

## The MHD (magnetohydrodynamic) equations

Want to develop a set of fluid equations that can be used to model plasma dynamics in a magnetic field. Will use the

Boltzmann equation with collision operator and Maxwell's Eqs.

To justify the fluid model, need to assume that the collision rate is high

⇒ high enough so that both species are drifting Maxwellians ⇒ same T  
 ⇒ don't require equal temperatures.

~~don't require equal temperatures~~

⇒ consider for simplicity have the masses to be of the same order so don't distinguish use same ordering for each species.

Assume the following

$$u \sim v_t \sim c_A \sim \text{Alfven speed} \sim \frac{B}{\sqrt{4\pi n}} \\ \omega \sim v_t / L, \quad \gamma \sim \Omega$$

Define a small parameter

$$\epsilon \equiv \frac{\omega}{\gamma} \sim \left( \frac{v_t}{c} \right)^2 \sim \frac{v_L}{L}$$

Boltzmann Equation for species  $\alpha, \beta$

$$\frac{\partial f_\alpha}{\partial t} + \mathbf{v} \cdot \nabla f_\alpha + \left(\frac{q}{m}\right)_\alpha \left(\mathbf{E} + \frac{1}{c} \mathbf{v} \times \mathbf{B}\right) \cdot \nabla_{\mathbf{v}} f_\alpha = C_{\alpha\alpha} + C_{\alpha\beta}$$

$\omega$        $\omega$        $\frac{q}{m} \frac{E}{v t}$        $\frac{q}{m} \frac{L}{c} B$        $\nu$        $\nu$

~~$\omega$~~        ~~$\omega$~~        $\frac{qB}{mc}$        $\left(\frac{cE}{Bvt}\right)$        $\Omega$        $\nu$        $\nu$

$\epsilon$        $\epsilon$        $1$        $1$        $1$        $1$        $1$

Gauss' Law

$$\nabla \cdot \mathbf{E} = 4\pi e (S_{\alpha} f_i - S_{\alpha} f_e)$$

$\frac{B}{c} \frac{E}{BL}$        ~~$\frac{B}{c} \frac{E}{BL}$~~        $4\pi e n_i$        $4\pi e n_e$

get rid of  $n$  through  $CA = B^2/4\pi m n$

$\frac{v t}{m 4\pi L} \frac{B^2 m c}{n_i q c^2 B}$        $1$        $1$

$\frac{v t^3}{c^2 L \Omega}$        $1$        $1$

$\left(\frac{v t}{c}\right)^2 \frac{r_L}{L}$        $1$        $1$

$\epsilon^2$        $1$        $1$        $\Rightarrow$  quasi-neutrality

Faradays Law

$$\frac{1}{c} \frac{\partial \mathbf{B}}{\partial t} + \nabla \times \frac{\mathbf{E}}{c} = 0$$

$$\frac{\omega}{c} \mathbf{B} \quad \mathbf{E} \frac{c}{L} \mathbf{B}$$

$$\omega, \frac{v_t}{L}$$

$$1, 1$$

Ampere's Law

$$\nabla \times \mathbf{B} = \frac{1}{c} \frac{\partial \mathbf{E}}{\partial t} + \frac{4\pi e}{c} \int dV \mathbf{v} (f_i - f_e)$$

$$\frac{B}{L} \quad \frac{\omega}{c} \frac{E}{L} \quad \frac{4\pi e}{c} n_i v_t \quad \frac{4\pi e}{c} \frac{v_t m}{B^2 c A^2} \quad \frac{B^2}{m}$$

$$\frac{B}{L} \frac{m}{B} \frac{c}{e} \quad \frac{\omega}{c^2} \frac{v_t}{L} \frac{m}{B} \frac{c}{e} \quad \frac{1}{v_t} \quad \frac{1}{v_t}$$

$$\frac{v_t}{L} \quad \frac{v_t^2}{c^2} \frac{\omega}{L} \quad 1, 1$$

$$e \quad e^2 \quad 1, 1$$

Lowest order

$$\textcircled{1} \quad \frac{q}{m} \left( \mathbf{E}_0 + \frac{1}{c} \mathbf{v} \times \mathbf{B}_0 \right) \cdot \frac{\partial}{\partial v} \epsilon_{\alpha 0} = C_{\alpha \alpha 0} + C_{\alpha \beta 0}$$

$$\textcircled{2} \quad 0 = n_i^0 - n_e^0 \quad \Leftarrow \text{quasi-neutrality}$$

$$\textcircled{3} \quad \frac{1}{c} \frac{\partial \mathbf{B}_0}{\partial t} + \nabla \times \mathbf{E}_0 = 0$$

$$\textcircled{4} \quad 0 = (n_i \mathbf{u}_i)_0 - (n_e \mathbf{u}_e)_0 \Rightarrow \mathbf{J}_0 = 0$$

From ①  $\Rightarrow$  Maxwellian with  $n_i = n_e$ ,  $\mathbf{u}_i^0 = \mathbf{u}_e^0 = \mathbf{u}_0$   
and  $T_e = T_i$

$\Rightarrow$  eliminates  $C_{\alpha \alpha 0}$ ,  $C_{\alpha \beta 0}$

$$\left( \mathbf{E}_0 + \frac{1}{c} \mathbf{v} \times \mathbf{B}_0 \right) \cdot \frac{\partial}{\partial v} (\mathbf{v} - \mathbf{u}_0) = 0$$

$$\mathbf{E}_0 \cdot (\mathbf{v} - \mathbf{u}_0) - \frac{1}{c} \mathbf{u}_0 \cdot \mathbf{v} \times \mathbf{B}_0 = 0$$

$$\left( \mathbf{E}_0 + \frac{1}{c} \mathbf{u}_0 \times \mathbf{B}_0 \right) \cdot \mathbf{v} = \mathbf{E}_0 \cdot \mathbf{u}_0$$

Valid for any  $\mathbf{v}$  so  $\mathbf{E}_0 + \frac{1}{c} \mathbf{u}_0 \times \mathbf{B}_0 = 0$

$\Rightarrow \mathbf{u}_0$  is the  $\mathbf{E}_0 \times \mathbf{B}_0$  drift.

$$\Rightarrow \mathbf{E}_{0 \parallel} \mathbf{u}_{0 \parallel} = 0$$

$$\Rightarrow \mathbf{E}_{0 \parallel} = 0$$

$$\boxed{\mathbf{E}_0 = -\frac{1}{c} \mathbf{u}_0 \times \mathbf{B}_0}$$



$\mathbf{E}_0 = 0$  in  
plasma frame

$\Rightarrow$  good  
conductor

First order in  $\epsilon$

$$\frac{df_{\alpha 0}}{dt} + v_m \cdot \nabla f_{\alpha 0} + \left(\frac{q}{m}\right)_\alpha \left[ \left( \mathbf{E} + \frac{1}{2} \mathbf{v} \times \mathbf{B} \right) \cdot \frac{d}{dt} f_\alpha \right]_1$$

$$= C_{\alpha 1} + C_{\alpha B 1}$$

$$0 = n_{i1} - n_{e1}$$

$$\frac{1}{c} \frac{dB_1}{dt} + \nabla \times \mathbf{E}_1 = 0$$

$$\nabla \times \mathbf{B}_0 = \frac{4\pi}{c} e \left[ (n_i v_i)_1 - (n_e v_e)_1 \right]$$

Recall:

$$\int dv C_{\alpha\alpha} \begin{pmatrix} 1 \\ m v_x \\ \frac{1}{2} m v^2 \end{pmatrix} = 0$$

$\Rightarrow$  self-collisions don't change #,  
momentum or energy

$$\int dv C_{\alpha\beta} \begin{pmatrix} 1 \\ m v_x \\ \frac{1}{2} m v^2 \end{pmatrix} = \begin{pmatrix} 0 \\ -R_{\alpha\beta} \\ -\left(\frac{\partial W}{\partial t}\right)_{\alpha\beta} \end{pmatrix}$$

$$R_{\alpha\beta} + R_{\beta\alpha} = 0$$

$$\left(\frac{\partial W}{\partial t}\right)_{\alpha\beta} + \left(\frac{\partial W}{\partial t}\right)_{\beta\alpha} = 0$$

~~First order~~

Can eliminate collision operators by integrating over  $u$

$$\frac{d}{dt} n_{\alpha} + \nabla \cdot n_{\alpha} u_{\alpha} = 0$$

suppress subscript  $\alpha$ .

~~Integrate over  $\int d^3u$~~

Operate with  ~~$\int d^3u$~~   $\int d^3u$

$$n_{\alpha} m_{\alpha} \frac{d}{dt} u_{\alpha} = -\nabla P_{\alpha} + n_{\alpha} e_{\alpha} \left( E_{\alpha} + \frac{1}{c} u_{\alpha} \times B_{\alpha} \right) + e_{\alpha} \left( n_{\alpha} e_{\alpha} E + \frac{1}{c} (n u)_{\alpha} \times B \right) - P_{\alpha} e_{\alpha}$$

$\Rightarrow$  sum over  $\alpha, \beta$  and use  $n_{\alpha} e_{\alpha} = n e$   
 $\Rightarrow$  define  $e = (m_{\alpha} + m_{\beta}) u$ ,  $P_{\alpha} + P_{\beta} = P$

$$e \frac{d}{dt} u = -\nabla P + \frac{1}{c} J_{\parallel} \times B$$

$$J_{\parallel} = e (n u)_{\parallel} - e (n u)_{\perp}$$

~~Note~~

Operate with  $\int d^3u v^2$

$$\frac{d}{dt} \left( \frac{1}{2} n m_{\alpha} u^2 + \frac{3}{2} P_{\alpha} \right) + \nabla \cdot \left( \frac{1}{2} n m_{\alpha} u^2 + \frac{3}{2} P_{\alpha} \right) u + \nabla \cdot (P_{\alpha} u) = e_{\alpha} \left[ n u_{\parallel} \cdot E_{\alpha} + (n u)_{\perp} \cdot E_{\alpha} \right] - \left( \frac{\partial \omega}{\partial t} \right)_{\alpha \beta}$$

Sum over species

$$\frac{\partial}{\partial t} \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) + \nabla \cdot \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) \underline{u} + \nabla \cdot (P \underline{u})$$

$$= \underline{J}_1 \cdot \underline{E} = - \underline{J}_1 \cdot \underline{u} \times \underline{B} \frac{1}{c}$$

$$\nabla \times \underline{B}_0 = \frac{4\pi e}{c} \underline{J}_1 = \underline{u} \cdot \frac{1}{c} \underline{J}_1 \times \underline{B}$$

~~$$\frac{\partial}{\partial t} \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) + \nabla \cdot \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) \underline{u} + \nabla \cdot (P \underline{u}) = \underline{u} \cdot \left( \frac{\partial e \underline{u}}{\partial t} + \nabla \cdot (e \underline{u}) \right)$$~~

$$\frac{\partial}{\partial t} \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) + \nabla \cdot \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) \underline{u} + \nabla \cdot (P \underline{u})$$

$$= \underline{u} \cdot \nabla P + e \underline{u} \cdot \frac{d \underline{u}}{dt}$$

$$e \frac{d}{dt} \frac{u^2}{2}$$

~~$$\frac{1}{2} u^2 \frac{\partial e}{\partial t} + \frac{\partial}{\partial t} \left( \frac{3}{2} P \right) + \underline{u} \cdot \nabla \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right)$$

$$+ \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) \nabla \cdot \underline{u} + \underline{u} \cdot \nabla P + P \nabla \cdot \underline{u}$$

$$= \underline{u} \cdot \nabla P + e \underline{u} \cdot \nabla \frac{u^2}{2}$$~~

~~$$-\frac{1}{2} u^2 \left( + e \nabla \cdot \underline{u} + \underline{u} \cdot \nabla e \right) + \frac{\partial}{\partial t} \left( \frac{3}{2} P \right) + \frac{1}{2} u^2 \underline{u} \cdot \nabla e$$

$$+ \frac{3}{2} \underline{u} \cdot \nabla P + \left( \frac{1}{2} e u^2 + \frac{3}{2} P \right) \nabla \cdot \underline{u} = 0$$~~

$$\cancel{\frac{dP}{dt}} + \frac{dP}{dt} + \frac{5}{3} P \nabla \cdot \underline{u} = 0$$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \rho \underline{u} = 0 \quad \Rightarrow \rho = (\rho_0 + \rho_1) u$$

$$\rho \frac{d}{dt} \underline{u} = -\nabla P + \frac{1}{c} \underline{J} \times \underline{B} = -\nabla \left( P + \frac{B^2}{8\pi} \right) + \frac{1}{4\pi} \underline{B} \cdot \nabla \underline{B}$$

$$\nabla \times \underline{B} = \frac{4\pi}{c} \underline{J}$$

$$\nabla \cdot \underline{B} = 0$$

$$\frac{1}{c} \frac{\partial}{\partial t} \underline{B} + \nabla \times \underline{E} = 0$$

$$\underline{E} = -\frac{1}{c} \underline{u} \times \underline{B}$$



## Ideal MHD Equilibria

In exploring the stability and dynamics of magnetized plasma, it is useful to start with states that are time stationary so that they are in a state of equilibrium. We will discuss several examples of such states. For simplicity we will limit this discussion to situations without flow although in many cases equilibria with flow can also be constructed. The basic equations are

momentum:

$$\textcircled{1} \quad 0 = -\nabla P + \frac{1}{c} \underline{J} \times \underline{B}$$

$$\textcircled{2} \quad \underline{E} = 0$$

$$\textcircled{3} \quad \nabla \cdot \underline{B} = 0$$

$$\textcircled{4} \quad \nabla \times \underline{B} = \frac{4\pi}{c} \underline{J}$$

Alternatively, we can write  $\textcircled{1}$  as

$$0 = -\nabla P + \frac{1}{c} \frac{c}{4\pi} \underbrace{(\nabla \times \underline{B}) \times \underline{B}}_{-\nabla \frac{B^2}{2} + \underline{B} \cdot \nabla \underline{B}}$$

$$\textcircled{1}' \quad 0 = -\nabla \left( P + \frac{B^2}{8\pi} \right) + \underline{B} \cdot \nabla \underline{B}$$

Egn ① requires that

$$\mathbf{B} \cdot \nabla P = 0$$

so that the pressure is constant along  $\mathbf{B}$ .

$\Rightarrow$  any parallel pressure gradients are relaxed by sound waves.

Egn ④ requires

$$\nabla \cdot \mathbf{J} = 0$$

which implies that charge can not build up. This would produce an  $\mathbf{E} \neq 0$  and since

$$\mathbf{E} = -\mathbf{u} \times \mathbf{B} \frac{1}{c}$$

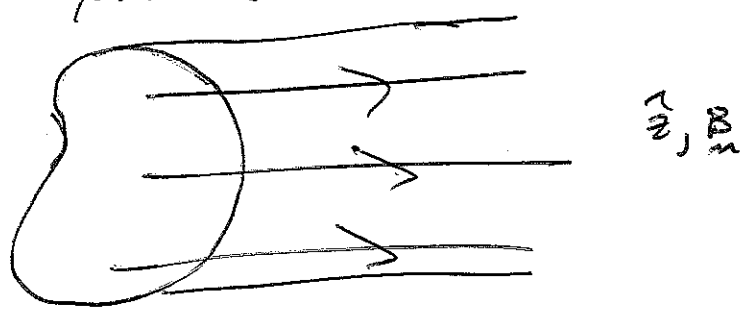
flows would be generated.

Case I  $\mathbf{B} \cdot \nabla \mathbf{B} = 0$

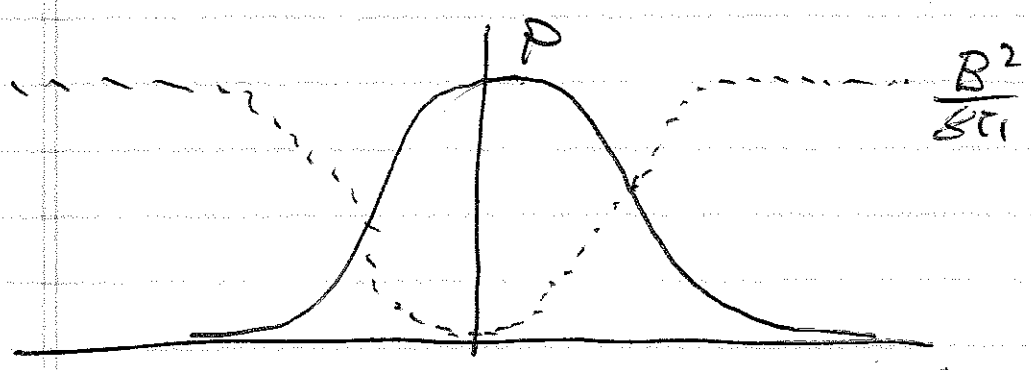
In this case we have straight field lines and

$$\nabla \left( P + \frac{B^2}{8\pi} \right) = 0$$

$\Rightarrow$  the total pressure, magnetic plus plasma must be constant

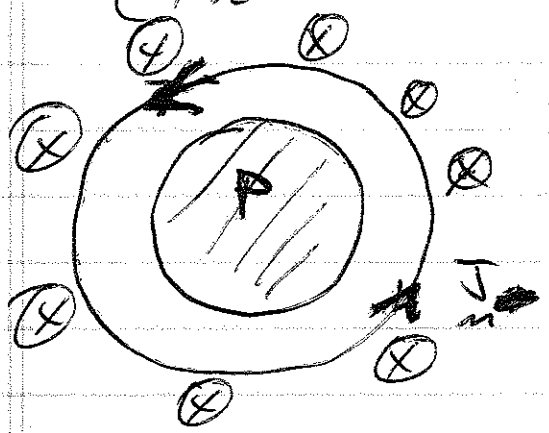


This is not a trivial state, e.g.,



$$\frac{\partial}{\partial r} \left( p + \frac{B^2}{8\pi} \right) = 0$$

Confinement experiments with such equilibria are called  $\theta$ -pinches (the currents are in the  $\theta$  direction)

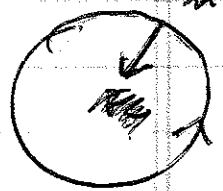


$$-\frac{\partial B_z}{\partial r} = \frac{4\pi}{c} J_\theta$$

Case II  $B_\theta \neq 0, B_z \neq 0$

Consider a pure azimuthal field  $B_\theta \hat{e}_\theta$

$$\begin{aligned} \nabla \cdot \mathbf{B} &= \nabla \cdot (B_\theta \hat{e}_\theta) = \frac{1}{r} \frac{\partial}{\partial r} (r B_\theta) \\ &= -\frac{B_\theta}{r} \hat{e}_r \end{aligned}$$



$$\Rightarrow R_0 = r$$

This ~~to occur~~ results from "magnetic tension", which can balance variations in the total plasma pressure

$$\frac{d}{dr} \left( p + \frac{B_\theta^2}{8\pi} \right) + \frac{B_\theta^2}{r} = 0$$

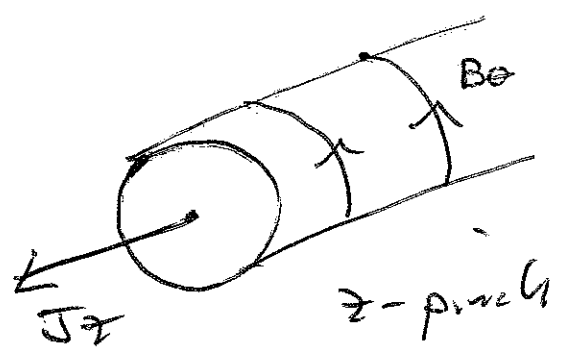
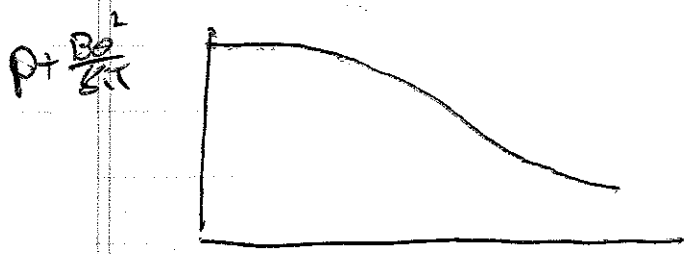
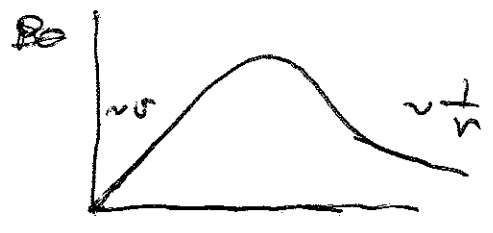
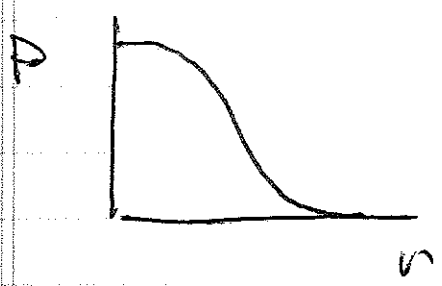
$$(\nabla \times B_\theta \theta)_z = \frac{4\pi}{c} J_z$$

⇒ current along z ⇒ z-pinch

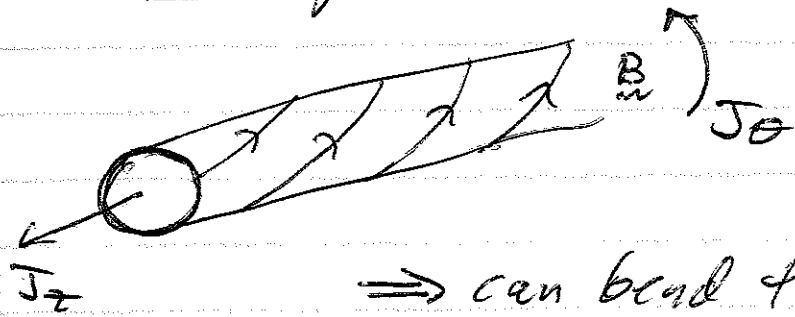
$$(\nabla \times r B_\theta \nabla \theta)_z = \frac{4\pi}{c} J_z$$

$$\underbrace{\left( \nabla (r B_\theta) \times \nabla \theta \right)}_z = \frac{4\pi}{c} J_z$$

$$\frac{1}{r} \frac{d}{dr} r B_\theta = \frac{4\pi}{c} J_z$$



A screw pinch has both  $B_\theta, B_z \neq 0$

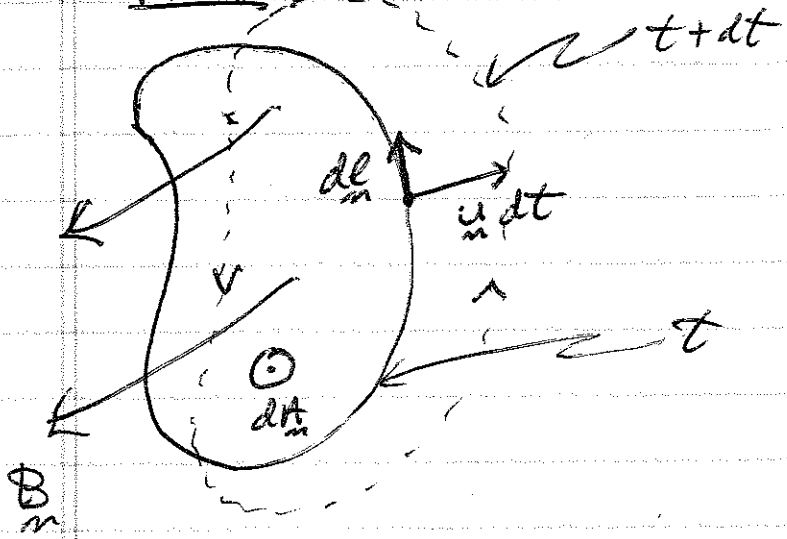


$\Rightarrow$  can bend this into a torus?  
~~Addressed later~~

The "frozen-in" theorem

In ideal MHD plasma, a plasma moves so that the magnetic flux through any connected fluid element is conserved.  $\Rightarrow$  not valid if have dissipation  $\Rightarrow$  e.g.  $\eta$  resistivity

Proof:



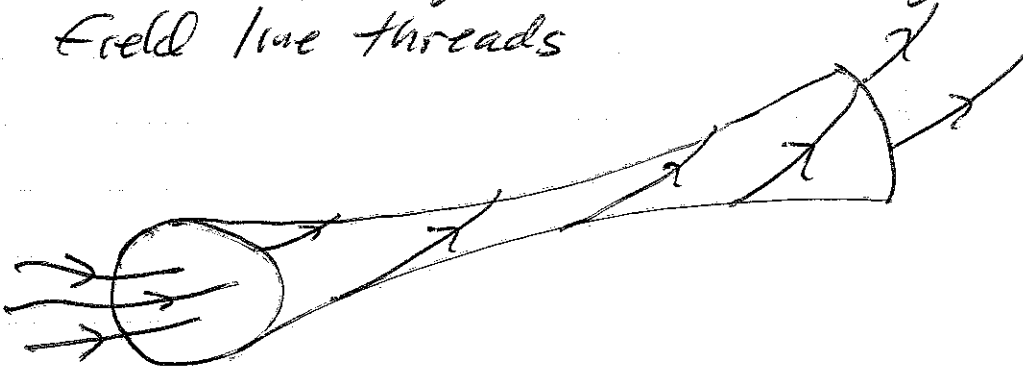
$$\Phi = \oint B_n \cdot dA$$

$$\frac{d\Phi}{dt} = \underbrace{\oint \frac{\partial B}{\partial t} \cdot dA}_{\text{change due to } \frac{\partial B}{\partial t}} - \underbrace{\frac{1}{dt} \oint dl_m \times u dt \cdot B_m}_{\text{change due to motion of loop}}$$

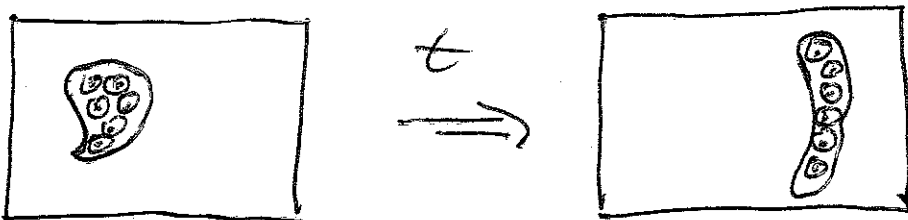
$$\begin{aligned} \dot{\Phi} &= \oint_{\partial V} \vec{B} \cdot d\vec{A} - \oint \vec{u} \times \vec{B} \cdot d\vec{l} \\ &= -c \oint \nabla \times \vec{E} \cdot d\vec{A} - \oint \vec{u} \times \vec{B} \cdot d\vec{l} \\ &= -c \oint (\vec{E} + \frac{1}{c} \vec{u} \times \vec{B}) \cdot d\vec{l} \\ &= 0 \quad \text{since} \quad \vec{E} = -\frac{1}{c} \vec{u} \times \vec{B} \end{aligned}$$

Note that must have  $E_{||} = 0$ .

~~Define a flux tube as a surface~~  
A "flux tube" is defined by a bounding surface through which no magnetic field line threads



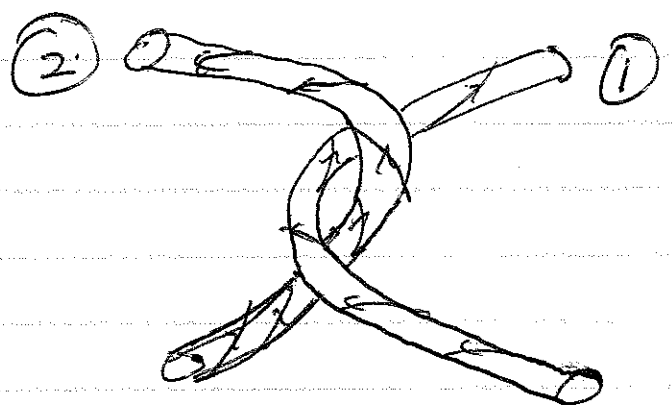
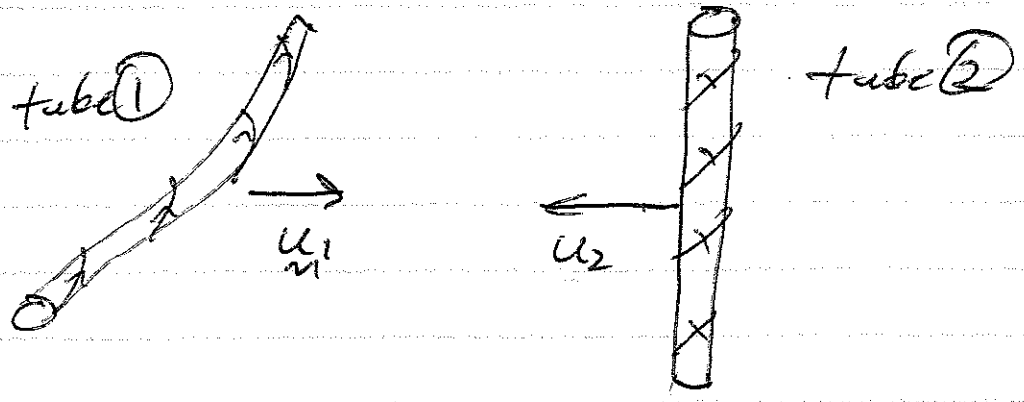
Corollary 1: when a plasma moves, it carries its flux tube with it.



Consider region of localized field and surrounding plasma. As the plasma moves the field lines move with ~~the~~ it.

Congruency 2

Two "crossed" flux tubes can not pass through each other.



$\Rightarrow$  flux tubes do not change topology