

Lecture 4 - Relativistic Correction of Schrödinger's Equation

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For a non-degenerate $|E_0\rangle$,

$$E^{(1)} = \langle E_0 | V | E_0 \rangle$$

↳ first-order perturbation just the expectation value

$$E^{(2)} = \sum_i \frac{|\langle i | V | E_0 \rangle|^2}{E_0 - E_i}$$

— where the second-order energy correction is always negative (repels the levels)

For a degenerate energy E_0 , you must diagonalize V (find the eigenvalues and eigenvectors) truncated to the degenerate subspace.

degenerate subspace: $\{|m\rangle\}$

with orthogonal complement $\{|i\rangle\}$

first-order perturbation: The secular equation

$$\sum_m \langle m | V | m' \rangle \langle m' | \psi \rangle = E^{(1)} \langle m | \psi \rangle$$

V truncated to the m -subspace

ex: Homework Hamiltonian

$$H_0 = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

degenerate subspace $|m\rangle$:

$$\begin{pmatrix} 1 \\ 0 \\ 0 \end{pmatrix}, \begin{pmatrix} 0 \\ 1 \\ 0 \end{pmatrix} \leftrightarrow E_0 = 0 \text{ — both vectors have the same energy (span the degenerate subspace)}$$

orthogonal complement $|i\rangle$:

$$\begin{pmatrix} 0 \\ 0 \\ 1 \end{pmatrix} \leftrightarrow E_0 = 1$$

Second Order Secular Equation

If $E^{(1)}$ is degenerate, then to find $E^{(2)}$ we must diagonalize the second-order secular equation

find the eigenstates of:

$$\sum_i \frac{\langle m | V | i \rangle \langle i | V | m \rangle}{E_0 - E_i}$$

where the eigenvalues of this operator are the potential perturbation values for $E^{(2)}$

mathematical note:

for σ_z , eigenvalues $= \pm 1$

↳ Pauli matrices, σ_x, σ_y as well

$$A = a\mathbb{1} + \vec{b} \cdot \vec{\sigma}$$

most general Hermitian

2×2 matrix

choose $\vec{b} = |\vec{b}|\hat{z}$ such that

$$A = a\mathbb{1} + |\vec{b}|\sigma_z$$

∴ Eigenvalues are $a \pm |\vec{b}|$

* for a 2×2 Hermitian matrix, you can find the eigenvalues by inspection by writing the operator as a scalar plus a vector dotted into a Pauli matrix

Relativistic Corrections

Let's talk about relativity for a minute...

The Schrödinger equation is NOT Lorentz invariant.

• in relativistic conditions, you're determining the potential for an infinitely large momentum

→ consequently, there's a chance the particle could be anywhere in the Universe

ex: The Free Particle ($\hbar=1$)

one time derivative

$$i\partial_t \psi = \underbrace{\frac{-1}{2m} \nabla^2 \psi}_{\text{two space derivatives}}$$

two space derivatives

→ add a time derivative or remove a space derivative to make invariant

Dirac proposed that we may remove a space derivative so that this becomes Lorentz invariant and yields the relativistic form of $E = P^2/2m$ in Special Relativity, we have instead $E^2 = p^2 c^2 + m^2 c^4$

$$i\partial_t \psi = \mathcal{H} \psi$$

$$= (\alpha^i p_i c + \beta m c^2) \psi$$

↳ just some unknown operator that is linear in momentum

For non-zero p , how do we obtain $E^2 = p^2 c^2 + m^2 c^4$?

→ What does $E^2 = p^2 c^2 + m^2 c^4$ imply about the operators α^i and β ?

($c=1$ now)

Assume: $\Psi = e^{-iEt} e^{i\vec{p} \cdot \vec{x}} \Psi_0$

$\uparrow \Psi_0 = \Psi(x=0, t=0)$ a constant vector in some as yet undetermined Hilbert space $\mathcal{H}_{\text{other}}$ $\rightarrow$ such that the full physical Hilbert space may be expressed as

$\mathcal{H}_{\text{phys}} = \mathcal{H}_{\text{space}} \otimes \mathcal{H}_{\text{other}}$
 $\swarrow$ acts non-trivially here & they commute

$\uparrow \alpha, \beta$ act non-trivially here

$$i\partial_t \Psi = E\Psi$$

$$(i\partial_t)^2 \Psi = E^2 \Psi = \mathcal{H}^2 \Psi$$

$$\begin{aligned} &= (\alpha^i p_i + \beta m)(\alpha^j p_j + \beta m) \Psi \quad / p_i, p_j \text{ interchangeable here} \\ &= [\alpha^i \alpha^j p_i p_j + (\alpha^i \beta + \beta \alpha^i) m p_i + \beta^2 m^2] \Psi \\ &\quad \text{this term must} = 0 \end{aligned}$$

to match the E^2 equation:

$$-\beta^2 = 1$$

$$-\alpha^i \beta + \beta \alpha^i = 0 \quad (\alpha^i \text{ and } \beta \text{ anti-commute})$$

$$-\alpha^i \alpha^j + \alpha^j \alpha^i = 2\delta^{ij} \quad \text{Kronecker delta}$$

*Note: this relation also holds for Pauli matrices

$$\sigma^i \sigma^j + \sigma^j \sigma^i = 2\delta^{ij}$$

$$\sigma^x \sigma^y + \sigma^y \sigma^x = 0; \quad \sigma^x \sigma^x + \sigma^y \sigma^y = 2\mathbf{I}_2$$

$$\rightarrow \alpha^i = \sigma^x \otimes \sigma^i$$

$\downarrow$ tensor product space

$$\beta = \sigma^z \otimes \mathbf{I}_2$$

$$\downarrow \beta^2 = \mathbf{I}_2 \otimes \mathbf{I}_2 = \mathbf{I}_4$$

$$\sigma^i \sigma^j = (\sigma^x)^2 \otimes \sigma^i \sigma^j$$

$$= \mathbf{I}_2 \otimes \sigma^i \sigma^j$$

a four-dimensional vector space

Check:

$$\begin{aligned} \alpha^i \beta + \beta \alpha^i &= (\sigma^x \otimes \sigma^i)(\sigma^z \otimes \mathbf{I}_2) + (\sigma^z \otimes \mathbf{I}_2)(\sigma^x \otimes \sigma^i) \\ &= \sigma^x \sigma^z \otimes \sigma^i + \sigma^z \sigma^x \otimes \sigma^i \\ &= 0 \quad \checkmark \end{aligned}$$

$\rightarrow$

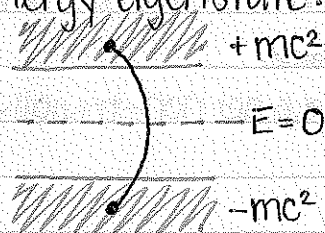
$$\beta = \sigma^z \otimes I_2 \longleftrightarrow \begin{pmatrix} I_2 & 0 \\ 0 & -I_2 \end{pmatrix}$$

$$\alpha^i = \sigma^x \otimes \sigma^i \longleftrightarrow \begin{pmatrix} 0 & \sigma^i \\ \sigma^i & 0 \end{pmatrix}$$

for the lowest energy eigenstate: $\vec{p} = 0$

$$\beta m \psi = E \psi$$

$$E = \pm mc^2$$



A wavefunction can't just predict matter, it must also predict antimatter.