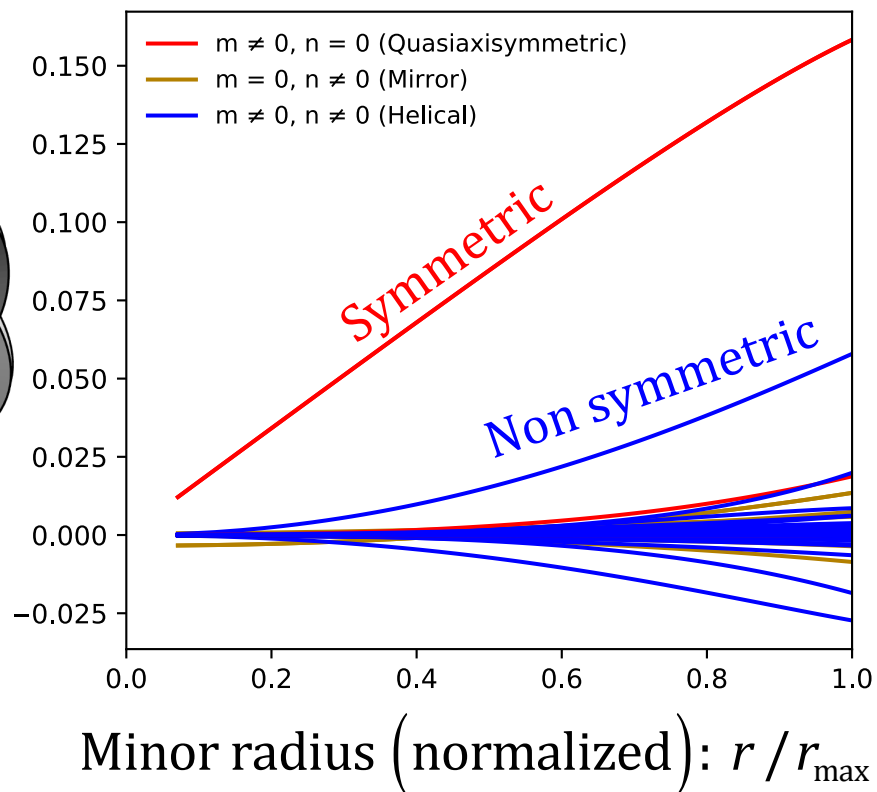
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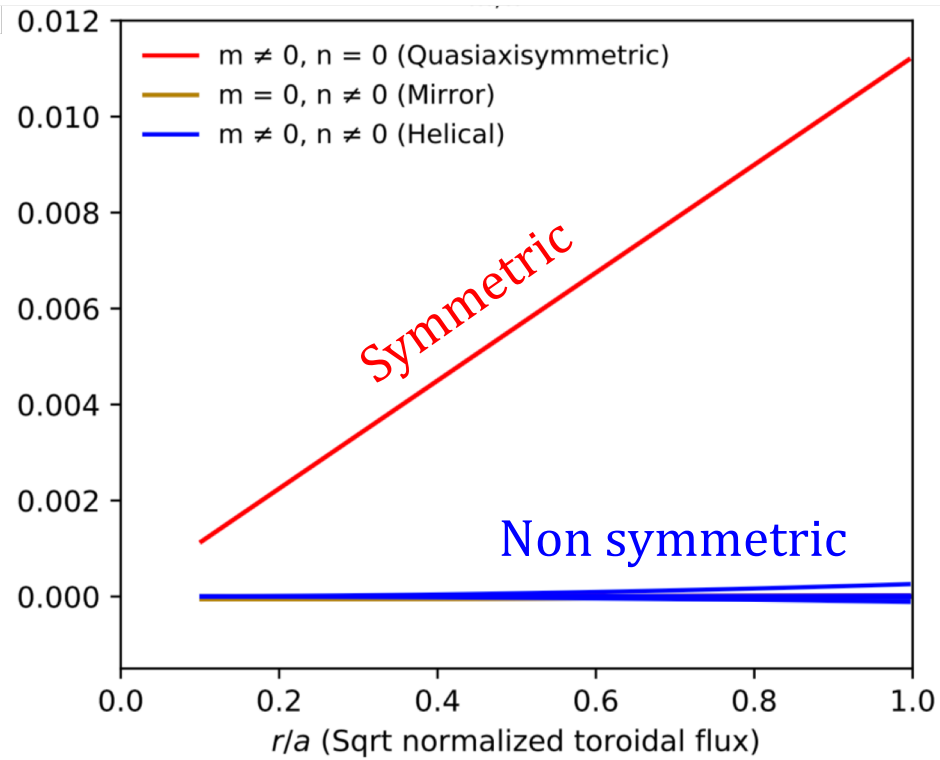
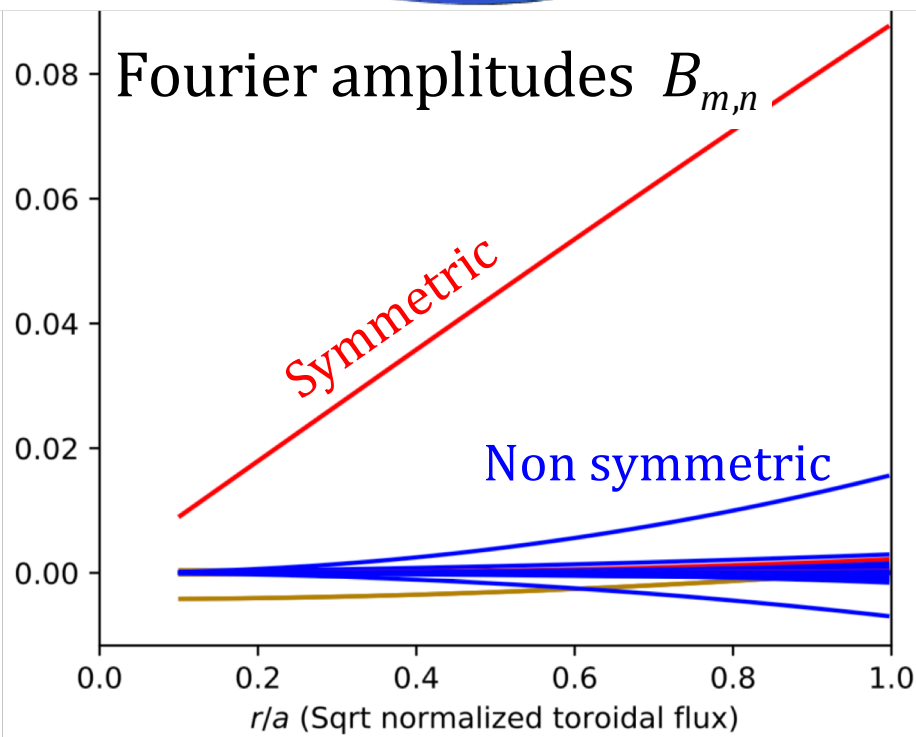
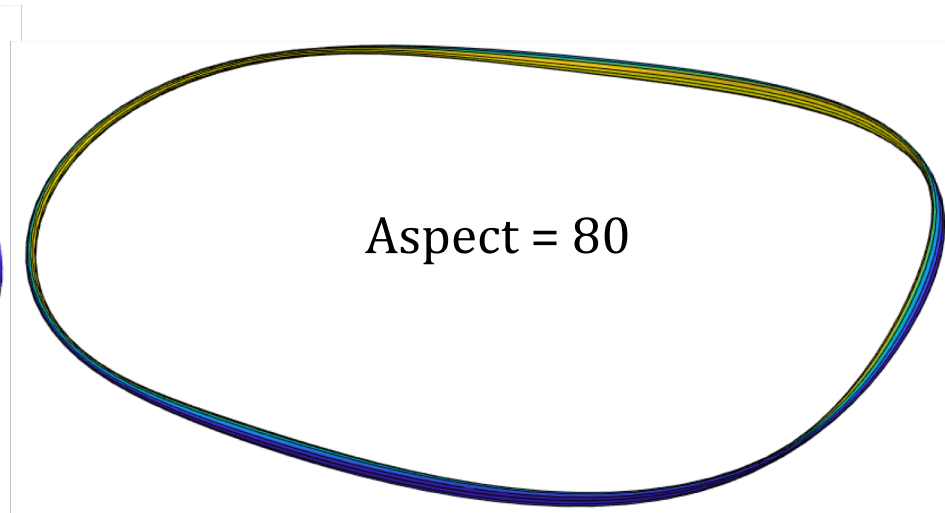
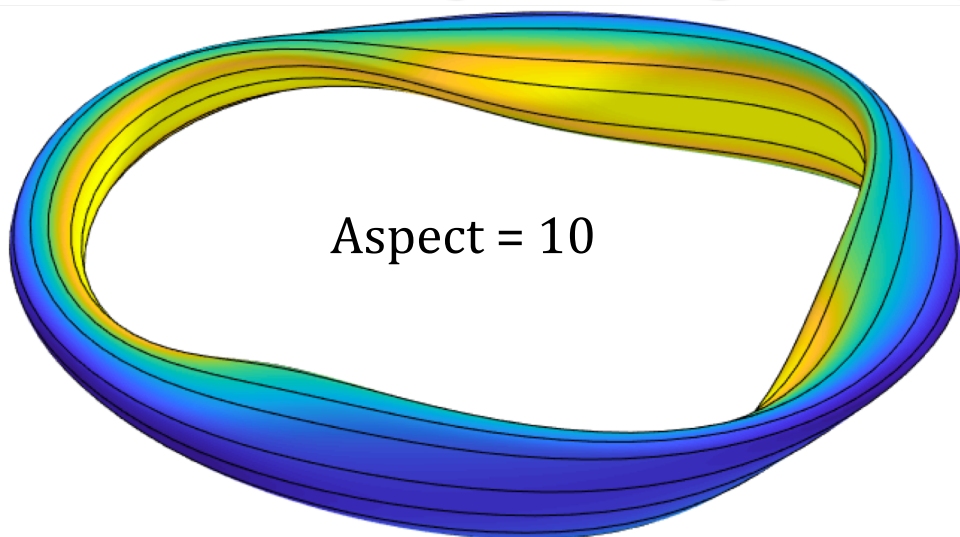
Realization with coils:



Fourier amplitudes $B_{m,n}$



Quasisymmetry can be achieved more accurately at high “aspect ratio”



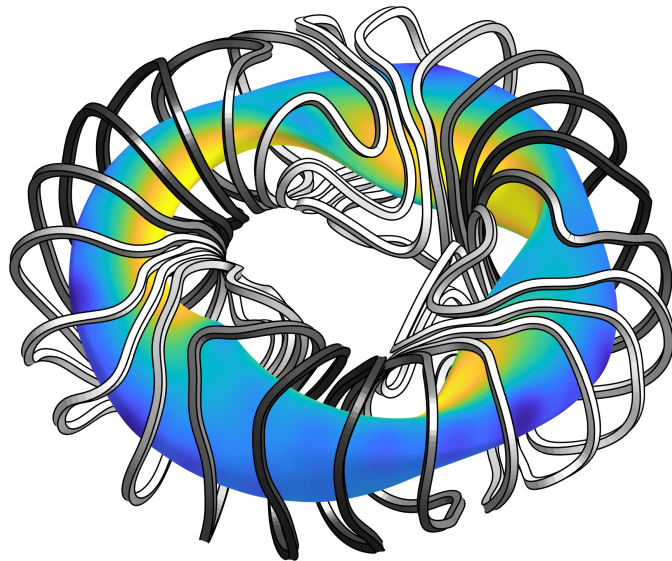
Conclusions:

Symmetry is important in magnetic confinement

- Magnetic confinement and symmetry are connected via canonical momentum conservation.
- Axisymmetry can yield robust confinement but requires an internal current \mathbf{J} . With nonaxisymmetric shaping you don't need \mathbf{J} but confinement is not automatic.
- Without axisymmetry, integrability of \mathbf{B} is not automatic.
- For large $|\mathbf{B}|$, a “hidden” symmetry can yield an approximate conserved quantity that implies particle confinement.

There are important outstanding questions about quasi-symmetry.

- Is there a coordinate-free way to see quasi-symmetry in the Lagrangian?
- Are there phenomena like quasi-symmetry in other physical systems?
- Can quasi-symmetric fields be produced with simply shaped coils far from the plasma?

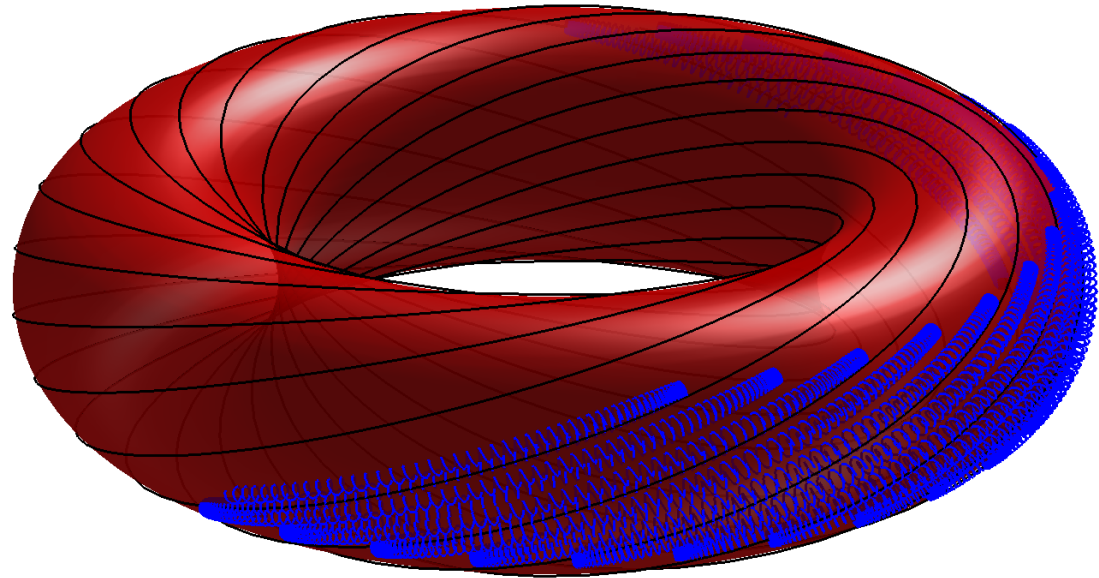


Extra slides

Tokamak vs stellarator

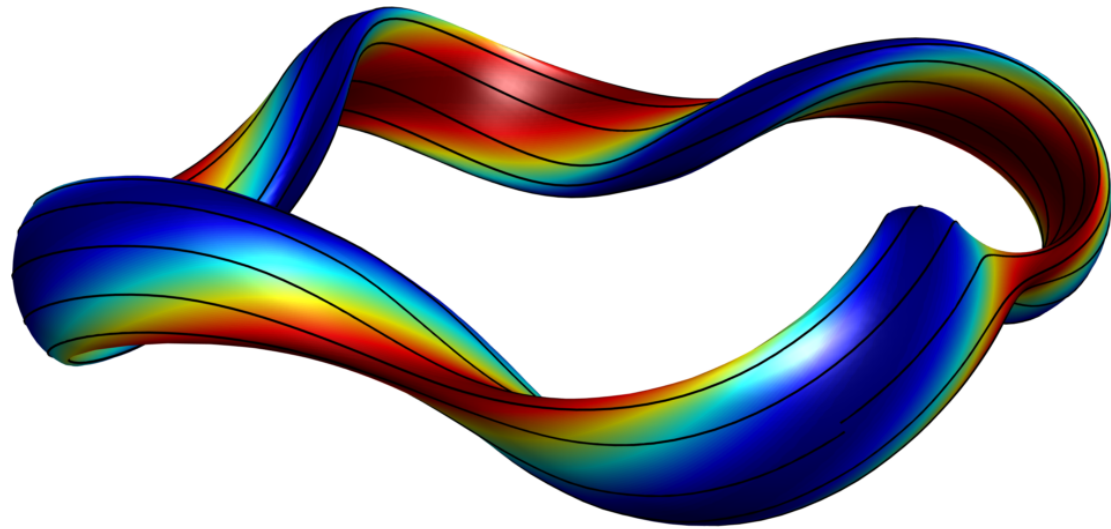
Tokamak:

- Axisymmetric
- Robust confinement
- Requires J_ϕ in plasma:
HUGE problem!



Stellarator:

- Nonaxisymmetric
- Requires careful shaping
to get confinement
- No \mathbf{J} required in plasma

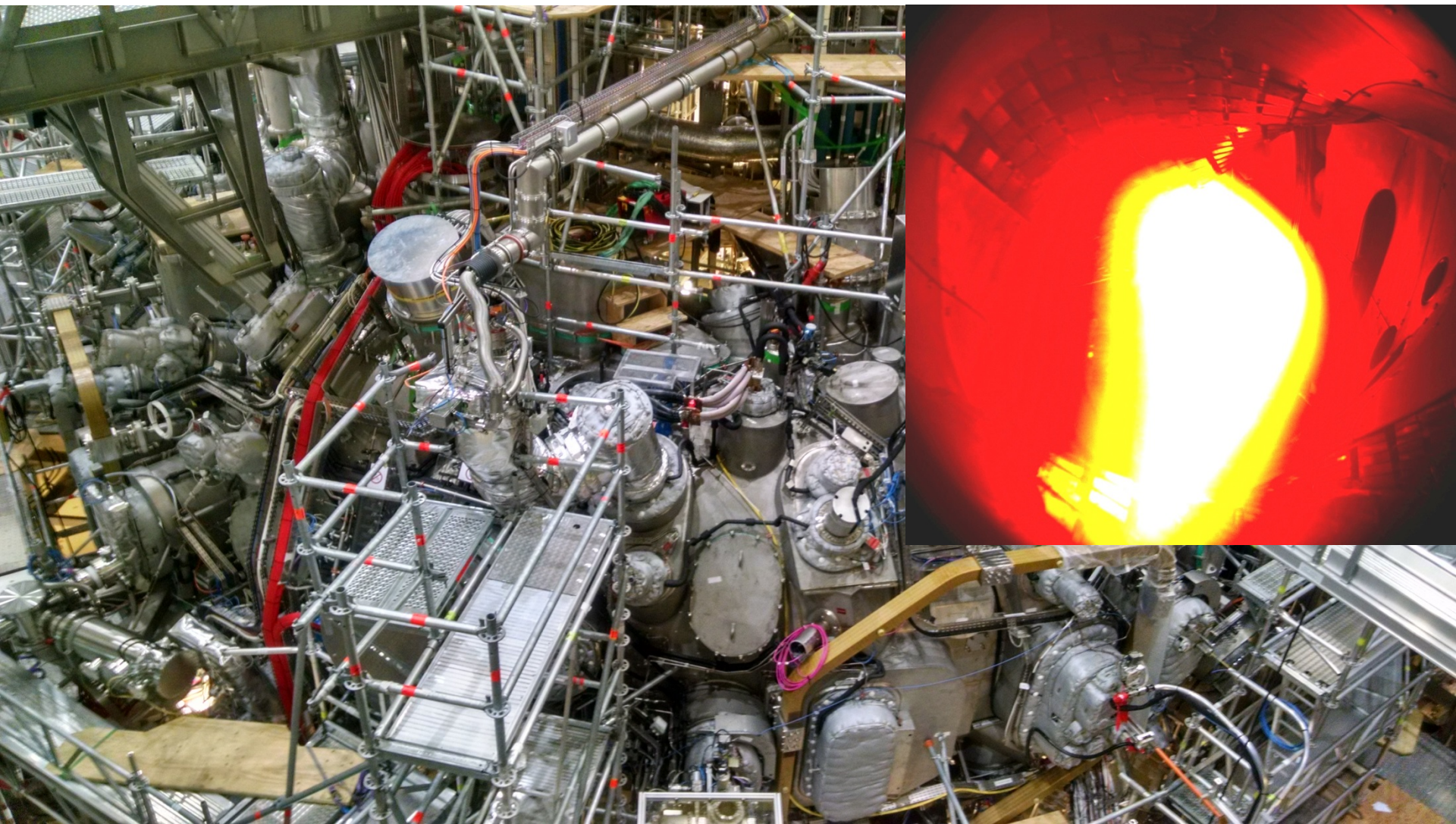


Example of very nonaxisymmetric magnetic confinement: W7-X (Germany)

Science

Oct 21, 2015

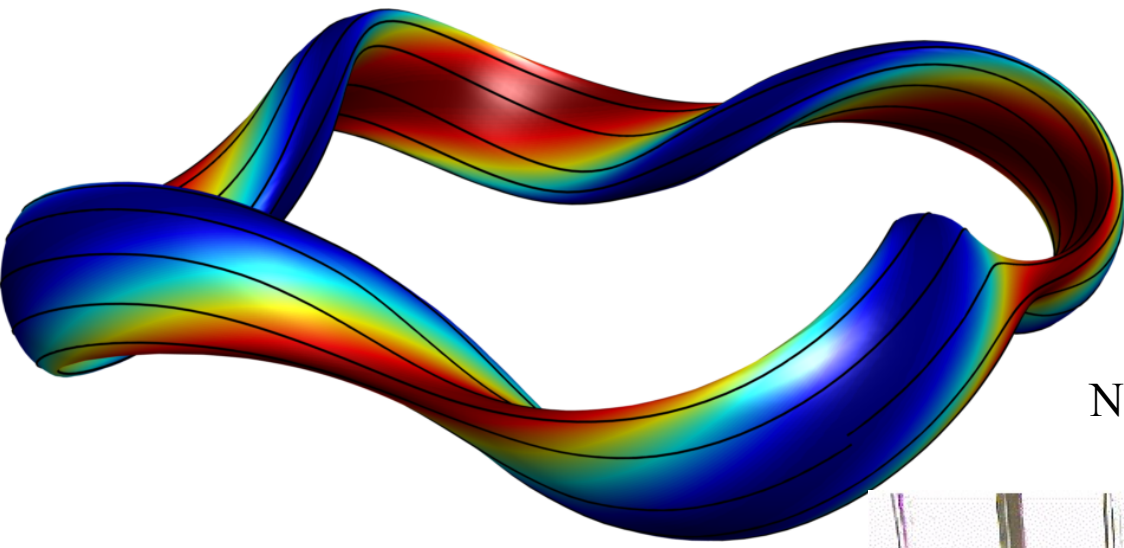
Good-enough particle confinement, but not perfect -
Not quasisymmetric.



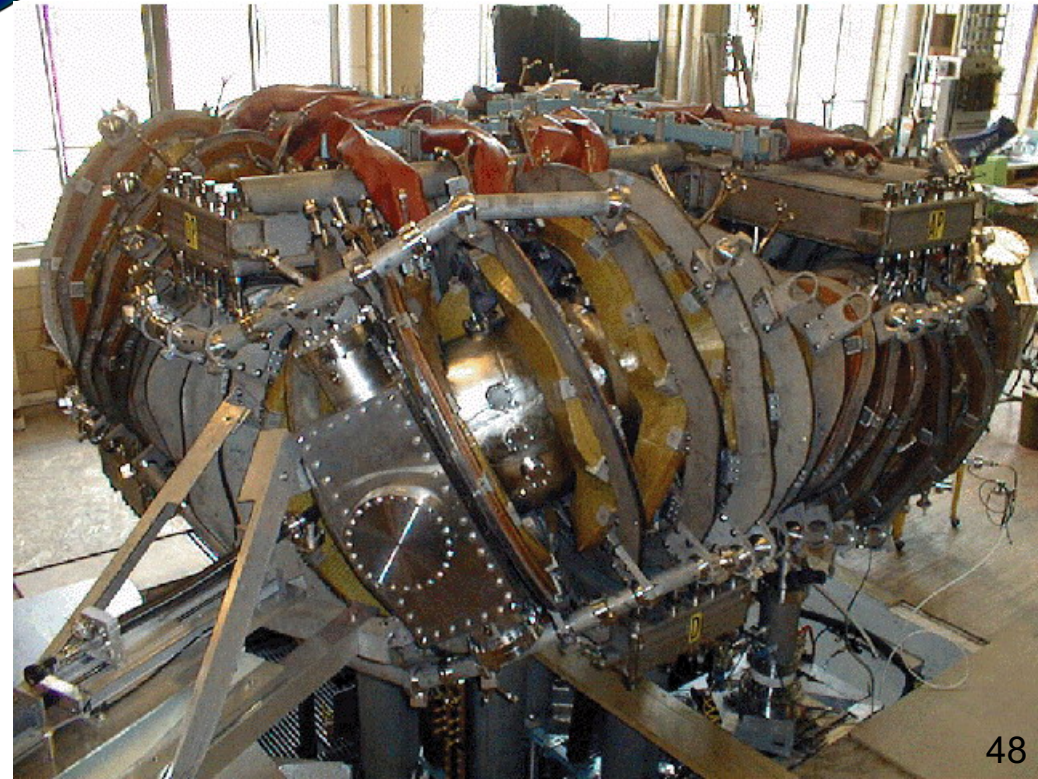
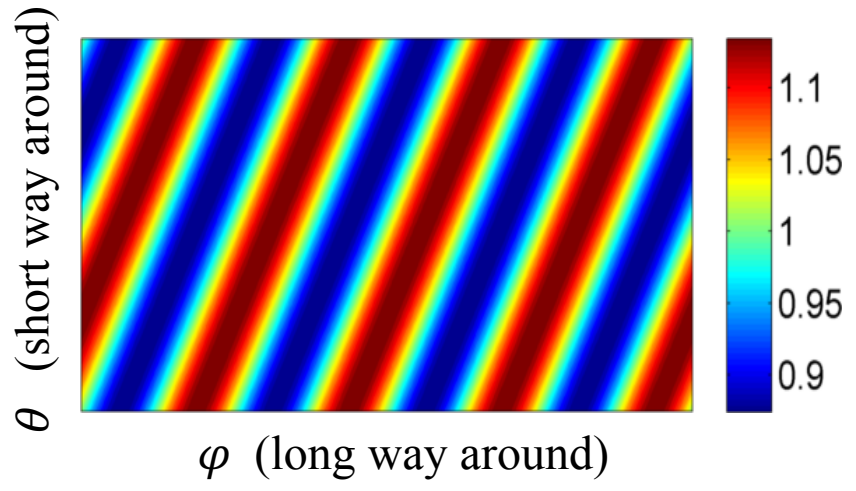
Quasi-symmetry can also be helical.

HSX:
Helically Symmetric eXperiment
(Univ. Wisconsin)

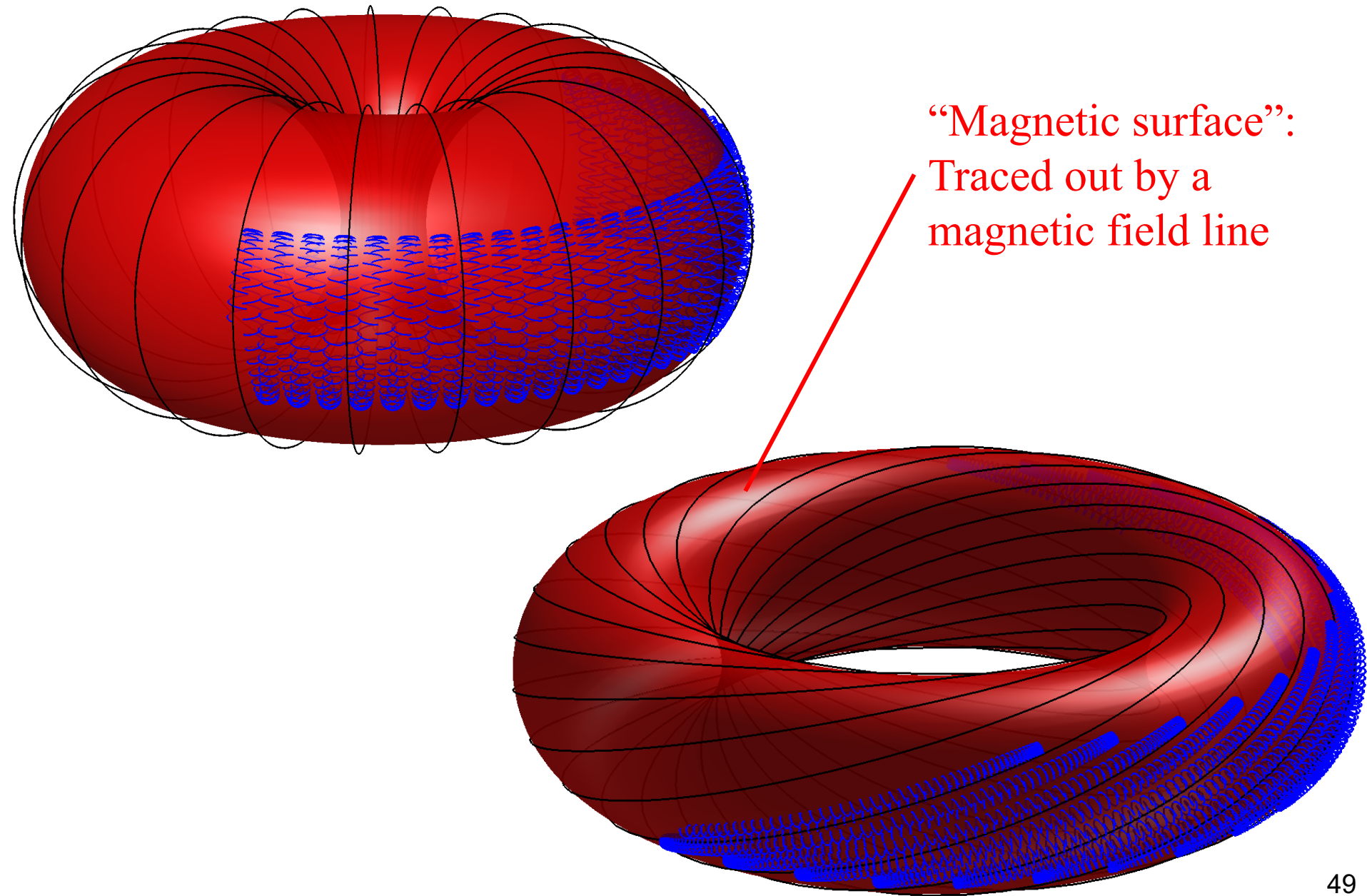
No \mathbf{J} required in plasma \Rightarrow Very stable.



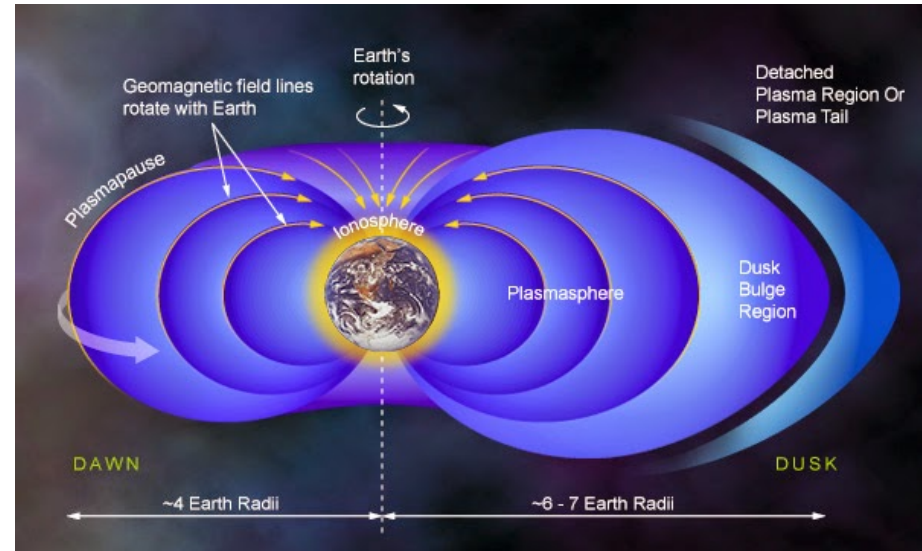
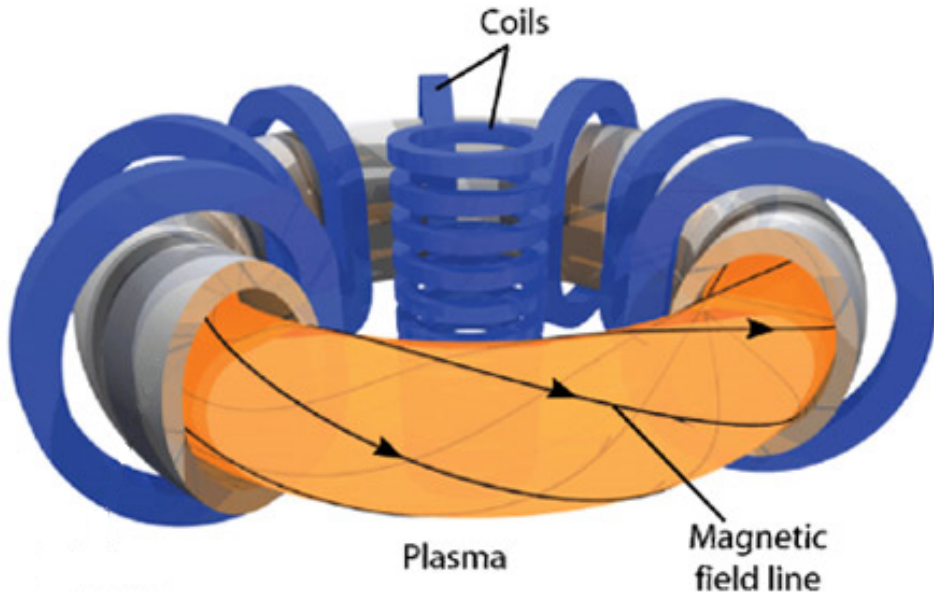
Magnetic field magnitude $|B|$ in Tesla



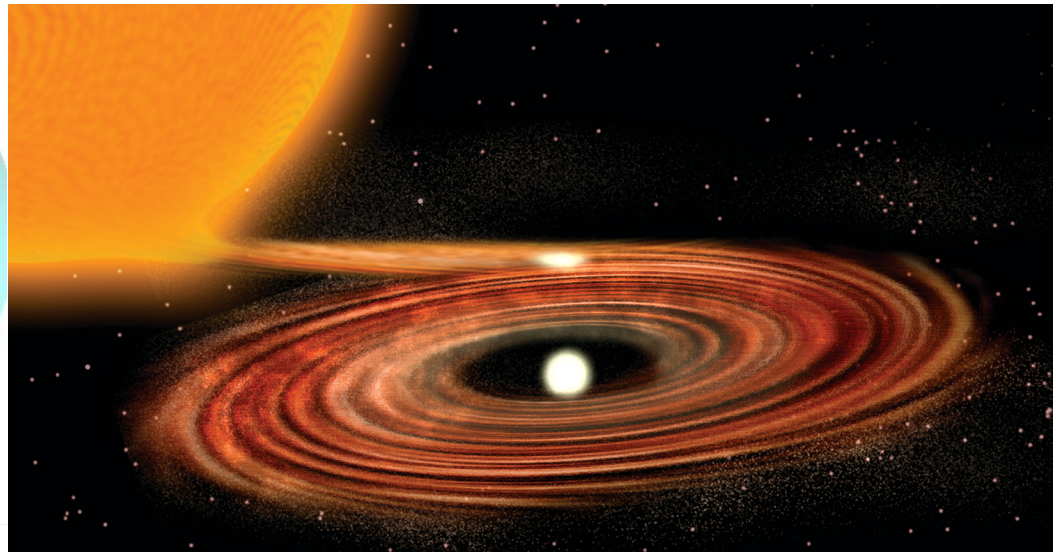
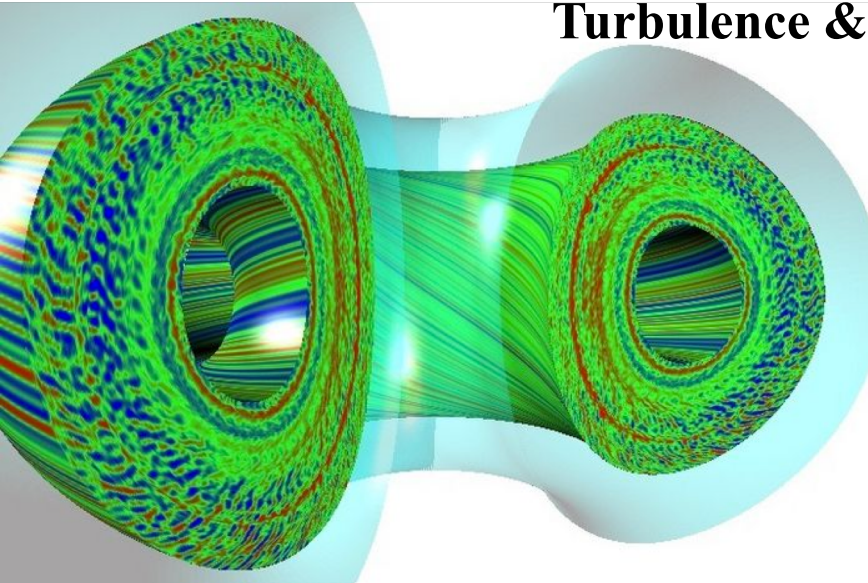
In axisymmetry, particles are confined (close) to $A_\phi R$ surfaces, despite complicated orbits.



No plasma is perfectly axisymmetric.



Turbulence & waves break symmetry:



Why not make the nested $A_\phi R$ surfaces spherical instead of toroidal?

$A_\phi R$ would need to depend on Z along the symmetry axis.

$$B_R = -\frac{\partial A_\phi}{\partial Z} = -\frac{1}{R} \frac{\partial (A_\phi R)}{\partial Z}$$

So B_R would diverge ($\propto 1/R$) along the symmetry axis.

Curl in cylindrical coordinates, assuming axisymmetry

$$B_R = -\frac{\partial A_\phi}{\partial Z} = -\frac{1}{R} \frac{\partial(A_\phi R)}{\partial Z}$$

$$B_\phi = \frac{\partial A_R}{\partial Z} - \frac{\partial A_Z}{\partial R}$$

$$B_Z = \frac{1}{R} \frac{\partial(A_\phi R)}{\partial R}$$