#### Particle-in-Cell (PIC) Simulations of Plasmas

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# Cartoon PIC

Adapted from

[https://www.lanl.gov/science/NSS/issue2\\_2010/story4.shtml](https://www.lanl.gov/science/NSS/issue2_2010/story4.shtml)



**E** and **B** are known on the grid. Particles move freely.

Why Doing Plasma Physics via Computer Simulations Using Particles Makes Good Physical Sense Inspired by Birdsall & Langdon, *Plasma Physics via Computer Simulation*

- **Debye length**  $\lambda_D = v_{th}/\omega_{pe} \ll L$ ; we care about  $\lambda \gtrsim \lambda_D$ .
- ► For a meaningful plasma  $N_D = n \lambda_D^3$  ≫ 1
- But that means

 $\frac{\mathsf{KE}}{\mathsf{PE}}$  (thermal kinetic energy)  $N_D^{2/3} \gg 1$ <br>PE (electrostatic potential energy)

- <sup>I</sup> ∴ *Particles interact collectively, not discretely*.
- $\triangleright$  Grids with  $\Delta x \leq \lambda_D$  capture the important physics without the unimportant inter-particle effects.

# Cartoon Timestep



### Updating **x**, **v**, **J**, **B**, and **E**

 $\blacktriangleright$  Field advancement:

$$
\frac{\partial \mathbf{B}}{\partial t} = -c \nabla \times \mathbf{E} \qquad \frac{\partial \mathbf{E}}{\partial t} = c \nabla \times \mathbf{B} - 4\pi \mathbf{J}
$$

 $\blacktriangleright$  Particle advancement:

$$
\frac{d\mathbf{x}}{dt} = \mathbf{v} \qquad \frac{d(\gamma \mathbf{v})}{dt} = \frac{q}{m} \left( \mathbf{E} + \frac{\mathbf{v}}{c} \times \mathbf{B} \right)
$$

 $\blacktriangleright$  Current density update:

$$
\mathbf{J} = \sum_i q_i \mathbf{v}_i S(\mathbf{X} - \mathbf{x}_i)
$$

where  $S(X - x)$  is a shape function.

#### Translating Between Particles and the Grid

Adapted from <https://www.particleincell.com/2010/es-pic-method/>



# Effective Particle Shapes (1D)

Adapted from <https://perswww.kuleuven.be/~u0052182/weather/pic.pdf>



 $\blacktriangleright$  Nearest gridpoint

 $\blacktriangleright$  First-order (cloud-in-cell)

 $\blacktriangleright$  Quadratic spline

#### Does PIC Satisfy  $\nabla \cdot \mathbf{B} = 0$  and  $\nabla \cdot \mathbf{E} = 4\pi \rho$ ?

$$
\text{Numerically, } \boldsymbol{\nabla} \cdot (\boldsymbol{\nabla} \times) = 0
$$

$$
\frac{\partial}{\partial t}(\nabla \cdot \mathbf{B}) = -c \, \nabla \cdot \nabla \times \mathbf{E} = 0
$$

If  $\nabla \cdot \mathbf{B} = 0$  at  $t = 0$ , it remains so (ignoring round-off)

In contrast,

$$
\frac{\partial}{\partial t}(\nabla \cdot \mathbf{E}) = c \, \nabla \cdot \nabla \times \mathbf{B} - 4\pi \nabla \cdot \mathbf{J} = -4\pi \nabla \cdot \mathbf{J}
$$

To satisfy Gauss's Law requires

$$
\frac{\partial \rho}{\partial t} + \mathbf{\nabla} \cdot \mathbf{J} = 0
$$

## Unfortunately  $\cdots$

Continuity is not, in general, satisfied

Corrections fall into two broad categories

- ► "Fix" **E**
- <sup>I</sup> "Fix" **J**

An approach of the first type: Suppose a  $\Phi$  exists such that

$$
\mathbf{E}' = \mathbf{E} - \nabla \Phi \qquad \text{where} \qquad \nabla \cdot \mathbf{E}' = 4\pi \rho
$$

Find Φ by solving

$$
\nabla^2 \Phi = \boldsymbol{\nabla} \boldsymbol{\cdot} \boldsymbol{E} - 4 \pi \rho \equiv b
$$

This ( $\nabla^2 \Phi = b$ ) is Poisson's equation and can be solved many different ways: FFTs, matrix methods, multigrid methods, · · ·

# An Alternative: Fluid vs. PIC Simulations

#### Fluid (MHD) Advantages:

- $\triangleright$  Correct on large scales
- $\triangleright$  Computationally fast

Disadvantages:

 $\triangleright$  Wrong at small scales

Kinetic (PIC) Advantages:

 $\blacktriangleright \approx$  All of the physics

Disadvantages:

- $\blacktriangleright$  Must resolve important scales
- $\triangleright$  Computationally painful



### Resolution for Explicit PIC

For timestep ∆*t*, grid spacing ∆*x*, and velocity *u* a general constraint is

▶ CFL (Courant-Friedrichs-Lewy):

$$
\frac{u\Delta t}{\Delta x}\leq 1
$$

For plasmas also need to resolve important physical scales

$$
\blacktriangleright \Delta x < (\lambda_D, \, \omega_{\text{pe}}, \, \rho_{\text{Le}})
$$

<sup>I</sup> ∆*t* < (ω*pe*, ω*ce*)

**Not resolving generally leads to numerical instability.**

#### Kinetic Scales

How painful?

- **►** Solar corona: *B* = 50 G, *n* = 10<sup>9</sup> cm<sup>-3</sup>, *L* ≈ 10<sup>9</sup> m,  $\tau \approx 10^3$  s
	- $\blacktriangleright$  *d<sub>p</sub>* ≈ 10 m
	- $\blacktriangleright$  Ωpc  $^{-1}$   $\approx$  2  $\times$  10<sup>-6</sup> s
	- **►**  $\omega_{pi}^{-1} \approx 2 \times 10^{-8}$  s
- ► Magnetosphere: *B* = 2 × 10<sup>-4</sup> G, *n* = 20 cm<sup>-3</sup>,  $L \approx 10^4$  km,  $\tau \approx 10^3$  s
	- $\blacktriangleright$  *d<sub>p</sub>* ≈ 50 km
	- $▶ \Omega_{\rho c}^{-1} \approx 0.5~\text{s}$
	- **►**  $\omega_{pi}^{-1} \approx 2 \times 10^{-4}$  s
- **F** Tokamak:  $B = 3 \times 10^4$  G,  $n = 2 \times 10^{13}$  cm<sup>-3</sup>,  $L \approx 10^2$  cm,  $\tau \approx 10^{-2}$  s
	- $\blacktriangleright$  *d<sub>p</sub>* ≈ 5 cm
	- <sup>I</sup> Ω −1 *pc* ≈ 3 × 10<sup>−</sup><sup>9</sup> s
	- **►**  $\omega_{pi}^{-1} \approx 2 \times 10^{-10}$  s

# The Annoyances of Reality

And How to Get Around Them

Besides real systems being much larger than kinetic scales, nature insists on making the situation worse.

- $\blacktriangleright$  *m<sub>p</sub>*/*m*<sub>e</sub> ≈ 1836
- $\blacktriangleright$  *c*/*v*<sub>A</sub>  $\gg$  1

The resulting separation of scales is computationally challenging. To combat it, artificial values are often used

$$
m_p/m_e = 400, 100, 25
$$

$$
\blacktriangleright \hspace{.1cm} c/v_A = 20 - 50
$$

Potential unwanted side-effects (e.g.,  $v_{th,e} \rightarrow c$ ) must be kept in mind.

# PIC on Supercomputers

Domain Decomposition



A useful simulation ( $\geq 10^{10}$  particles) needs many cores working in parallel. Communication should be minimized.

## Supercomputer Performance



# Brief Notes on PIC-Related Topics

#### Accurate Numerical Differentiation

Not PIC-Specific From the Taylor series

$$
f(x_0+\Delta x)=f(x)+\Delta x\left.\frac{df}{dx}\right|_{x_0}+\frac{(\Delta x)^2}{2}\left.\frac{d^2f}{dx^2}\right|_{x_0}+\mathcal{O}(\Delta x^3)
$$

comes the approximation

$$
\frac{df}{dx} = \frac{f(x + \Delta x) - f(x)}{\Delta x} + \mathcal{O}(\Delta x)
$$

Incorporating a variation

$$
f(x_0 - \Delta x) = f(x) - \Delta x \left. \frac{df}{dx} \right|_{x_0} + \frac{(\Delta x)^2}{2} \left. \frac{d^2 f}{dx^2} \right|_{x_0} + \mathcal{O}(\Delta x^3)
$$

gives something more accurate

$$
\frac{df}{dx} = \frac{f(x + \Delta x) - f(x - \Delta x)}{2\Delta x} + \mathcal{O}(\Delta x^2)
$$

#### Symmetry Reduces Errors and Helps Stability

From <https://www.particleincell.com/2011/velocity-integration/>

Basic leapfrog algorithm



# Gridding Systems

Adapted from <https://commons.wikimedia.org/wiki/File:Yee-cube.svg>

The Yee lattice is a popular  $-$  but not the only  $-$  choice. **E** is known on edges, **B**/**H** on faces.



The finite-difference versions of Maxwell's equations are nice, but bookkeeping is an annoyance.

### Explicit Versus Implicit Algorithms

Consider

$$
\frac{\partial u}{\partial t} = D \frac{\partial^2 u}{\partial x^2}
$$

Explicit discretization:

$$
\frac{u_j^{n+1}-u_j^n}{\Delta t}=D\left[\frac{u_{j+1}^n-2u_j^n+u_{j-1}^n}{(\Delta x)^2}\right]
$$

Implicit discretization:

$$
\frac{u_j^{n+1} - u_j^n}{\Delta t} = D \left[ \frac{u_{j+1}^{n+1} - 2u_j^{n+1} + u_{j-1}^{n+1}}{(\Delta x)^2} \right]
$$

Implicit is typically much more stable but requires much more work to solve.