

Physics 401 Homework1---Due September 14

- 1) In class it was stated that the spectral density for black-body radiation was given by Planck's

formula:
$$\rho(\nu, T) = \left(\frac{8\pi}{c^3} \right) \frac{\nu^2}{e^{\frac{h\nu}{kT}} - 1} .$$

- a) We argued in class that for low frequencies ($\frac{h\nu}{kT} \ll 1$) this should reduce to the classical statistical mechanical result of Rayleigh and Jeans----

$$\rho_{R-J}(\nu, T) = \frac{8\pi\nu^2}{c^3} . \text{ Show that it does.}$$

- b) Prior to the years immediately preceding the discovery of the Planck formula data for black-body radiation was restricted to relatively high frequencies ($\frac{h\nu}{kT} \gg 1$)

and was found to empirically be fit by Wien's law $\rho(\nu, T) = a\nu^3 e^{-\frac{b\nu}{kT}}$, where a and b are parameters. Starting with the Planck law, find the parameters a and b in terms of fundamental constants.

- 2) Using the conservation of energy and momentum and assuming a two-body collision between a photon and an electron initially at rest derive the Compton

formula:
$$\lambda_s - \lambda_i = \frac{h}{mc} (1 - \cos(\theta))$$
 where the subscripts i and s indicate the

wavelength of the incident and scattered radiation respectively. (Hint: You may want to use relativity to first solve the problem in the center of mass frame and then boost).