Introduction to Nonequilibrium Statistical Physics, C. Jarzynski, Spring 2020

## A few useful mathematical tools

These notes provide a brief, informal, and non-rigorous summary of some mathematical concepts and tools that will be useful in this course. The main topics covered are: the moments and cumulants of a probability distribution, the law of large numbers, the central limit theorem, large deviation theory, and Poisson processes.

Throughout these notes, f(x) denotes a probability distribution of a real variable x. We assume f(x) is normalized to unity,  $\int_{-\infty}^{+\infty} f(x) dx = 1$ , and all of its cumulants (defined below) have well-defined, non-infinite values. We refer to x as a random variable, which is sampled from the distribution f(x). Angular brackets  $\langle \cdots \rangle$  denote averages with respect to values of x sampled from this distribution.

The moments of the distribution are averages of integer powers of x:

$$\mu_n = \langle x^n \rangle = \int dx f(x) x^n \quad , \quad n = 0, 1, 2, \cdots$$
 (1)

 $\mu$  (without a subscript) and  $\sigma^2$  will denote the mean and variance of the distribution:

$$\mu = \mu_1 \qquad , \qquad \sigma^2 = \mu_2 - \mu_1^2 \tag{2}$$

The cumulant-generating function, or simply generating function, is given by

$$g(\lambda) = \ln \langle e^{\lambda x} \rangle = \ln \int dx \, f(x) e^{\lambda x}$$
 (3)

where  $\lambda$  is a real number. The *cumulants*  $\kappa_m$  are defined through the expansion

$$g(\lambda) = \sum_{m=1}^{\infty} \kappa_m \frac{\lambda^m}{m!} = \kappa_1 \lambda + \frac{1}{2} \kappa_2 \lambda^2 + \cdots$$
 (4)

(Note that g(0) = 0.) Equivalently,

$$\kappa_m = \left(\frac{\mathrm{d}^m g}{\mathrm{d}\lambda^m}\right)_{\lambda=0} \quad , \quad m \ge 1 \tag{5}$$

Computing the first two derivatives of  $g(\lambda)$  (Eq. 3),

$$g'(\lambda) = \frac{\mathrm{d}g}{\mathrm{d}\lambda} = \frac{\int f(x) x e^{\lambda x}}{\int f(x) e^{\lambda x}} \tag{6}$$

$$g''(\lambda) = \frac{\mathrm{d}^2 g}{\mathrm{d}\lambda^2} = \frac{\int f(x) \, x^2 \, e^{\lambda x}}{\int f(x) \, e^{\lambda x}} - \left[ \frac{\int f(x) \, x \, e^{\lambda x}}{\int f(x) \, e^{\lambda x}} \right]^2 \tag{7}$$

and evaluating them at  $\lambda = 0$  gives us

$$\kappa_1 = g'(0) = \mu_1 \quad , \quad \kappa_2 = g''(0) = \mu_2 - \mu_1^2$$
(8)

Thus the first and second cumulants are the mean,  $\kappa_1 = \mu$ , and variance,  $\kappa_2 = \sigma^2$ . Higher cumulants such as the *skewness*,  $\kappa_3$  and the *kurtosis*,  $\kappa_4$ , can similarly be evaluated.

Scaling property: if y = ax, where a > 0 is a constant, then

$$\mu_m[y] = a^m \,\mu_m[x] \quad , \quad \kappa_m[y] = a^m \,\kappa_m[x] \tag{9}$$

The notation  $\mu_m[y]$  denotes the m'th moment of the random variable y, etc.

The cumulants (but not the moments) satisfy the useful property

$$\kappa_m[X] = N \,\kappa_m[x] \tag{10}$$

where  $X = \sum_{n=0}^{N} x_n$  is the sum of N independent samples from the distribution f(x). This can be seen by comparing the corresponding generating functions:

$$G(\lambda) = \ln\langle e^{\lambda X} \rangle = \ln\langle e^{\lambda x_1} e^{\lambda x_2} \cdots e^{\lambda x_N} \rangle = N \ln\langle e^{\lambda x} \rangle = Ng(\lambda)$$
(11)

Eq. 10 then follows from Eq. 4.

Now consider the sample mean,  $\bar{x} = N^{-1} \sum_{n=1}^{N} x_n = X/N$ , which is the average over N independent samples from f(x). Let  $f_N(\bar{x})$  denote the probability distribution of the sample mean. Combining Eqs. 9 and 10 we get

$$\kappa_m[\bar{x}] = \frac{1}{N^m} \kappa_m[X] = \frac{1}{N^{m-1}} \kappa_m[x] \tag{12}$$

For the mean and variance, this gives us the familiar (hopefully!) results

$$\mu_{\bar{x}} = \mu_x \equiv \mu \qquad , \qquad \sigma_{\bar{x}}^2 = \frac{1}{N} \sigma_x^2$$
 (13)

Thus as  $N \to \infty$ ,  $f_N(\bar{x})$  becomes infinitely narrow:

$$\lim_{N \to \infty} f_N(\bar{x}) = \delta(\bar{x} - \mu) \tag{14}$$

This is the law of large numbers (LLN), which states that the sample mean  $\bar{x}$  converges to the ensemble mean  $\mu$  in the limit of infinitely many samples.

## Central limit theorem (CLT)

The CLT states, roughly, that  $f_N(\bar{x})$  looks like a Gaussian for sufficiently large N. For a Gaussian distribution,  $\kappa_m = 0$  for all  $m \geq 3$ . (Proof assigned as exercise!) Moreover, any distribution that is *not* a Gaussian has infinitely many non-zero cumulants (Marcinkiewicz, 1939). To establish the CLT, consider a distribution f(x) with mean  $\mu = 0$  and variance  $\sigma^2 > 0$ . Now define a "rescaled" sample mean,

$$y = \sqrt{\frac{N}{\sigma^2}} \,\bar{x} \tag{15}$$

where  $\bar{x} = N^{-1} \sum_{n=1}^{N} x_n$  as before. Using Eqs. 9 and 10 we get

$$\kappa_m[y] = \frac{1}{N} \left(\frac{N}{\sigma^2}\right)^{m/2} \kappa_m[x] \tag{16}$$

hence as  $N \to \infty$  we get  $\kappa_2[y] = 1$ , and  $\kappa_m[y] \to 0$  for all  $m \ge 3$ . Thus the distribution of the variable y converges toward a Gaussian whose variance is unity. From Eq. 15 we conclude that the distribution of the sample mean  $\bar{x}$  tends toward a Gaussian of variance  $\sigma^2/N$ .

Dropping the assumption  $\mu = 0$ , we summarize LLN and CLT as follows. As  $N \to \infty$ ,  $f_N(\bar{x})$  becomes ever sharper and ever more Gaussian:

$$f_N(\bar{x}) \approx \sqrt{\frac{N}{2\pi\sigma^2}} \exp\left[-\frac{N(\bar{x}-\mu)^2}{2\sigma^2}\right]$$
 (17)

## Large deviation theory (LDT)

Both LLN and CLT follow naturally from large deviation theory. For our purposes, LDT is a set of tools built around the following statement. For a well-behaved distribution f(x), the distribution of the sample mean  $\bar{x}$  obeys

$$f_N(\bar{x}) \sim e^{-NI(\bar{x})}$$
 as  $N \to \infty$  (18)

The notation  $\sim$  indicates that Eq. 18 captures the dominant dependence on N. That is, if  $f_N(\bar{x}) = \exp{-N[I(\bar{x}) + \text{other terms}]}$ , then the "other terms" are *subdominant*, i.e. they vanish as  $N \to \infty$ .<sup>2</sup> Somewhat more precisely, LDT states that the limit

$$\lim_{N \to \infty} -\frac{1}{N} \ln f_N(\bar{x}) \equiv I(\bar{x}) \tag{19}$$

<sup>&</sup>lt;sup>1</sup> There is no loss of generality in assuming  $\mu = 0$ . If  $\mu \neq 0$  then we can define a new random variable  $x' = x - \mu$  whose mean is zero, and which is trivially related to x by a simple shift.

 $<sup>^{2}</sup>$  Subdominant terms might arise from an N-dependent normalization factor, such as the one in Eq. 17.

exists. The definition of "well-behaved" is a bit logically circular: if the above limit exists, then we say that f(x) "satisfies the large deviation principle". In practice, most of the distributions we will encounter in this course will satisfy this principle.

The large deviation function (or Cramér function)  $I(\bar{x})$  has these properties:

- 1. It is *convex*, that is  $I''(\bar{x}) \geq 0$  for all  $\bar{x}$ .
- 2. The unique minimum of  $I(\bar{x})$  occurs at  $\bar{x} = \mu$ , where  $I(\mu) = 0$ .
- 3.  $I''(\mu) = 1/\sigma^2$ .

As  $N \to \infty$ , the distribution  $f_N(\bar{x}) \sim e^{-NI(\bar{x})}$  becomes negligible except in the immediate vicinity of the minimum at  $\bar{x} = \mu$ . Expanding around this minimum, we have

$$I(\bar{x}) \approx \frac{(\bar{x} - \mu)^2}{2\sigma^2} \tag{20}$$

using properties 2 and 3. Combining with Eq. 18 we see that at large N,  $f_N(\bar{x})$  becomes approximately a Gaussian of mean  $\mu$  and variance  $\sigma^2/N$ , in agreement with Eq. 17.

Note that Eq. 20 is valid only near  $\bar{x} = \mu$ , in other words  $f_N(\bar{x})$  is Gaussian only in this "central" region. The tails of the distribution are described by the shape of  $I(\bar{x})$  away from this minimum. In other words, the central limit theorem does not describe these tails, but only the central region near  $\bar{x} = \mu$ , where most of the probability resides.

The function  $I(\bar{x})$  can be computed either directly from Eq. 19, or else via the Legendre transform of the generating function:

$$I(\bar{x}) = \max_{\lambda} \{\lambda \bar{x} - g(\lambda)\} = \lambda^* \bar{x} - g(\lambda^*) = \text{Legendre transform of } g(\lambda)$$
 (21)

where  $\lambda^* = \lambda^*(\bar{x})$  is defined by the condition  $g'(\lambda^*) = \bar{x}$ . I won't derive Eq. 21 here. For a pedagogical introduction to Legendre transforms, see Zia, Redish and McKay, "Making sense of the Legendre transform",  $Am.\ J.\ Phys.\ 77,\ 614\ (2009)$ .

Recall that g(0) = 0 and  $g'(0) = \mu$ , and note that  $g''(\lambda) \geq 0$  for all  $\lambda$  (this follows from Eq. 7). Exercise: using these properties of the generating function  $g(\lambda)$ , show that its Legendre transform  $I(\bar{x})$  satisfies properties 1-3 listed above.

Large deviation theory plays an important role in both equilibrium and nonequilibrium statistical physics, as well as in statistics, finance and other fields. A review of LDT can be found in Touchette, "The large deviation approach to statistical mechanics", *Phys. Reports* **478**, 1 (2009).

## Poisson processes

In a Poisson process, some event occurs repeatedly, with a fixed probability rate, r. That is, during every infinitesimal time interval  $\delta t$ , the probability that an event occurs is  $r \cdot \delta t$ . An example is the emission of  $\alpha$ -particles from a radioactive sample with a long half-life. Each emission of an  $\alpha$ -particle is an event. If our period of observation is much shorter than the half-life, we can model these events as a Poisson process.

For a Poisson process with rate r, let's compute  $p(n;\tau)$ , the probability to observe exactly n events during a given time interval of duration  $\tau$ . We divide the interval into  $K \gg 1$  sub-intervals of duration  $\delta t = \tau/K$ . By considering all the ways that n events can be distributed among K bins, we have

$$p(n;\tau) = \lim_{K \to \infty} \frac{K!}{n!(K-n)!} (r \cdot \delta t)^n (1 - r \cdot \delta t)^{K-n}$$
(22)

$$= \frac{1}{n!} \lim_{K \to \infty} \left[ \frac{K(K-1)\cdots(K-n+1)}{K^n} \right] \alpha^n \left( 1 - \frac{\alpha}{K} \right)^{K-n}$$
 (23)

$$= \frac{\alpha^n}{n!} \lim_{K \to \infty} \left[ \left( 1 - \frac{\alpha}{K} \right)^K \right]^{1 - (n/K)} \tag{24}$$

where  $\alpha \equiv r\tau$ . If we now use  $\lim_{N\to\infty} [1-(x/N)]^N = e^{-x}$ , we arrive at the result

$$p(n;\tau) = \frac{\alpha^n}{n!} e^{-\alpha}$$
 ,  $\alpha = r\tau$  (25)

It is easy to verify that normalization is satisfied:  $\sum_{n=0}^{\infty} p(n;\tau) = 1$ .

To compute the generating function  $g(\lambda)$ , we first evaluate

$$\langle e^{\lambda n} \rangle = \sum_{n=0}^{\infty} e^{\lambda n} \frac{\alpha^n}{n!} e^{-\alpha} \tag{26}$$

$$= \sum_{n=0}^{\infty} \frac{\left(\alpha e^{\lambda}\right)^n}{n!} \cdot e^{-\alpha e^{\lambda}} \cdot e^{\alpha \left(e^{\lambda} - 1\right)}$$
(27)

$$= e^{\alpha(e^{\lambda} - 1)} \sum_{n=0}^{\infty} \frac{\beta^n}{n!} e^{-\beta} \qquad , \qquad \beta = \alpha e^{\lambda}$$
 (28)

Since the sum on the last line converges to unity, we arrive at

$$g(\lambda) = \ln \langle e^{\lambda n} \rangle = \alpha \left( e^{\lambda} - 1 \right)$$
 (29)

Eq. 5 then gives us

$$\kappa_m[n] = \alpha \qquad , \qquad m = 1, 2, \cdots$$
(30)

<sup>&</sup>lt;sup>3</sup> We can motivate this by noting that  $\ln[1-(x/N)]^N=N\ln[1-(x/N)]\to -x$  as  $N\to\infty$ .

Thus all of the cumulants have the same value,  $\alpha = r\tau$ . In particular the mean and variance of the number of events occurring in the interval  $\tau$  are

$$\mu = \langle n \rangle = r\tau$$
 ,  $\sigma^2 = \langle n^2 \rangle - \langle n \rangle^2 = r\tau$  (31)

The result  $\langle n \rangle = r\tau$  makes sense; we could have guessed this without doing any calculations. The quantity  $\sigma = \sqrt{r\tau}$  characterizes the width of the distribution  $p(n;\tau)$ , thus it provides a measure of the size of typical fluctuations in n. For very long time intervals,  $r\tau \gg 1$ , both  $\langle n \rangle$  and  $\sigma$  become large, but the *relative* size of the fluctuations  $\sigma/\langle n \rangle = 1/\sqrt{r\tau}$ , goes to zero, in agreement with the law of large numbers.

Here's how to simulate a Poisson process using a standard random number generator:

- 1. Set  $t_0 = 0$  and k = 0.
- 2. Generate a random number  $\xi$  between 0 and 1.
- 3. Set

$$t_{k+1} = t_k - \frac{\ln \xi}{r} \tag{32}$$

4. Update  $k \to k+1$  and go back to step 2.

The sequence  $t_1, t_2, t_3, \cdots$  will represent the times at which the events occur.

Exercise: Convince yourself that this algorithm generates a Poisson process.