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Rayleigh - Taylor instability

Let's add gravity to our system of eqn's

$$\rho \frac{d\mathbf{v}}{dt} = -\nabla p + \rho \mathbf{g} + \frac{\mathbf{J} \times \mathbf{B}}{c}$$

$$\frac{dp}{dt} + \rho \nabla \cdot \mathbf{v} = 0$$

$$\frac{4\pi}{c} \mathbf{J} = \nabla \times \mathbf{B}$$

$$\frac{dp}{dt} + \rho \nabla \cdot \mathbf{v} = 0$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{v} \times \mathbf{B})$$

$\mathbf{g}$ = acceleration due to gravity

$\rho \mathbf{g}$ = force density

• why add gravity

equilibrium

$$+\nabla \left(p + \frac{B^2}{8\pi} \right) = \frac{B^2}{4\pi} \mathbf{k} + \rho \mathbf{g}$$

gravity simulates the effect of curved field lines w/out having complications of geometry

equilibrium (x, y, z)

$$\underline{B} = (0, 0, B_0)$$

$$p = p(x) \quad \swarrow$$

$$\underline{J} = 0$$

$$\rho = \rho(x)$$

$$\underline{g} = (g, 0, 0) \quad \swarrow x$$

$$\frac{dp}{dx} = \rho g \quad \text{Equal}$$

let's assume incompressible and eliminate

p equation

$$\nabla \cdot \underline{\xi} = 0$$

instead of $p_i + \underline{\xi} \cdot \nabla p_0 + \rho p \nabla \cdot \underline{\xi} = 0$

$$\underline{\xi} = (\xi_x, \xi_y, 0)$$

assume $\rho \gg 1$
 $\frac{\rho \delta \cdot \nabla p}{\rho p}$
 $\nabla \cdot \underline{\xi} \approx \frac{\rho \delta \cdot \nabla p}{\rho p}$

all quantities vary as $\exp[ik_y y - i\omega t]$

$$\underline{B}_1 = \nabla \times \underline{\xi} \times \underline{B}_0$$

$$= \underline{\xi} \nabla \cdot \underline{B}_0 + \underline{B}_0 \cdot \nabla \underline{\xi} - \underline{B}_0 \nabla \cdot \underline{\xi} - \underline{\xi} \cdot \nabla \underline{B}_0$$

$$= 0 \quad \text{no perturbed magnetic field}$$

$$\underline{J}_1 = 0$$

$$\nabla \times (\underline{\xi} \times \underline{B})$$

~~Equation of momentum balance~~

Equation of momentum balance

$$-\omega^2 \rho \underline{\xi} = -\nabla p_1 + \rho_1 \underline{g}$$

$\hat{z} \cdot \nabla \times ()$ elm

to eliminate p_1 $\hat{z} \cdot \nabla \times \{ \text{above} \}$

$$\hat{z} \cdot \nabla \times (-\omega^2 \rho \underline{\xi}) = \frac{d}{dx} (-\omega^2 \rho \xi_y) + ik \omega^2 \rho \xi_x$$

but $ik \xi_y + \frac{d}{dx} \xi_x = 0$ $\xi_y = -\frac{1}{ik} \frac{d}{dx} \xi_x$

$$\frac{\omega^2}{ik} \left[\frac{d}{dx} \rho \frac{d}{dx} \xi_x - k^2 \rho \xi_x \right] = \hat{z} \cdot \nabla \times (-\omega^2 \rho \underline{\xi})$$

O.K.

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$$\vec{z} \cdot \nabla \times (-\nabla \tilde{\phi}_1 + \rho_1 \vec{g})$$

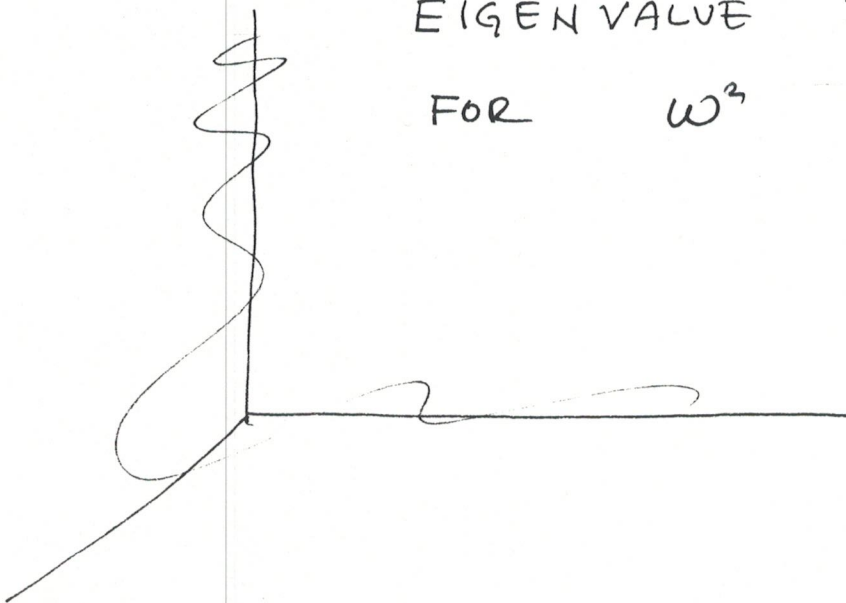
$$= \vec{z} \cdot \nabla \times \rho_1 \vec{g} = -ikg \tilde{\rho}_1$$

$$\boxed{\tilde{\rho}_1 + \tilde{z}_x \frac{\partial \rho_0}{\partial x} = 0}$$

FINALLY

$$\omega^2 \left[\frac{d}{dx} \rho_0 \frac{d}{dx} \tilde{z}_x - k^2 \rho_0 \tilde{z}_x \right] = -k^2 g \frac{\partial \rho_0}{\partial x} \tilde{z}_x$$

EIGENVALUE EQUATION
FOR ω^2



LETS NEGLECT THE x derivative and
solve locally $k^2 \gg \frac{d^2}{dx^2}$

$$\omega^2 = \frac{g \frac{\partial \rho_0}{\partial x}}{\rho_0} = \frac{g}{r_n}$$

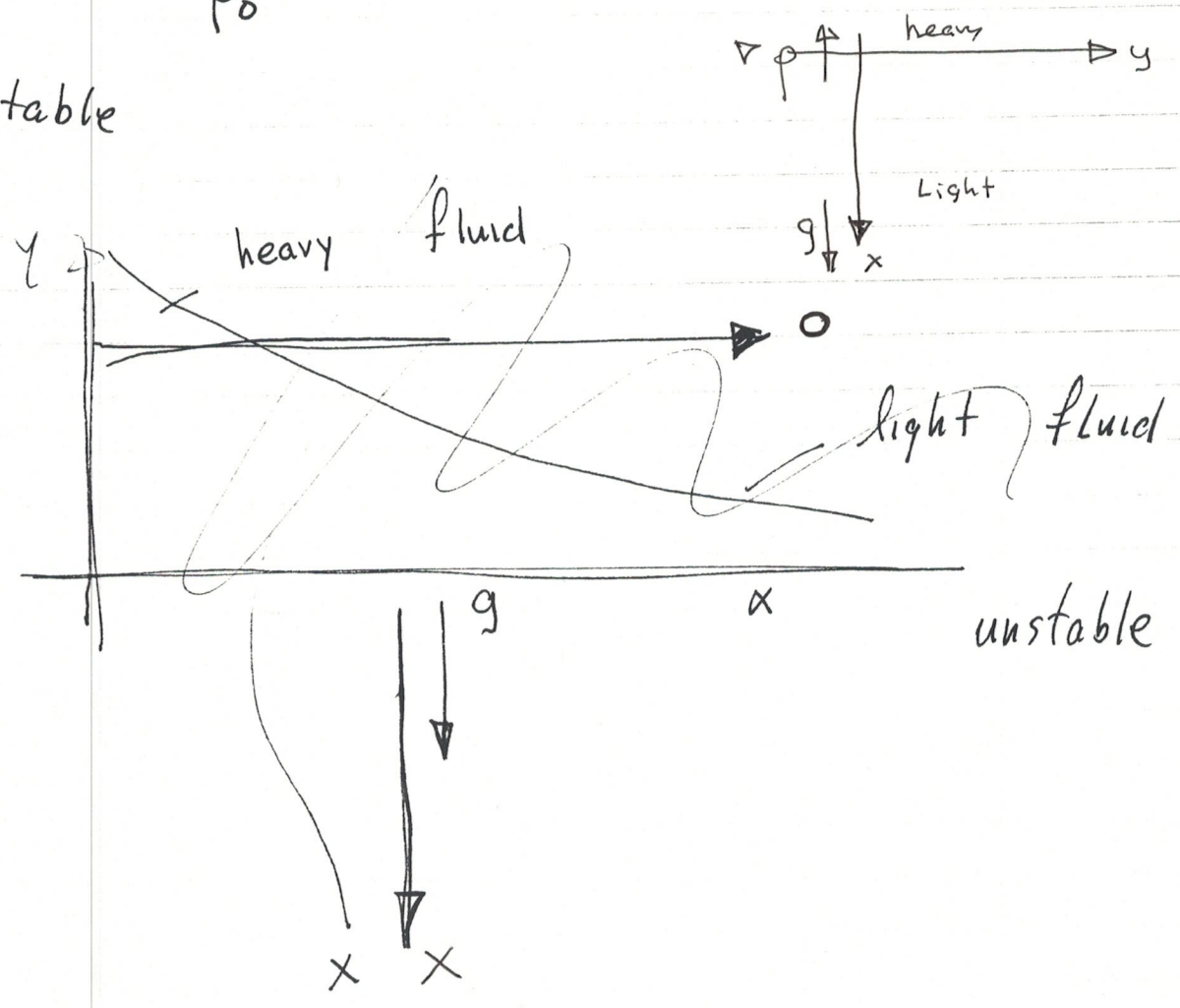
IF

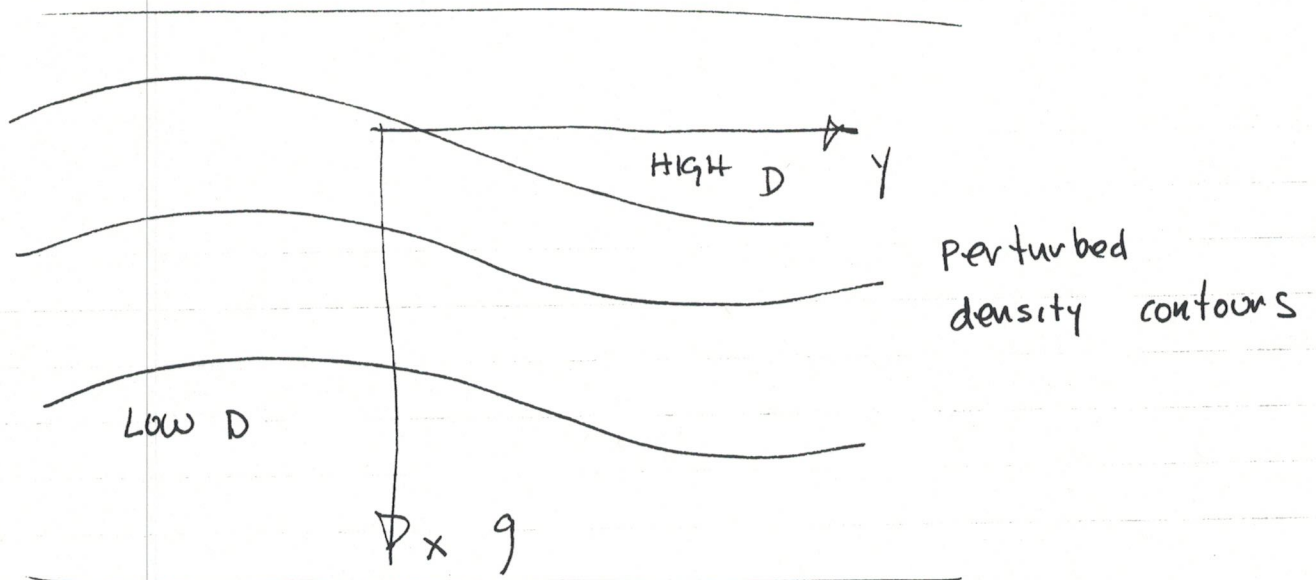
$$\frac{g}{\rho_0} \frac{\partial \rho_0}{\partial x} < 0$$

$$\omega^2 < 0$$

heavy fluid supported by light fluid

unstable





system moves to a lower energy state

Quadratic form

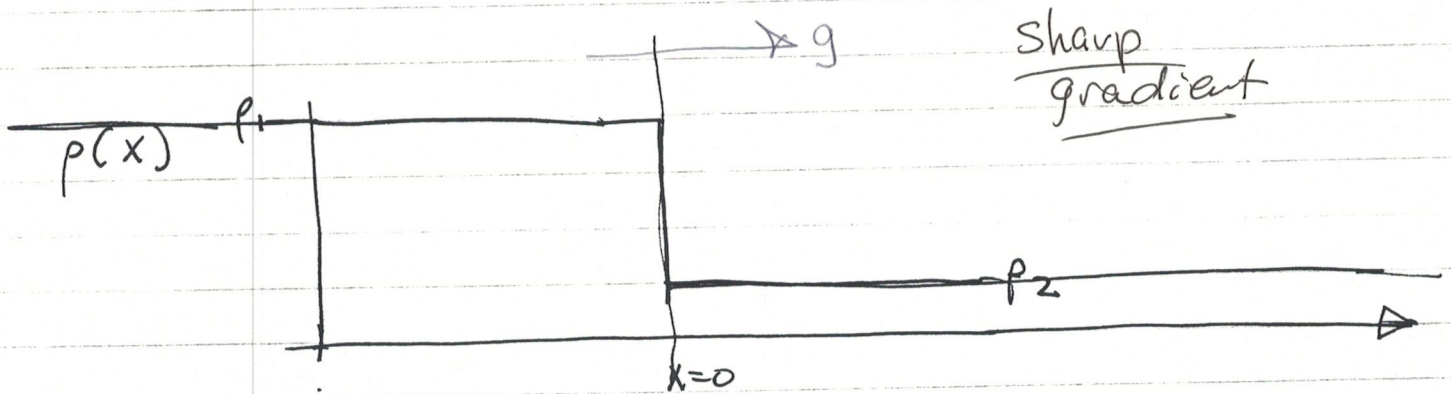
$$\int_{x_1}^{x_2} dx \bar{z}_x \omega^2 \left[\frac{d}{dx} \rho_0 \frac{d}{dx} \bar{z}_x - k^2 \rho_0 \bar{z}_x \right] = - \int_{x_1}^{x_2} k^2 g \frac{d\rho_0}{dx} \bar{z}_x^2 dx$$

$$\omega^2 = \frac{\int_{x_1}^{x_2} g dx k^2 g \frac{d\rho_0}{dx} \bar{z}_x^2}{\int_{x_1}^x dx \rho_0 \left[\left(\frac{d\bar{z}_x}{dx} \right)^2 + k^2 \bar{z}_x \right]}$$

a variation WRT TO $\bar{z}_x$ reproduces

THE ω^2 are extreme values of

Two limits are interesting



for $x > 0$

$$p_2 \omega^2 \left(\frac{d^2}{dx^2} - k^2 \right) \xi_x = 0$$

$$\xi_x = \xi_x(0+) \exp(-kx)$$

for $x < 0$

$$p_1 \omega^2 \left(\frac{d^2}{dx^2} - k^2 \right) \xi_x = 0$$

$$\xi_x = \xi_x(0-) \exp(kx)$$

ξ is continuous

$$\xi_x(0+) = \xi_x(0-)$$

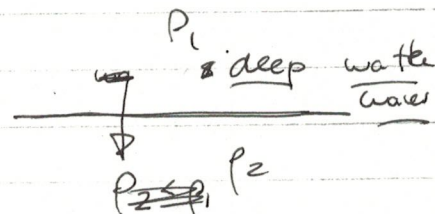
integrate
across discontinuity

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$$\omega^2 \left[\rho_2 \frac{d\zeta_x}{dx} \Big|_{0+} - \rho_1 \frac{d\zeta_x}{dx} \Big|_{0-} \right] = -k^2 g (\rho_2 - \rho_1) \zeta_x(0)$$

$$-k \zeta_x(0) - k \zeta_x(0)$$

$$\omega^2 = \frac{+ k^2 g (\rho_2 - \rho_1)}{\rho_2 + \rho_1}$$

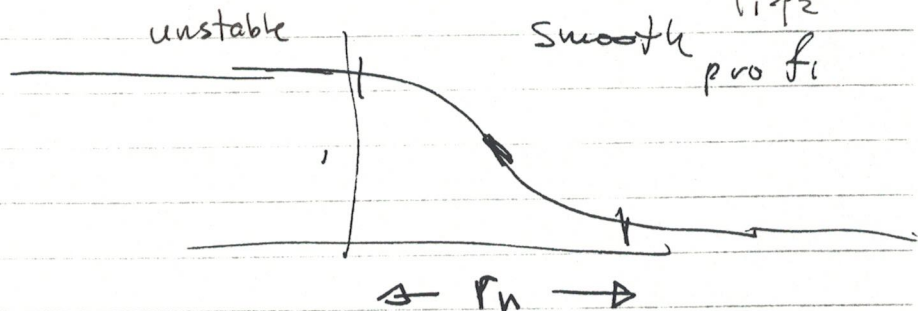


$\rho_2 > \rho_1$

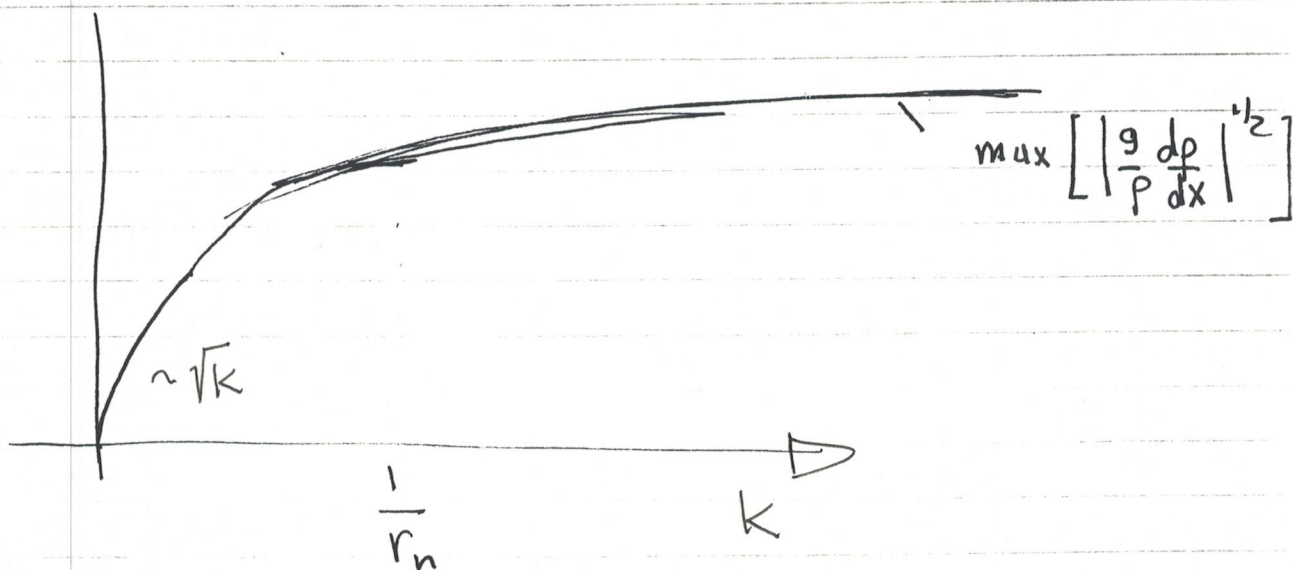
$\omega \sim \sqrt{k}$ Rayleigh
 $\rho_1 > \rho_2$
smooth profile

IF $\rho_2 < \rho_1$

$$\gamma = \sqrt{\frac{g k \Delta \rho}{2 \bar{\rho}}}$$



$\gamma(k)$



notice all this happened and

B_0 was ~~powerless~~ powerless to

stop the instability magnetic field

In fact perturbation had ~~no~~ no
perturbed magnetic field ~~per~~ "flute
instability"

$$(\xi_x, \xi_y, 0)$$

Lets include a $k_{||} B_0$

$$\nabla \cdot \xi = 0$$

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$$\underline{B}_1 = \underline{B}_0 \cdot \nabla \xi = i k_{||} B_0 \xi_{x/y}$$

$$\underline{J}_1 = \frac{c}{4\pi} \nabla \times \underline{B}_1 = \frac{c}{4\pi} \underline{J}_{1z} = \frac{c}{4\pi} \left[\frac{d}{dx} B_{1y} - i k B_{1x} \right]$$

$$\begin{aligned} \hat{z} \cdot \nabla \times \frac{\underline{J}_1 \times \underline{B}_0}{c} &= \nabla \cdot \left(\frac{\underline{J}_1 \times \underline{B}_0}{c} \right) \times \hat{z} = \nabla \cdot \left(\frac{B_0 \underline{J}_1}{c} \right) \\ &= \frac{B_0 i k_{||}}{4\pi} \left[\frac{d}{dx} \underbrace{i k_{||} B_0 \xi_y}_{B_y} - i k \underbrace{i k_{||} B_0 \xi_x}_{B_{1x}} \right] \end{aligned}$$

$\nabla \cdot \underline{J}_1 = 0$

$$i k \xi_y + \frac{d}{dx} \xi_x = 0$$

$$\xi_y = -\frac{1}{i k} \frac{d}{dx} \xi_x$$

$$\hat{z} \cdot \nabla \times \frac{\underline{J}_1 \times \underline{B}_0}{c} = -i \frac{k_{||}^2 B_0^2}{4\pi k} \left[\frac{d^2}{dx^2} \xi_x - k^2 \xi_x \right]$$

(k)

$$\frac{d}{dx} \left(\omega_p^2 - \frac{k_{\parallel}^2 B_0^2}{4\pi} \right) \frac{d}{dx} \xi_x - k^2 \left(\omega_p^2 - \frac{k_{\parallel}^2 B_0^2}{4\pi} \right) \xi_x = - \underbrace{k^2 g \frac{\partial \rho_0}{\partial x}}_{\text{instability}} \xi_x$$

notice shear Alfvén dispersion relations

Local limit $k^2 \gg \frac{d}{dx}$
shear Alfvén wave

$$\omega_p^2 - \frac{k_{\parallel}^2 B_0^2}{4\pi} = - k^2 g \frac{\partial \rho_0}{\partial x}$$

unstable if $\frac{k_{\parallel}^2 B_0^2}{4\pi} + k^2 g \frac{\partial \rho_0}{\partial x} < 0$

$\frac{k_{\parallel}^2 B_0^2}{4\pi}$

$k^2 g \frac{\partial \rho_0}{\partial x}$

energy released
by rearing mass
density

stabilizing

term due to

field line bending

Quadratic Form

(12)

ξ_x times above & integrate

$$+ \omega^2 \int dx \rho \left[\left(\frac{d\xi_x}{dx} \right)^2 + k^2 \xi_x^2 \right]$$

Kinetic energy of flow

Bending energy

$$= \int dx \left[\frac{k_{\parallel}^2 B_0^2}{4\pi} \left[\left(\frac{d\xi_x}{dx} \right)^2 + k^2 \xi_x^2 \right] + k^2 g \frac{\partial \rho_0}{\partial x} \xi_x^2 \right]$$

energy released

by gravity + displ

Energy principle