

Fitzpatrick Kinetic.pdf
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Electrostatic Waves in a warm Plasma

Recall for a cold plasma we found

$$\epsilon = 1 - \sum_{e,i} \frac{\omega_{pe,i}^2}{\omega^2}$$

where $\omega_{pe,i}^2 = \frac{4\pi N_{e,i} q_{e,i}^2}{m_{e,i}}$

$$\omega_{pe}^2 \gg \omega_{pi}^2$$

normal electrostatic mode $\epsilon = 0$ $\omega \approx \pm \omega_{pe}$

* We will investigate the case of non-zero Temperature

* How will we do this

* Solve Vlasov Equation

(20)
(14)

Step #1

find an equilibrium solution in which $f_{e,i}$ is independent of time

we will assume $E = 0$ in equilibrium

¶

$$\sum_{e,i} q_{e,i} n_{e,i,0} = 0$$

CHARGE NEUTRALITY

STEP #2

We will assume a small electric field is present in the form of a wave

$$\tilde{E} = -\nabla \phi$$

$$\phi = \text{Re} \left\{ \hat{\phi} e^{i(k \cdot \vec{x} - \omega t)} \right\}$$

$\hat{\phi}$ = complex amplitude

k = vector wave number

ω = frequency

STEP #3

We will solve the

linearized VE to

find the perturbed distribution function

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$$f(x, \omega) = f_0 + \text{Re} \left\{ \hat{f} e^{i(k-x-\omega t)} \right\}$$

↑ equilibrium
↓ perturbation

assume $|\hat{f}| \ll |f_0|$

Step #4

We will compute the perturbed charge density

$$\rho = \text{Re} \left\{ \hat{\rho} e^{i(k-x-\omega t)} \right\}$$

Step #5

We will solve the Poisson equation

$$-\nabla^2 \cdot \text{Re} \left\{ \hat{\phi} e^{i(k-x-\omega t)} \right\} = 4\pi \text{Re} \left\{ \hat{\rho} e^{i(k-x-\omega t)} \right\}$$

From which we will extract $\epsilon(\omega, k)$

Step #1

Find equilibrium solution
corresponding to $E = 0$

$$\frac{\partial f}{\partial t} + v \cdot \frac{\partial f}{\partial x} + \frac{q}{m} E \cdot \frac{\partial f}{\partial v} = 0$$

$$f = f_0 + \text{Re} \left\{ \hat{f} e^{i(kx - vt)} \right\}$$

$$E = -\nabla \phi \quad \phi = \text{Re} \left\{ \hat{\phi} e^{i(kx - vt)} \right\}$$

Equilibrium

$$\left. \begin{array}{l} \frac{\partial f_0}{\partial t} = 0 \\ \frac{\partial f_0}{\partial x} = 0 \\ E = 0 \end{array} \right\} \quad \text{arbitrary function of velocity}$$

$$f_0 = f_0(v)$$

arbitrary function of velocity

independent of space

$$n_0 = \int d^3 v f_0$$

we require

$$\sum_{e,i} q_{e,i} \int d^3 v f_{e,i,0} = 0$$

Step #2

$$\underline{E} = -\nabla \phi$$

$$\phi = \operatorname{Re} \left\{ \hat{\phi} e^{i(\underline{k} \cdot \underline{x} - \omega t)} \right\}$$

$$\underline{E} = \operatorname{Re} \left\{ -ik \hat{\phi} e^{i(\underline{k} \cdot \underline{x} - \omega t)} \right\}$$

Step #3 solve the linearized VE

$$\frac{\partial f}{\partial t} + \underline{v} \cdot \nabla f + \frac{q}{m} \operatorname{Re} \left\{ -ik \hat{\phi} e^{i(\underline{k} \cdot \underline{x} - \omega t)} \right\} \cdot \frac{\partial}{\partial \underline{x}} f = 0$$

$$f = f_0 + \text{perturbation} \quad \text{perturbation} \propto \phi$$

DROP SECOND ORDER IN ϕ TERM

$$\left(\frac{\partial}{\partial t} + \underline{v} \cdot \nabla \right) \operatorname{Re} \left\{ \hat{f} e^{i(\underline{k} \cdot \underline{x} - \omega t)} \right\} + \operatorname{Re} \left\{ -ik \frac{q\hat{\phi}}{m}, \frac{\partial f_0}{\partial \underline{x}} \right\} = 0$$

$$\operatorname{Re} \left\{ -i(\omega - \underline{k} \cdot \underline{v}) \hat{f} e^{i(\underline{k} \cdot \underline{x} - \omega t)} - ik \frac{q\hat{\phi}}{m}, \frac{\partial f_0}{\partial \underline{x}} \right\} = 0$$

$$-i(\omega - \underline{k} \cdot \underline{v}) \hat{f} - i \cancel{\frac{q\phi}{m}} \underline{k} \cdot \frac{\partial f_0}{\partial \underline{v}} = 0$$

so

$$\boxed{\hat{f} = -\frac{q}{m} \frac{\hat{\phi}}{(\omega - \underline{k} \cdot \underline{v})} - \underline{k} \cdot \frac{\partial f_0}{\partial \underline{v}}}$$

$\omega - \underline{k} \cdot \underline{v}$ = doppler shifted frequency
for particles with velocity \underline{v}

Step #4

$$p = \phi \operatorname{Re} \left\{ \hat{p} e^{i(\underline{k} \cdot \underline{x} - \omega t)} \right\}$$

$$\hat{p} = \sum_{e,i} \int d^3r q_{q,i} \hat{f}_{e,i}$$

$$\hat{\rho} = - \sum_{e,i} \int d^3v \frac{q_{e,i}^2}{m_{e,i}} \frac{k \cdot \frac{\partial f_0}{\partial v}}{(w - k \cdot v)} \hat{\phi}$$

note: $\hat{\rho}$ is proportional to $\hat{\phi}$

Step #5 Poisson Eqn.

$$-\nabla^2 \operatorname{Re} \{ \hat{\phi} e^{i(k \cdot x - wt)} \} = 4\pi \operatorname{Re} \{ \hat{\rho} e^{i(k \cdot x - wt)} \}$$

↓

$$k^2 \hat{\phi} = 4\pi \hat{\rho} = - \sum_{e,i} \int d^3v \frac{4\pi q_{e,i}}{m_{e,i}} \frac{k \cdot \frac{\partial f_0}{\partial v}}{(w - k \cdot v)} \hat{\phi}$$

OR

$$k^2 \left(1 + \frac{1}{k^2} \sum_{e,i} \int d^3v \frac{4\pi q_{e,i}}{m_{e,i}} \frac{k}{(w - k \cdot v)} \cdot \frac{\partial f_0}{\partial v} \right) \hat{\phi} = 0$$

$$\underbrace{\quad}_{\epsilon(\omega, k)}$$

$$\boxed{\epsilon(\omega, k) = 1 + \frac{1}{k^2} \sum_{e,i} \int d^3v \frac{4\pi q_{e,i}}{m_{e,i}} \frac{k \cdot \frac{\partial f_0}{\partial v}}{(w - k \cdot v)}}$$

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Reality check

FIRST do Velocity integral by parts

$$E(\omega, \vec{k}) = 1 - \frac{1}{\pi k^2} \int d^3v \frac{4\pi q_{e,i}^2}{m_{e,i}} f_0 \vec{k} \cdot \frac{\partial}{\partial \vec{v}} \frac{1}{(\omega - \vec{k} \cdot \vec{v})}$$

$$\vec{k} \cdot \frac{\partial}{\partial \vec{v}} \frac{1}{(\omega - \vec{k} \cdot \vec{v})} = - \frac{1}{(\omega - \vec{k} \cdot \vec{v})^2} \vec{k} \cdot \frac{\partial}{\partial \vec{v}} (-\vec{k} \cdot \vec{v})$$

$$\vec{k} \cdot \sum_{ij} k_i \frac{\partial}{\partial v_i} k_j v_j = \sum_{ij} k_i k_j \frac{\partial v_j}{\partial v_i}$$

$\Delta \delta_{ij} = \begin{cases} 1 & i=j \\ 0 & i \neq j \end{cases}$

$$= 2 \sum_i k_i^2 = k^2$$

$$E(\omega, \vec{k}) = 1 - \sum_{ij} \int d^3v \frac{4\pi q_{e,i}^2}{m_{e,i}} \frac{f_0(v)}{(\omega - \vec{k} \cdot \vec{v})^2}$$

Suppose plasma is cold,

$$f_0(v) \rightarrow 0 \quad \text{when} \quad |v| > V_{\text{typical}} \sim \sqrt{\frac{2T}{m}}$$

if $kV_{\text{typical}} \ll \omega$ than can neglect Doppler shift

$$\epsilon(\omega, \underline{k}) = 1 - \sum_{e,i} \frac{4\pi q_{e,i}^2}{m_e \omega^2} \int d^3v f_0(\underline{v})$$

$$= 1 - \sum_{e,i} \frac{w_{pe,i}}{\omega} \quad \text{our old result}$$

OLD Result applies when $V_{typ} \ll \frac{\omega}{|k|}$

Doppler shift is small

Interpretation each group of electrons acts as a density $d\Omega = d^3v f_0(\underline{v})$

has Doppler shift $\omega - \underline{k} \cdot \underline{v}$

$$\epsilon(\omega, \underline{k}) = 1 - \sum_{e,i} \int \frac{4\pi q_{e,i}^2 d\Omega_{e,i}}{m_e (\omega - \underline{k} \cdot \underline{v})^2}$$

$$d\Omega_{e,i} = \frac{4\pi q_{e,i}^2 d\Omega}{m_e}$$

Note, for electrostatic dielectric only velocities in direction of \underline{k} are important

$$\epsilon(\omega, \underline{k}) = 1 - \sum_{e,i} \frac{4\pi q^2}{m} \int \frac{d^3v f(v)}{(\omega - \underline{k} \cdot \underline{v})^2}$$

Take \underline{k} to be in x-direction

$$\epsilon(\omega, k) = 1 - \sum_{e,i} \frac{4\pi q^2}{m} \int \frac{dv_x dv_y dv_z f(v_x, v_y, v_z)}{(\omega - kv_x)^2}$$

Do v_y & v_z integrals

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$$\epsilon(\omega, k) = 1 - \sum_{e,i} \frac{4\pi q^2}{m} \int \frac{dv_x f(v_x)}{(\omega - kv_x)^2}$$

$$f_{1D}(v_x) = \int dv_y dv_z f(v_x, v_y, v_z)$$

NEXT STEP CONSIDER FIRST APPROXIMATION

$$\frac{1}{\omega - k \cdot \underline{v}} \underset{\text{assume}}{\approx} \frac{1}{\omega - k \cdot \underline{u} - k \cdot \delta \underline{v}} = \frac{1}{\omega_d - k \cdot \delta \underline{v}}$$

$$\frac{1}{\omega - k \cdot \underline{v}} = \frac{1}{\underbrace{\omega - k \cdot \underline{u} - k \cdot \delta \underline{v}}_{\text{doppler freq}}} = \frac{1}{\omega_d - k \cdot \delta \underline{v}}$$

EXPAND DENOMINATOR

$$\left(\frac{1}{(\omega_d - k \cdot \delta \underline{v})} \right)^2 \approx \frac{1}{\omega_d^2} \left[1 + 2 \frac{k \cdot \delta \underline{v}}{\omega_d} + \frac{3(k \cdot \delta \underline{v})^2}{\omega_d^2} + \dots \right]$$

$$\int d\underline{v} \frac{f_0(\underline{v})}{(\omega_d - k \cdot \delta \underline{v})} \approx \frac{1}{\omega_d^2} \int d\underline{v} f_0 \left(1 + 2 \frac{k \cdot \delta \underline{v}}{\omega_d} + \frac{3(k \cdot \delta \underline{v})^2}{\omega_d^2} + \dots \right)$$

$$= \frac{1}{\omega_d^2} \left[n_0 + \frac{0}{\omega_d} + \frac{3}{\omega_d^2} \int d\underline{v} (k \cdot \delta \underline{v})^2 f_0 \right]$$

$$\int d\underline{v} (k \cdot \delta \underline{v})^2 f_0 = \frac{1}{m} k \cdot \underline{n} \cdot k = \cancel{k^2 \cancel{n}} = \frac{k^2 P}{m}$$

for TE

USE $P = nT$

$$\epsilon \approx 1 - \frac{\omega_p^2}{\omega^2} \left(1 + \frac{3k^2 T}{m\omega^2} \right)$$

$\frac{T}{m}$ has units # of velocity^2

$$\frac{T}{m} = V_{th}^2 \quad V_{th} = \text{typical thermal velocity}$$

small correction if $|kV_t| \ll \omega$

or $V_t \ll \frac{\omega}{K}$ thermal velocity much less than phase velocity at wave.

Lets solve for ω for a given K

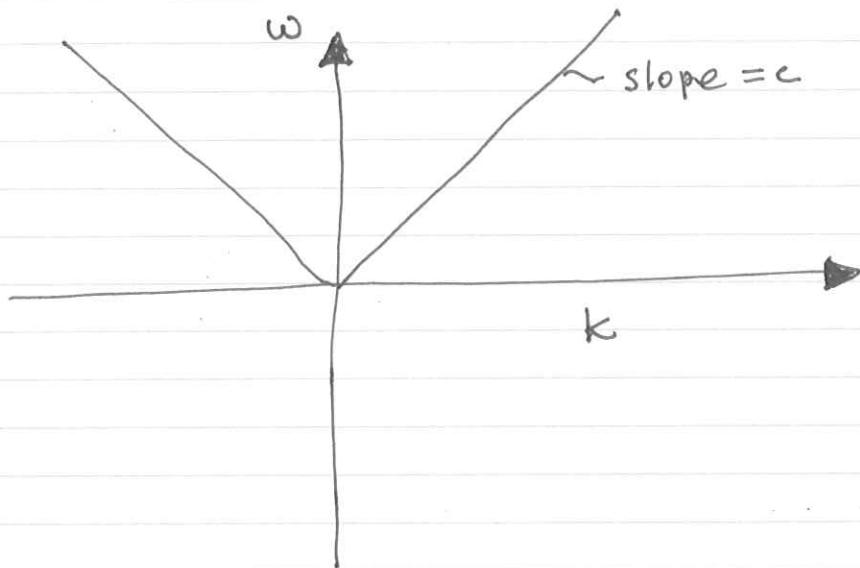
$$\epsilon(\omega, K) = 0 \Rightarrow \omega(K)$$

Normal mode frequency ~~is~~

This is known as a dispersion relation

For Light waves in Vacuum

$$\omega^2 = k^2 c^2$$



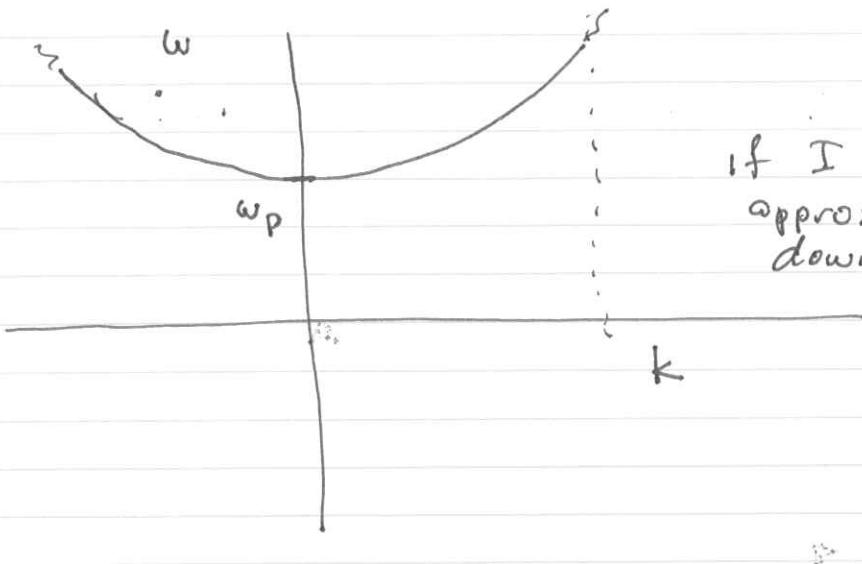
Solve $\epsilon \approx 0$

Appears to be a quadratic relation

$$\omega^2 \approx \omega_p^2 \left(1 + 3 \frac{k^2 T}{m \omega_p^2} \right) \approx \omega_p^2 + 3 \frac{k^2 T}{m}$$

~~$$\omega \approx \sqrt{\omega_p^2 + 3 \frac{k^2 T}{m}}$$~~

$$\omega = \sqrt{\omega_p^2 + 3 \frac{k^2 T}{m}} \approx \omega_p \left(1 + \frac{3}{2} \frac{k^2 T}{m \omega_p^2} \right)$$



Can we understand this based on fluid equations?

Ans: yes

$$\omega = \omega_p^2 \left(1 + \frac{3}{2} k^2 \lambda_0^2 \right)$$

$$\lambda_0^2 = \frac{T}{4\pi n e^2} = \frac{T}{m \omega_p^2}$$

good as long as $k \lambda_0 \ll 1$

wave lengths long compared with
Debye length

⊗ Apply same procedure to solution of fluid equations

continuity

$$\frac{\partial n}{\partial t} + \frac{\partial}{\partial x} n \underline{u} = 0$$

momentum

$$(mn) \left[\frac{\partial \underline{u}}{\partial t} + \underline{u} \cdot \nabla \underline{u} \right] = -\nabla p + q \overset{n}{\underline{E}}$$

$$\underline{E} = -\nabla \phi$$

internal energy

ratio of specific heats

$$\left(\frac{\partial}{\partial t} P + \underline{u} \cdot \nabla P \right) + \gamma_s P \nabla \cdot \underline{u} = 0$$

Equilibrium:

$$P = P_0 \quad n = n_0 \quad - \text{constants}$$

$$\underline{u}_0, \phi_0 = 0$$

no flow no potential

Perturbation

$$P = P_0 + \text{Re} \{ \hat{P} e^{ikx-i\omega t} \}$$

$$n = n_0 + \text{Re} \{ \hat{n} e^{ikx-i\omega t} \}$$

$$\underline{u} = \text{Re} \{ \hat{\underline{u}} e^{ikx-i\omega t} \}$$

$$\phi = \text{Re} \{ \hat{\phi} e^{ikx-i\omega t} \}$$

continuity

drop $\hat{n}\hat{u}$ (non linear term)

$$-i\omega \hat{n} + i\cancel{k \cdot n_0} \hat{u} = 0$$

$$m n_0 (-i\omega \hat{u}) = -i\cancel{k} \hat{p} - q n_0 i k \hat{\phi}$$

$$-i\omega \hat{p} + \gamma_s p_0 i \cancel{k \cdot u} = 0$$

eliminate $\hat{p} \& \hat{u}$ to find \hat{n} in terms of $\hat{\phi}$

$$\frac{\hat{p}}{p_0} = \gamma_s \frac{\hat{n}}{n_0}$$

$$-i\omega m n_0 \hat{u} = -i\cancel{k} \left(p_0 \gamma_s \frac{\hat{n}}{n_0} + q n_0 \hat{\phi} \right)$$

$$+ i\omega \hat{n} = i n_0 \cancel{k \cdot u}$$

$$\frac{p_0}{n_0} = T$$

$$\omega^2 m \hat{n} = k^2 \left(\gamma_s \frac{p_0}{n_0} \hat{n} + q n_0 \hat{\phi} \right)$$

$$\hat{n} = \frac{k^2 q n_0 \hat{\phi}}{\omega^2 - \frac{k^2 \gamma_s T}{m}}$$

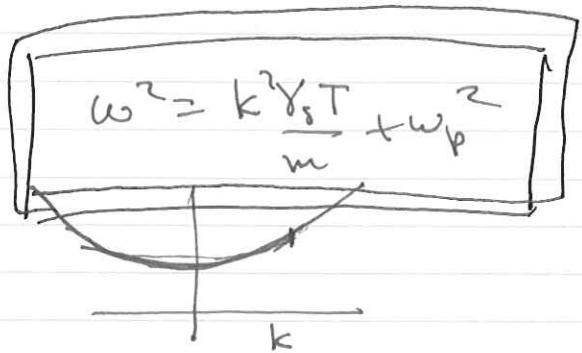
charge density


$$4\pi \hat{p} = 4\pi q \hat{n}$$

Poisson Equation

$$k^2 \hat{\phi} = 4\pi q \hat{n} = \frac{k^2 \omega_p^2}{\omega^2 - k^2 \gamma_s T/m}$$

$$\epsilon = 1 - \frac{\omega_p^2}{\omega^2 - k^2 \gamma_s T/m}$$



if we assume $k^2 \gamma_s T/m \ll \omega^2$ as before

$$\epsilon = 1 - \frac{\omega_p^2}{\omega^2} \left[1 + \gamma_s \frac{k^2 T}{m \omega^2} + \dots \right]$$

compare with



the same

$$\epsilon = 1 - \frac{\omega_p^2}{\omega^2} \left[1 + 3 \frac{k^2 T}{m \omega^2} + \dots \right]$$

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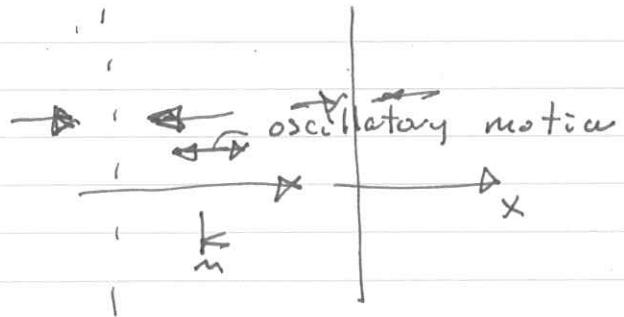
$$\gamma_s = \frac{n_f + 2}{n_f}$$

To get $\gamma_s = 3$, we must have $n_f = 1$
(not 3)

Explanation: ~~collisionless plasma~~

~~compression~~

Wave perturbation involves
a compression in 1D



This increases
internal
energy associated
with motion in
x direction

In a collisional fluid ~~beats~~ raises

$$\frac{1}{2} m s v_x^2$$

collisions quickly redistribute
this energy to other degrees of freedom

$$n_f = 3$$

$$\frac{1}{2} m s v_y^2 + \frac{1}{2} m s v_z^2$$

In collisionless plasma

energy remains
in $\frac{1}{2} m s v_x^2$

not shared

$$n_f = 1$$

9V

Landau Damping

We have been avoiding the a
big problem - the cut-off

$$\epsilon = 1 + \frac{4\pi q^2}{m k^2} \int d^3v \frac{k \cdot \frac{\partial f_0}{\partial v}}{(\omega - k \cdot v)}$$

OR

$$= 1 - \frac{4\pi q^2}{m} \int d^3v \frac{f_0}{(\omega - k \cdot v)^2}$$

what happens when $k \cdot v \approx \omega$

what happens when $\omega = k \cdot v$?

Integral is singular

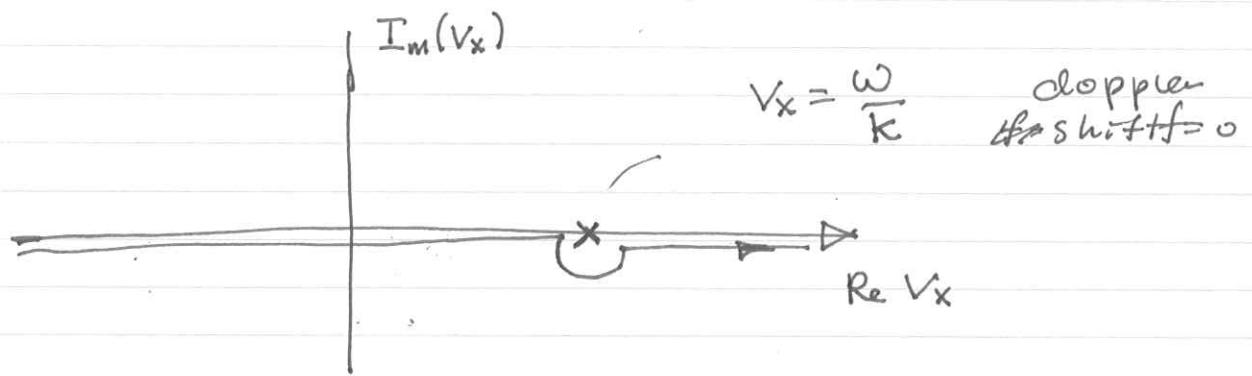
~~Waves~~

Focus on 1D

$$\epsilon = 1 + \frac{4\pi q^2}{m k^2} \int dv_x \frac{k \frac{\partial f_0}{\partial v_x}}{\omega - k v_x}$$

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SHORT ANSWER MUST PERFORM
CONTOUR OF INTEGRATION



E becomes complex

dissipation — waves are damped

Landau damping